

Nuclear and Particle Physics Experiments for Schools and Universities

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Increasing Interest in Outreach and Education in Physics

Recent trends show a significant increase in outreach and educational programs developed by universities and research institutions.

Focus on making physics and other scientific disciplines more accessible and engaging to a broader audience.

Goals of Outreach Initiatives

- Stimulate Interest in Science: Create engaging content and hands-on activities to capture the curiosity of diverse audiences.
- Promote Understanding: Simplify complex concepts to enhance comprehension of fundamental principles of physics.
- Inspire Future Scientists: Encourage young students to pursue careers in science through inspirational programs and role models.

Interactive Workshops

Public Lectures and Seminars

Collaborative Projects



- Increased Engagement: Higher participation rates in science fairs, workshops, and educational programs.
- Enhanced Understanding: Improved comprehension of scientific concepts among participants.
- Future Scientists: Growing interest in STEM careers among young students.

Impact

Educational Project

CAEN brings the experience acquired in more than 45 years of collaboration with the **High Energy & Nuclear Physics** community into the educational laboratories Worldwide.

CAEN enters the world of learning and training by providing **modern physics experiments for Advanced Labs** based on the latest technologies and instrumentation.



Inspire students and guide them towards the analysis and comprehension of different physics phenomena with a series of experiments based on **state-of-the art technologies**, instruments and methods.

Target the experiment depending on the student educational level. With this approach, the experiments proposed can be performed at high school level (grade 11,12) science classes up to undergraduate physics laboratory and PhD courses.



Educational Events



PhD schools

Outreach and Masterclass

University and PhD Laboratory courses

RESEARCH TOOLS

Tailored courses to meet trainee's needs

High school Laboratory courses

Training courses for High School teachers

Courses and schools also available on-site

CAEN Educational Training

CAEN offers training courses for different types of experiments used in educational laboratories by targeting them depending on:

- Students' educational level (from high school to PhD)
- Applications (from pulse processing electronics to nuclear safety)
- Expert users

2018 Erasmus Mundus JMD on Nuclear Physics
Co-funded by the Erasmus+ Programme of the European Union

2019 School on Waveform Digitizer
Tuesday, November 28 at 2:30 p.m. / 5 p.m. IST
Register Now

2° School on Waveform Digitizer

Nuclear and Particle Physics Experiments for Schools and Universities WEBINAR
Thursday, September 19 at 6:30 p.m. / 8:30 p.m.
Register Now

Webinars

2024

Courses:

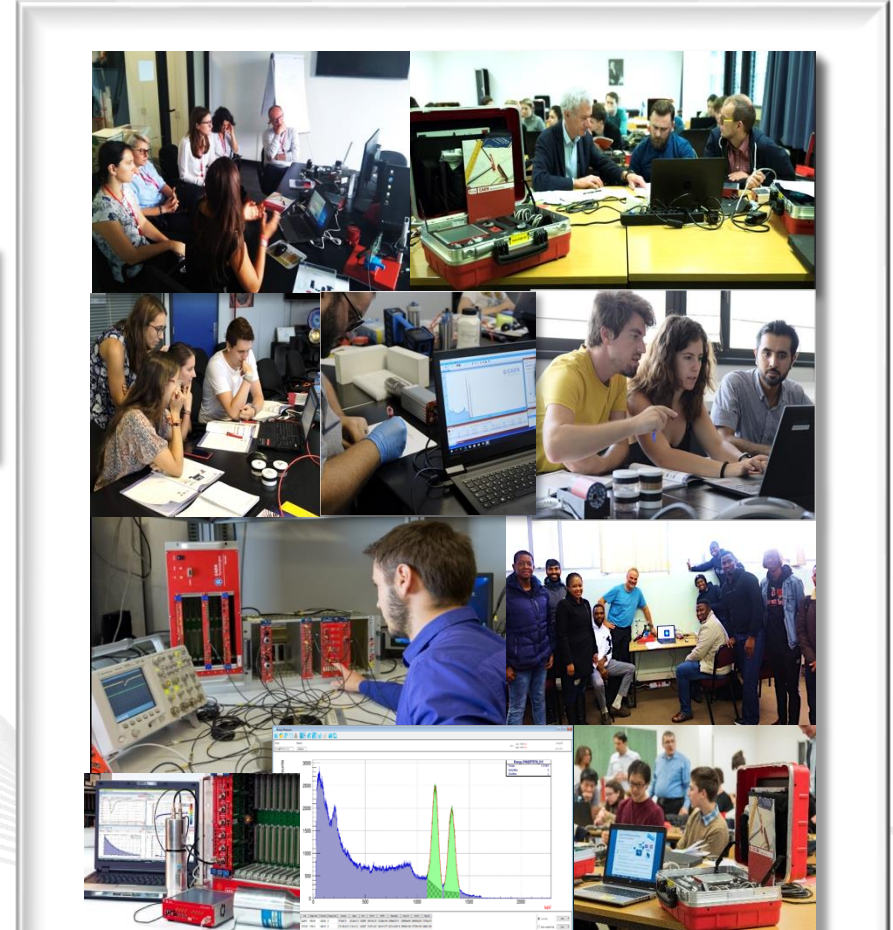
- Nuclear Physics
- Quantum Physics
- Environmental Radioactivity
- Cosmic Rays
- Emulation systems
- Nuclear Imaging
- Detectors Characterization
- Statistics
- Digital Pulses Processing
- FPGA Programming (Sci-Compiler based)
- Electronic Products

PRISMA SCHOOL 2018
Photosensors and Signal Processing in Particle Detectors
Mainz, 12 - 16 March 2018

PRISMA+ SCHOOL 2022
Photon Detection and FPGAs in Particle Detectors
Mainz, 26 - 30 September 2022

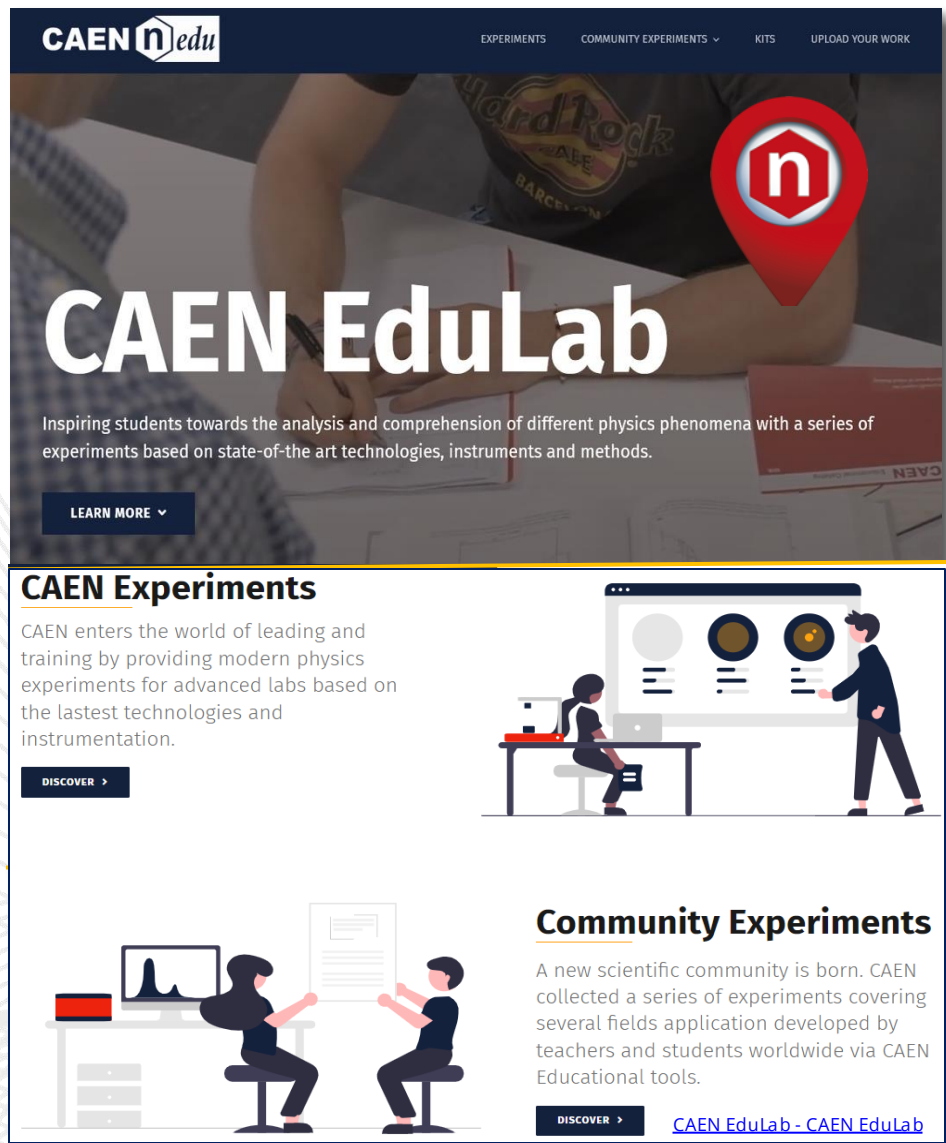
ANIMMA INTERNATIONAL SUMMER SCHOOL PISA 2023
The 1st Anima International School of Physics

INTERNATIONAL MOBILITY - IMSci-Nu Master
25 - 28 February 2024
Viareggio, Italy



All courses, taught by expert instructors and academics, are balanced between software, theoretical lessons and practical lab exercises to provide the maximum benefits:

- Discussion and constructive interaction with the other users and the expert staff
 - Practical Hands-on focused on the concepts covered in class lesson
 - Practical exercises on CAEN hardware and software tools
- Reference materials are also provided!



NEW Educational Website

Innovative Scientific Network

Caen introduces a global platform for Modern & Nuclear Physics education, enabling seamless sharing of experiments for students and professionals.

Interactive and User-Friendly Resources

The platform also features interactive tools that enhance communication and collaboration among members, along with comprehensive guides tailored specifically to each user's needs.

CAEN Educational Products

A wide range of experiments covering *Nuclear and Particle Physics fields!*

From the radioactive decays (β and γ) to the cosmic rays, from the light quanta to the advanced statistics and from the nuclear imaging to the emulation of the radioactive processes. Moreover, a new product line is fully focused on environmental radiation (indoor and outdoor) and on FPGA programming.

Nuclear Physics and Radioactivity

- γ Spectroscopy
- β -Radiation
- Nuclear Imaging – PET
- Γ Environmental Radioactivity (indoor)
- Γ Environmental Radioactivity (outdoor)
- GM detectors

Advanced Statistics based on Silicon Photomultiplier Detectors

Particle Detector Characterization

- Silicon Photomultiplier (SiPM)
- Photomultiplier Tube (PMT)

Particle Physics

- Photons
- Cosmic Rays

Electronics:

- Pulse Processing
- FPGA Programming



Nuclear Physics and Radioactivity

- γ Spectroscopy***
- ✓ Detecting γ -Radiation
 - ✓ Poisson and Gaussian Distributions
 - ✓ Energy Resolution
 - ✓ System Calibration: Linearity and Resolution
 - ✓ A comparison of different scintillating crystals: Light Yield, Decay Time and resolution
 - ✓ γ -Radiation Absorption
 - ✓ Photonuclear cross-section/Compton Scattering cross-section
- β-Radiation***
- ✓ Response of a Plastic Scintillating Tile
 - ✓ β Spectroscopy
 - ✓ β -radiation: Transmission through Matter
 - ✓ β -Radiation as a Method to Measure Paper Sheet Grammage and thin layer thickness
- Nuclear Imaging - PET***
- ✓ Basic Measurements: γ Spectroscopy and System Linearity
 - ✓ Positron Annihilation Detection
 - ✓ Two-dimensional Reconstruction of Source
 - ✓ Spatial Resolution
- γ Environmental Radioactivity (outdoor)***
- ✓ Environmental monitoring in land field
 - ✓ Ground Coverage Effect on the Environmental Monitoring
 - ✓ Human Body Radioactivity
 - ✓ Environmental detection as a function of the soil distance
 - ✓ Radioactivity maps production
 - ✓ Radiological evaluation of the building materials
 - ✓ Geochemical and mineral exploration
- γ Environmental Radioactivity (indoor)***
- ✓ Energy calibration of System based on LYSO crystal
 - ✓ Background Measurements
 - ✓ Fertilizer and photopeak identification
 - ✓ Identifications Sample Test
 - ✓ Soil sample identification
 - ✓ Samples Comparison
 - ✓ Radon passive measurements

- GM Detectors***
- ✓ Statistics: Uncertainty as a function of live time
 - ✓ Environmental Background
 - ✓ Lead Shielding Effect on Environmental Radioactive Background
 - ✓ Detecting Ionizing-Radiation
 - ✓ Samples Comparison

Particle Physics

- Photons***
- ✓ Quantum Nature of Light
 - ✓ Hands-on Photon Counting Statistics
- Cosmic Rays***
- ✓ Statistics
 - ✓ Muons Detection
 - ✓ Muons Spectrum
 - ✓ Muons Vertical Flux on Horizontal Detector
 - ✓ Zenith Dependence of Muons Flux
 - ✓ Random Coincidence
 - ✓ Detection Efficiency
 - ✓ Cosmic Flux as a function of the altitude
 - ✓ Cosmic Shower Detection
 - ✓ Environmental and Cosmic Radiation
 - ✓ Absorption Measurements
 - ✓ Solar Activity Monitoring

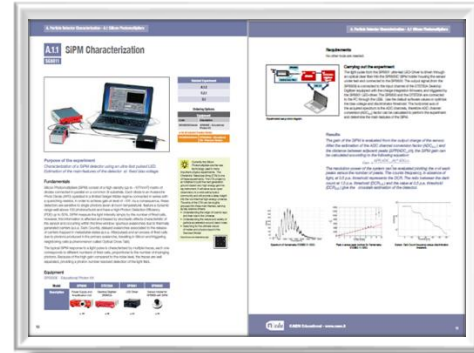
Particle Detector Characterization

- Silicon Photomultiplier (SiPM)***
- ✓ SiPM Characterization
 - ✓ Dependence of the SiPM Properties on the bias voltage
 - ✓ Temperature Effects on SiPM Properties
- Photomultiplier Tube (PMT)***
- ✓ Measurement of Photomultiplier Plateau Curves

Pulse Processing: Open FPGA

- ✓ Analog signal acquisition and waveform Visualization
- ✓ Waveform digitizer with leading edge trigger.....

Advanced Statistics..



Short Guide

Main Topics:

- Experiment task
- Short description
- Equipment list
- Requirements
- Quick guide
- Experimental results



Detailed Guide

Guide Topics:

- General Information
- Introduction
- Physics Pills
- Required Equipment
- Getting Started
- Experimental Procedure
- Results
- Links related to this topic

A lot of experiments on handy!

Nuclear Physics and Radioactivity

γ Spectroscopy

- ✓ Detecting γ -Radiation
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Nuclear Imaging - PET

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GM Detectors

- ✓ Statistics: Uncertainty as a function of live time
- ✓ Environmental Background
- ✓ Lead Shielding Effect on Environmental Radioactivity
- ✓ Detecting Ionizing-Radiation
- ✓ Samples Comparison



Particle Physics

Photons

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Cosmic Rays

- ✓ Statistics
- ✓ Muons Detection
- ✓ Muons Spectrum
- ✓ Muons Vertical Flux on Horizontal Detector
- ✓ Zenith Dependence of Muons Flux
- ✓ Random Coincidence
- ✓ Detection Efficiency
- ✓ Cosmic Flux as a function of the altitude
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- ✓ Environmental and Cosmic Radiation
- ✓ Absorption Measurements
- ✓ Solar Activity Monitoring



Particle Detector Characterization

Silicon Photomultiplier (SiPM)

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Photomultiplier Tube (PMT)

- ✓ Measurement of Photomultiplier Plateau Curves

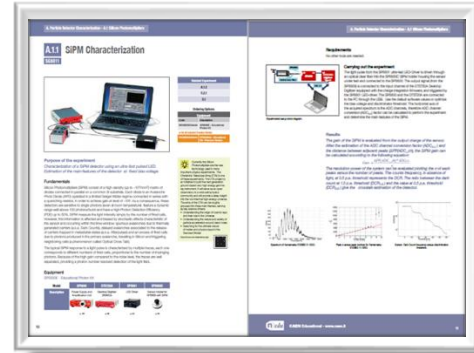


Pulse Processing: Open FPGA

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Advanced Statistics..



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Detailed Guide

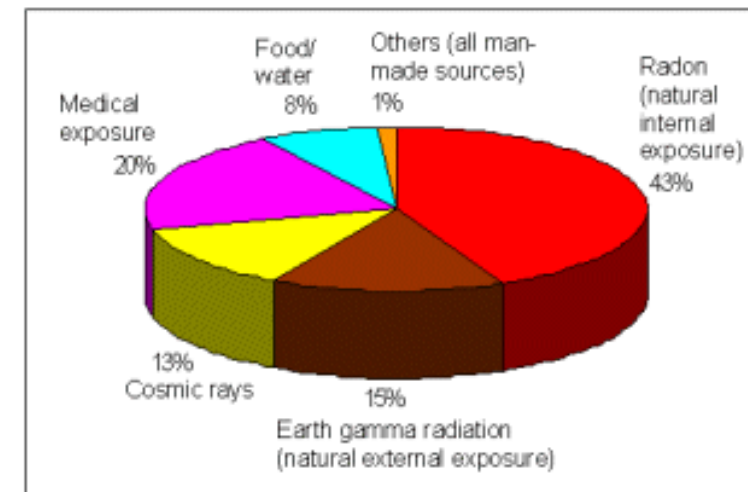
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- Results
- Links related to this topic

Radiation is a **natural part** of our environment!

Radioactive sources:

- Natural: NORM (Naturally Occurring Radioactive Material), soil, water, air and food contribute to our exposure to ionizing radiation
- Cosmic rays
- Industrial: nuclear elements produced by industry
- Medicine: nuclear medicine
- Military



https://www.who.int/ionizing_radiation/env/en/

Radioisotope types

Natural radionuclides

[mean abundance]:

- ^{40}K [2-2.5] %
- $^{238,235}\text{U}$ [2-2.5] ppm
- ^{232}Th [8 -12] ppm
- All has an half life $T_{1/2} >$ than the age of the solar system

Cosmic rays

Muons -> The intensity depends on the altitude and direction (N/S vs E/W)

Radionuclides from cosmic rays

- $^{14}\text{C}, ^7\text{Be}, ^3\text{He}$

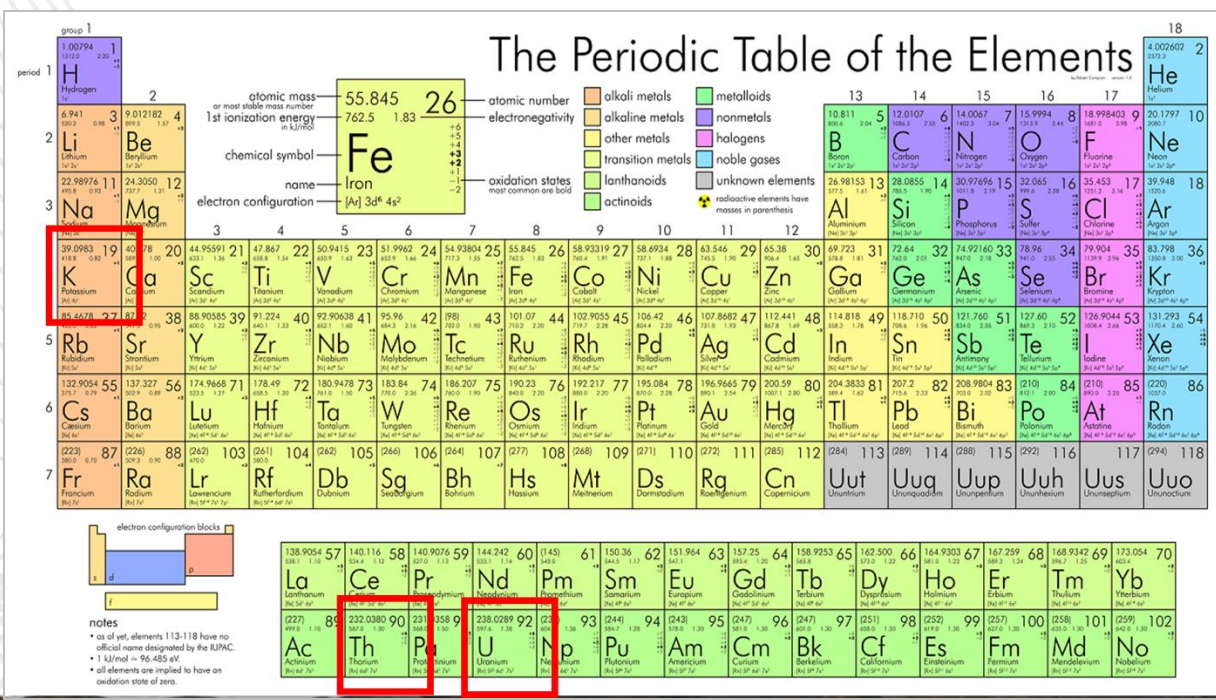
Artificial radionuclides

- From bombs or Nuclear power plants (ex. ^{137}Cs , actinides)
- Industrial (^{133m}Xe , ^{133}Ba , ^{241}Am) and medical (^{19}F , ^{67}Ga) radioisotopes

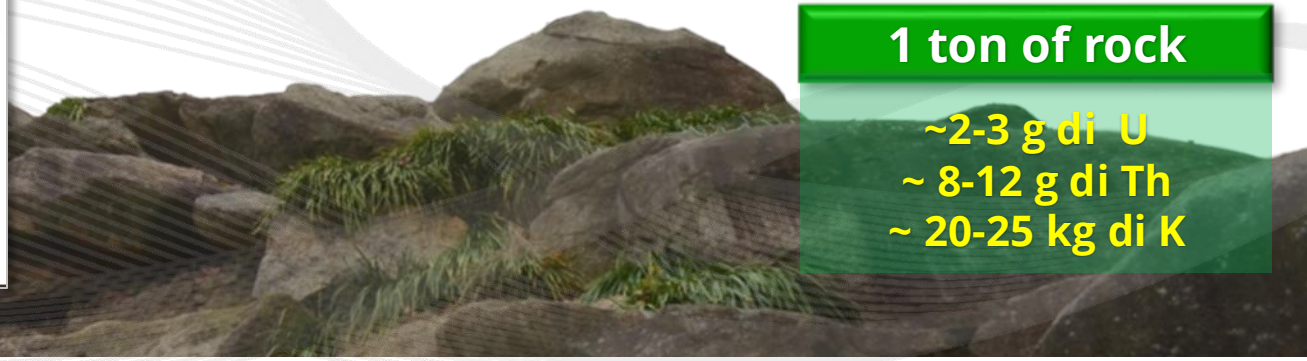
During the creation of the Earth, most of the elements initially produced were radioactive and they have been decayed to more stable forms.

The original radioactive elements still present on Earth are those that have a halftime comparable to the Earth. They are responsible for environmental radioactivity and internal warming of the planet and originate from elements very heavy without stable isotopes.

They mostly decay through the α and β channels



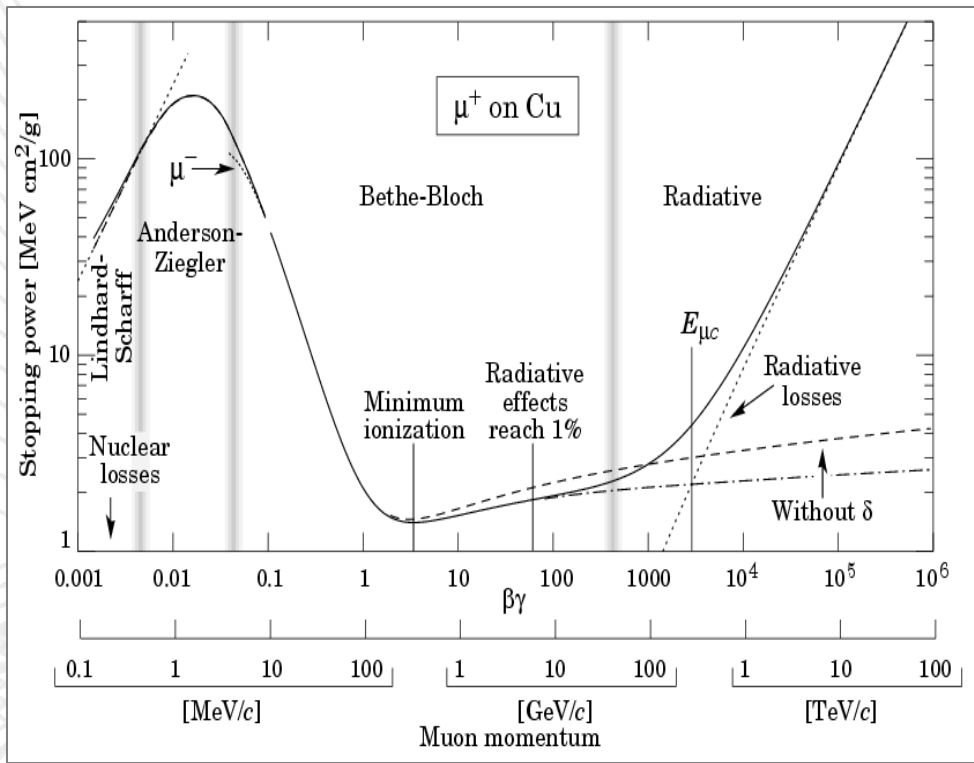
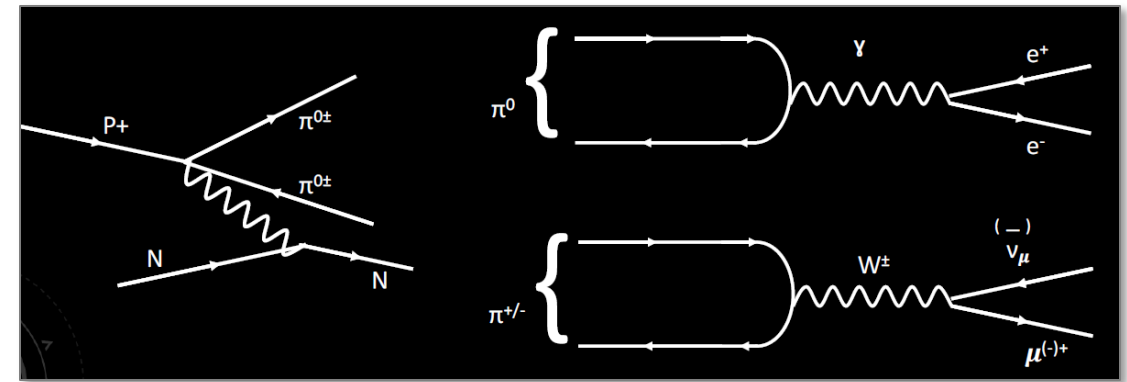
Element	Radioisotopes	Isotopic Adundance	Half time	Tipycal Adundance
Potassium	⁴⁰ K	0.012%	1.3 × 10 ⁹ anni	0.02 g/g [2%]
Uranium	²³⁸ U	99.3 %	4.5 × 10 ⁹ anni	2-3 µg/g [ppm]
Thorium	²³² Th	100 %	14.1 × 10 ⁹ anni	8-12 µg/g [ppm]



1 ton of rock
 ~2-3 g di U
 ~ 8-12 g di Th
 ~ 20-25 kg di K

High energy cosmic protons collide with atmospheric nuclei and produce **hadrons showers**. The only decay that is interesting due to long lifetime are pion channels.

π^0 channel decay produce e^\pm but in atmosphere they do not have a long range. Otherwise π^\pm decay μ^\pm and they have a longer range



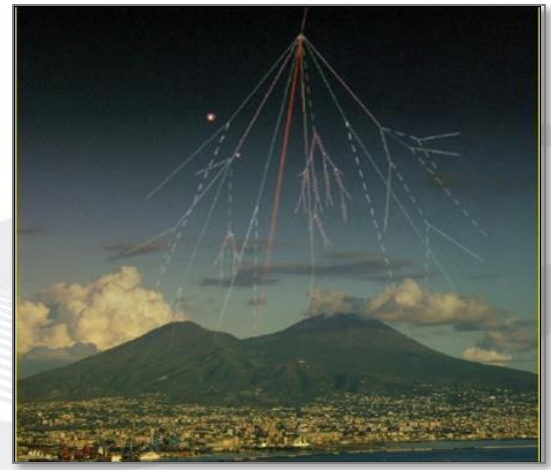
The Cosmic rays at sea level are made by high energy muons (**mean energy 4 GeV**), pions, proton and other soft component.

- The soft component is about 30% of total CR flux (according to PDG, 2017)
- The high energy protons fraction is 3.5% of total high energy muon flux
- Energy spectrum flat above 1GeV, than from range 10-100 GeV flux goes as $E^{-2.7}$

Muons $\{m_\mu = 105.7 \text{ MeV}/c^2\}$

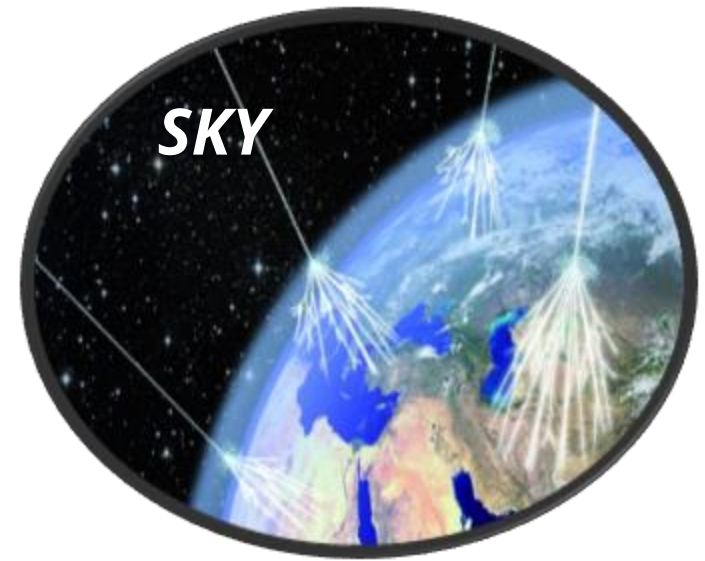
- Electromagnetically interaction (or weakly)
- Not stable $\rightarrow \tau(\mu) = 2.2 \times 10^{-6} \text{ s}$
- Produced at $\sim 15 \text{ km}$ altitude
- Mean energy @ground $\sim 4 \text{ GeV}$
- Mean high μ energy flux at sea level (PDG)

$$I_\nu = 0.82 \cdot 10^{-2} \frac{1}{\text{sr} \cdot \text{s} \cdot \text{cm}^2}$$



Educational Kits Suitable for Young Students Also!

Interesting educational program focused on the environment that surrounds us!



NEW Educational kits suggested for the environmental radioactivity experiments

Indoor solutions!

SP5630EN
Environmental kit



SP5630ENP
Environmental kit Plus



SP5660
RockyRAD

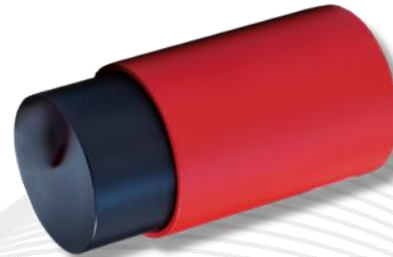


SP5640
Gamma EDU

Outdoor solutions!

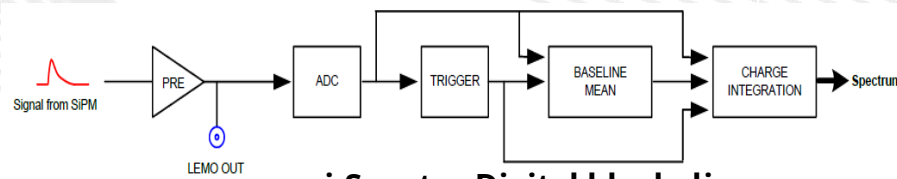
NEW SP5630 Environmental kit

Environmental gamma radiation measurements with SiPM based instrumentation!



S2570 - i-Spector Digital

- The system is based on a SiPM area $18 \times 18 \text{ mm}^2$ **Hamamatsu S14160-60520HS**. All SiPMs of the area are connected in parallel to increase the active area of the matrix.
- It integrates a shaper, a peak stretcher and a peak ADC to implement a simple MCA (4K).
- Scintillator Crystal: CsI $18 \times 18 \times 30 \text{ mm}^3$
- Energy Range: 30 keV to 3 MeV
- Energy Resolution (FWHM): $<6\%$ @ 662 keV
- Connectivity: Ethernet
- Software: Web GUI

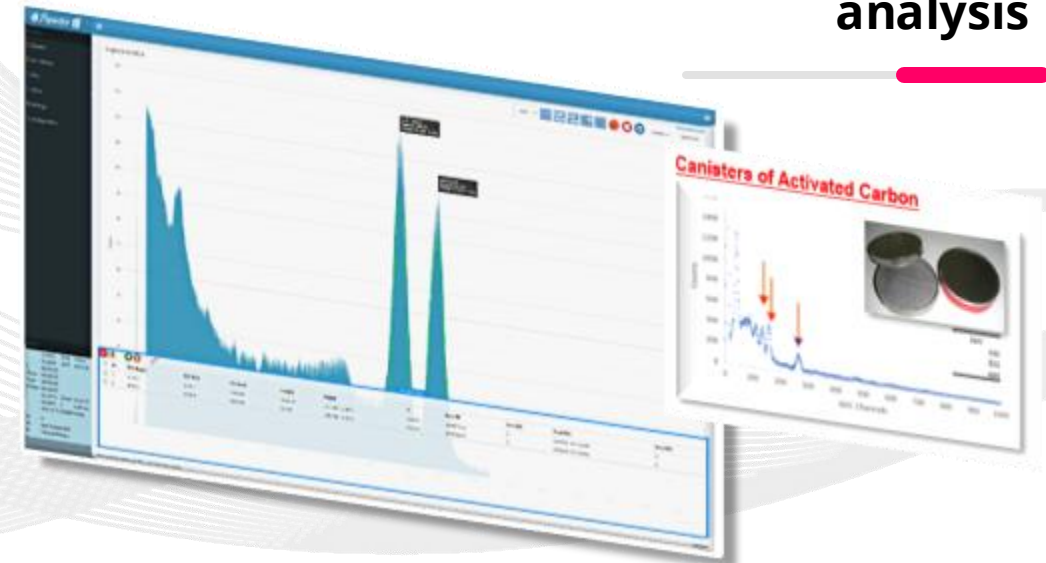


i-Spector Digital block diagram

Samples



Web-based GUI for unit control and data analysis



NEW SP5630ENP Environmental kit Plus

Gamma Radiation and Shielding Laboratory

The **SP5630ENP Environmental kit Plus** guide the users towards the development of complementary measurement techniques based on counting and on the analysis of the spectrum. The main goal is the study of the absorption of the gamma rays passing through matter thicknesses and the related observations about the different crossed materials. It is a user-friendly system for Advanced Labs based on the latest technologies and instrumentation.



S2570D i-Spector Digital
18x18mm² - CsI ASSEMBLY



BGO scintillating crystal



**SP5630EN
Environmental kit Plus**

Samples



Empty Beaker & Test Sample



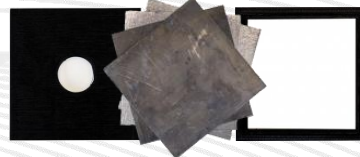
Fertilizer and Rock Samples



Canisters of Activated Carbon



Calibration Crystal
(Lu_{1.8}Y₂SiO₅:Ce)



Shielding Kit

Web-based Software



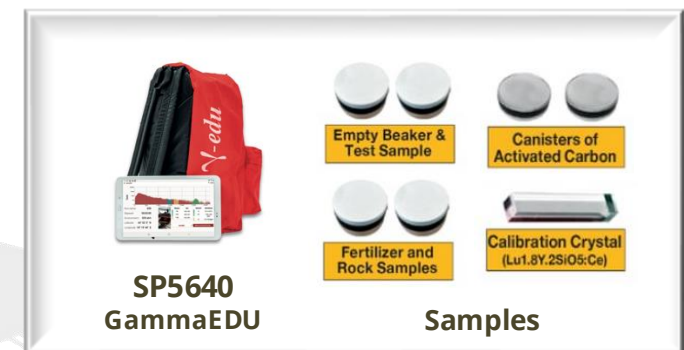
γ Environmental Radioactivity (indoor)

Section	Subsection	Experiment	Equipment											
			SP5600C	SP5600D	SP5600E	SP5600AN	SP5600EMU	SP5700	SP5701	SP5620CH	SP5630EN	SP5640	SP5630ENP	SP5650
Nuclear Physics and Radioactivity	γ Environmental Radioactivity (indoor)	Energy calibration of System based on LYSO crystal									*		*	
		Background Measurements									*		*	
		Fertilizer and photopeak identification									*		*	
		Identification Test Sample									*		*	
		Soil sample identification									*		*	
		Samples Comparison									*		*	
		Radon passive measurements									*		*	

★ Recommended kits



Alternative Choice



Hands-on

SP5630EN – Environmental Kit



NEW Educational kits suggested for the environmental radioactivity experiments

Indoor solutions!

SP5630EN
Environmental kit



SP5630ENP
Environmental kit Plus



SP5660
RockyRAD



SP5640
Gamma EDU

Outdoor solutions!



NEW GammaEDU - SP5640

Portable detection backpack for environmental radioactivity!

- NaI(Tl) (0.3 liter) Scintillator Crystal coupled to a PMT
- Power Supply included
- Identification of Natural Radiation [^{238}U , ^{232}Th , ^{40}K]
- Autonomy up to 6-8 hours
- Tablet included with GammaEDU Application
- Bluetooth and Wi-Fi Connectivity
- Geolocation and ability to view the map on Google Earth

Suitable for High School Students!



NEW GammaEDU - SP5640

Tablet

Tablet 10' with GammaEDU Application



Digital MCA Unit - S2580 - GAMMASTREAM

- High Voltage Power Supply (0 ÷ +1500V/500 μ A) - Charge Sensitive Preamplifier - digital Multi-Channel Analyzer (12-bit and 62.5 MHz ADC) for scintillation spectroscopy
- Coupled with NaI(Tl) with a 14-pin PMT
- Full stand-alone operation with embedded CPU, data storage (SSD) unit, and power supply for up to 6/8 hours operation
- Wired and wireless connectivity via USB, Ethernet, Wifi and Bluetooth
- Acquisition modes: PHA, PHA with time stamp, Signal Inspector

General Properties

Density(g/cm ³)	3.67
Melting point(K)	924
Wavelength of emission peak(nm)	415
Light output(Photons/Mev)	40,000
Decay time(ns)	264
Cleavage plane	(100)
Hygroscopic	Yes
Refractive index	1.85
Hardness(Mho)	2

NaI(Tl) Scintillator

Thallium doped sodium iodide, NaI(Tl), is the most widely used scintillation material, it has the greatest light output and convenient emission range

Dimension: 0.3 l

Applications

- Environmental Gamma detection and spectroscopy
- Mapping of potential radon-prone areas
- Environmental monitoring in land field
- Geochemical and mineral exploration
- Statistics
- Customs protection and border control

NEW GammaEDU App

Measurement Results

Geolocation

Picture

^{40}K , ^{214}Bi , ^{208}Tl Isotopes CPS

^{40}K , ^{238}U , ^{232}Th Abundances

Data

Displayed on Google Earth

.kmz file

Easy to share

Interface

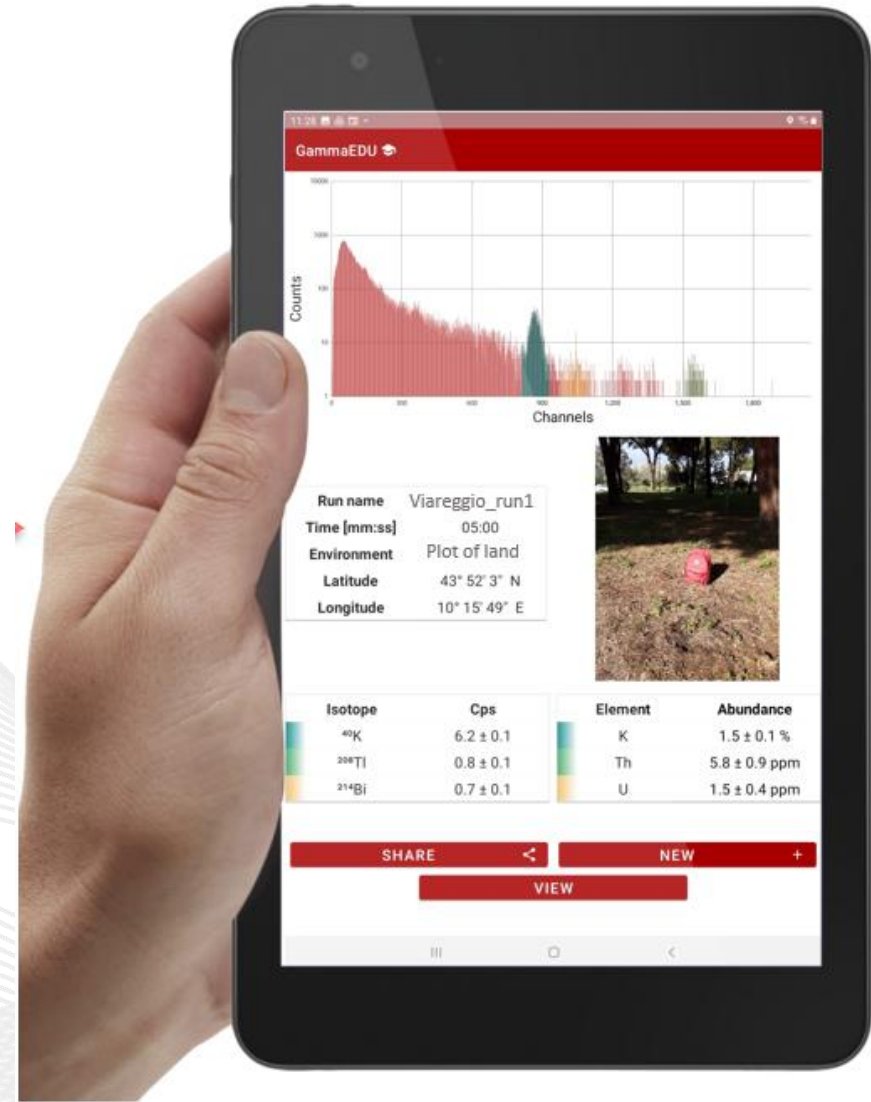
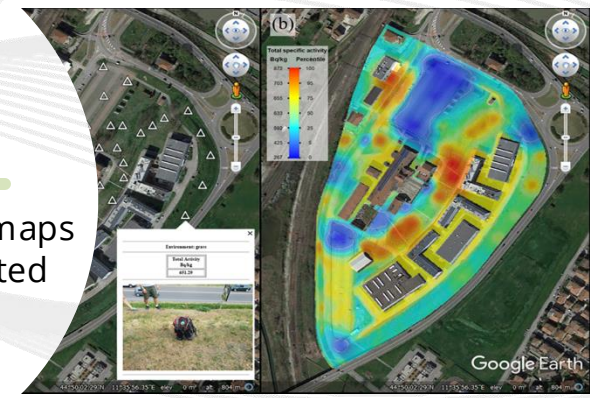
Bluetooth connection

Maps development

Studies in the field of earth sciences

**Coming
SOON**

New software for creating maps with color intensity associated with local radioactive contribution!

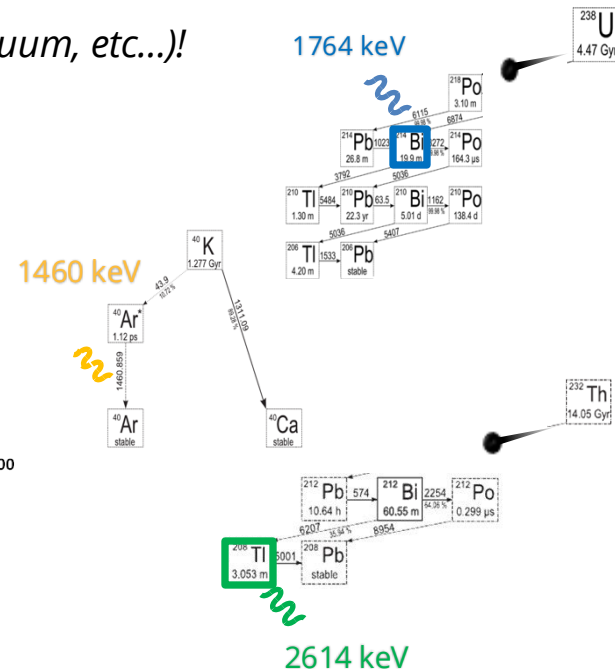
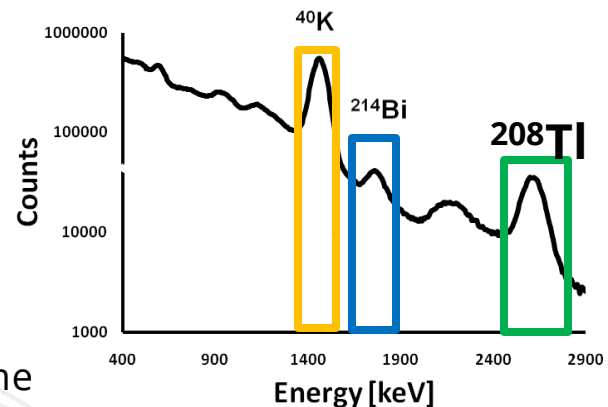


ROI Definition & Spectra Analysis

Definition of the **Region(s) of Interest (ROI)** of the energy spectra. These windows are used to define the photopeak regions required to calculate the correspondant areas (integral of the ROI).

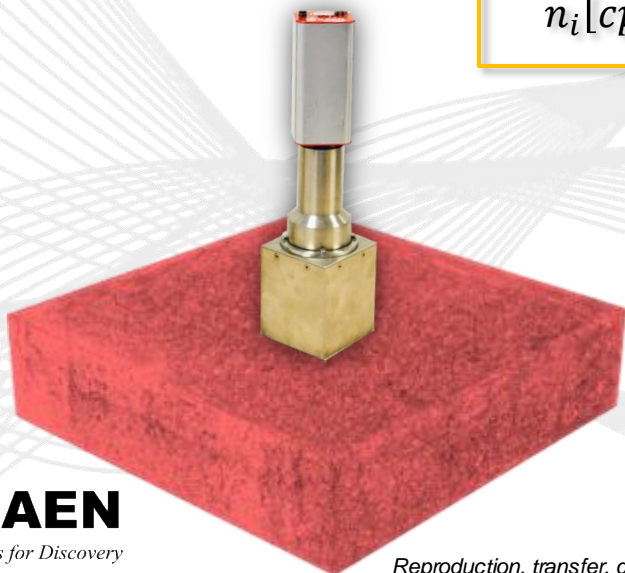
NOTE: In every ROI there are different contribution effects (photopeak, compton continuum, background continuum, etc...!)

Window	Isotope	Photopeak Energy (keV)	ROI (keV)
Potassio	^{40}K	1460	1370-1570
Uranio	^{214}Bi	1765	1660-1860
Torio	^{208}Tl	2614	2410-2810



Count rate: number of counts per unit of time

$$n_i[\text{cps}] = \frac{N_i[\text{conteggi}]}{T[\text{s}]}$$



Calibration: from counts to the abundancies

$$N(\text{cps}) = A \cdot S$$

$$S = A^{-1} \cdot N$$

S is the sensitivity coefficient of the detector

$$A = N \cdot S^{-1}$$

Calibration site characterized by known abundancies (A) of U, Th and U.

γ Environmental Radioactivity (outdoor)

Section	Subsection	Experiment	Equipment									★ SP5640		
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Nuclear Physics and Radioactivity	γ Environmental Radioactivity (outdoor)	Environmental monitoring in land field											★	
		γ Environmental Detection as a function of the soil distance												★
		Radioactivity maps production										★	★	
		Mapping of potential radon-prone areas										★	★	
		Geochemical and mineral exploration										★	★	

★ Recommended kits



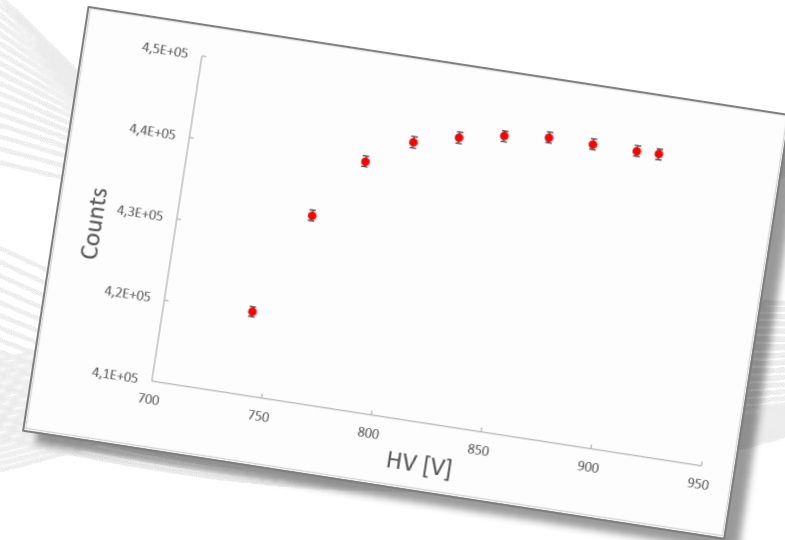
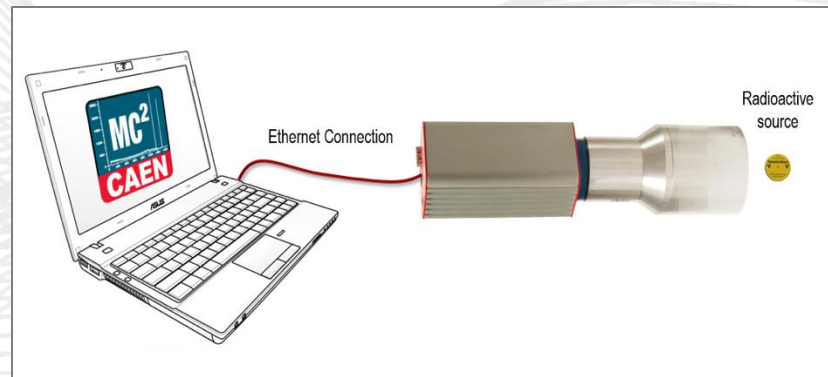
Alternative Choice



Photomultiplier Tube (PMT)

★ Recommended kits

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Particle Detector Characterization	Photomultiplier Tube (PMT)	<i>Measurement of Photomultiplier Plateau Curves</i>										☼



Hands-on

SP5640 – GammaEDU



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Environmental kit



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RockyRAD



SP5640
Gamma EDU

Outdoor solutions!

NEW



RockyRAD

Portable geiger counter for nuclear radiation

Geiger module

- Geiger Module GM J305
- Display OLED 128x64 pixels 1.54"
- Measuring range: 10 nSv/h – 50 μ Sv/h
- Sensitivity: 44 CPS/10 μ Sv/h (relative to ^{60}Co)
- Size: 71 * 136 * 43.8 mm³
- Bluetooth connection

Battery-powered

- Rechargeable Li-Ion battery, 3.7Vdc, 10Ah, 37Wh
- Power Consumption ~ 0.05W
- Power Supply: 5 V (USB-C)

Set of Rocks and Minerals included

Each RockyRAD kit includes diverse rock samples for immediate detection experiments.

NEW App



Key Features

- Start new measurement sessions
- Capture photographic evidence linked to specific readings
- Specify measurement intervals
- Organize and review collected data effectively

Display information

- Total Counts
- Counts Per Minute
- Equivalent Dose Rate

Interface

- Bluetooth connection
-

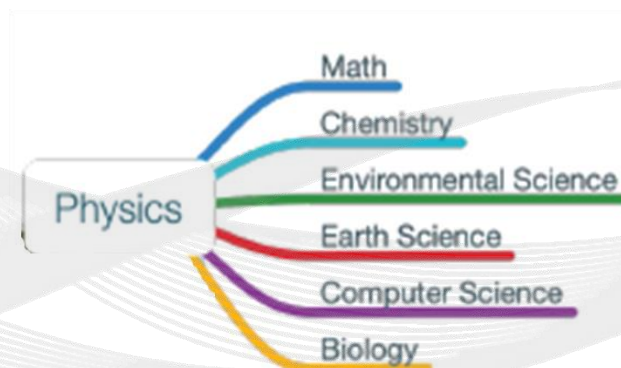


RockyRAD application requires Android™ mobile with Android™ release from 11.

GM Detector

Section	Subsection	Experiment	Equipment										
			SP5600C	SP5600D	SP5600E	SP5600AN	SP5600EMU	SP5700	SP5701	SP5620CH	SP5630EN	SP5640	SP5660
Nuclear Physics and Radioactivity	GM Detector	Statistics: Uncertainty as a function of live time											*
		Environmental Background											*
		Lead Shielding Effect on Environmental Radioactive Background											*
		Detecting Ionizing Radiation											*
		Samples Comparison											*

Recommended kits



RockyRAD_ References

CPM	Estimated Dose Rate [nSv/h]	Environment and Comments
14-34	114-274	Typical natural background radiation, corresponding to an annual dose of $1-2.4 \cdot 10^3 \mu\text{Sv}$.
28-48	228-388	Represents an increase of $10^3 \mu\text{Sv}$ per year compared to the typical background, aligning with the public exposure limit.
281	2282	The standard annual exposure limit for radiation workers in the nuclear industry, equating to $2 \cdot 10^4 \mu\text{Sv}$.
493	4000	Exposure at this level during high-altitude flights is generally safe due to the limited duration of flights. The same dose rate near a ground-based source would necessitate caution and protective measures, highlighting the importance of exposure context.
6000	48720	Maximum detectable scale by the Rockyrad device.

^[1] United Nations Scientific Committee on the Effects of Atomic Radiation. Sources and effects of ionizing radiation, united nations scientific committee on the effects of atomic radiation (UNSCEAR) 2008 report, volume I: Report to the general assembly, with scientific annexes A and B-sources. United Nations, 2010.

^[2] ICRP, International Commission on Radiological Protection. "Dose limits". ICRPedia. ICRP. Retrieved 26 April 2022.

^[3] Directive 2013/59/Euratom - protection against ionising radiation". European Agency For Safety And Health At Work. European Agency For Safety And Health At Work. Retrieved 26 April 2024.

^[4] Shea, M.A.; Smart, D.F. (August 2001). Comment on Galactic Radiation Dose to Air Crews. 27th International Cosmic Ray Conference. Bibcode:2001ICRC...10.4071S.

Hands-on

SP5660 – RockyRAD



NEW Educational kits dedicated to the Cosmic Rays detection



SP5620CH - Cosmic Hunter



SP5622B - Detection System Plus

NEW Cosmic Hunter – SP5620CH



A very simple cosmic Muon telescope!

The system consists of two scintillating tiles and a central board that counts the coincidences between the detectors, displaying the count on a numeric display.

SP5621 Coincidence Unit



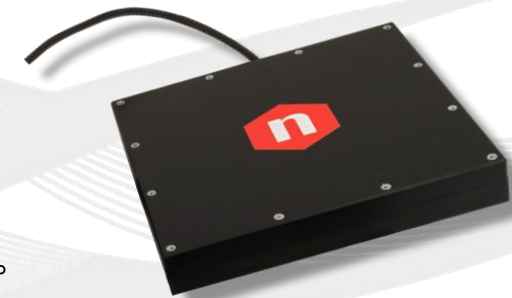
- The main unit houses a microcontroller based on the ESP32, the e-book display, and some interface and coincidence circuits.
- The output of the electronics is LVDS, and the board is powered by 5V.
- The operational commands relate to the type of coincidence (double, single, or even triple), the integration time, and the commands via the START, STOP, and RESET buttons.

- **Based on SiPM detectors and plastic scintillating tiles.**
- **Up to 3 scintillating tiles management**
- **No fixed geometry**
- **No Needs SW interface**
- **SD card to download data**

SP5622 Detection System

Each unit consists of:

- Plastic scintillator (15 x 15 x 1 cm²)
- Front-end electronic board (transconductance amplifier and a fast discriminator)
- SiPM (4 x 4 mm²) mounted in the tile corner at 45°

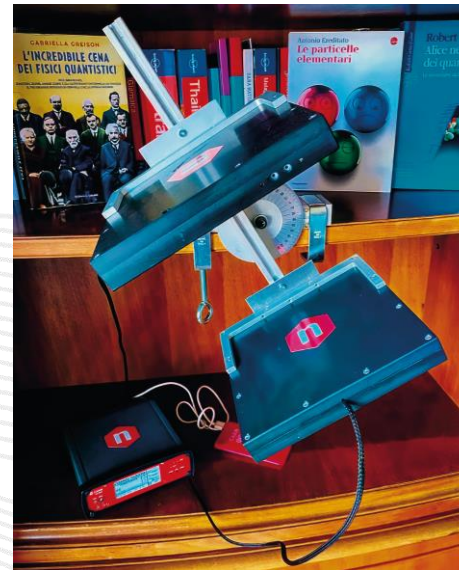


NEW Cosmic Hunter – SP5620CH



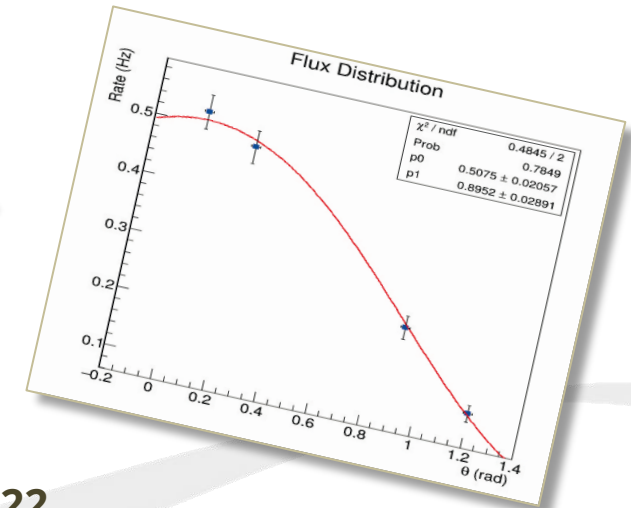
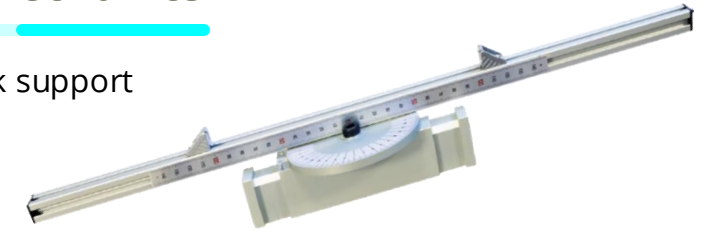
Additional Tools

- Based on SiPM detectors and plastic scintillating tiles.
- Up to 3 scintillating tiles management
- No fixed geometry
- No Needs SW interface
- SD card to download data



SP5609 Telescope Mechanics

- Rotary axis with desk support
- Clamps with screws
- Angle brackets kit



SP5622 Detection System

Each unit consists of:

- Plastic scintillator ($15 \times 15 \times 1 \text{ cm}^2$)
- Front-end electronic board
- SiPM ($4 \times 4 \text{ mm}^2$) mounted in the tile corner at 45°



NEW Cosmic Hunter – SP5620CH



num	coinc	date	time	sec	RecTime	TOP	BOTTOM	EXT	COINC
1	T-B	14/1/2019	16:53:45	601	600	6552	8395	0	529
2	T-B	14/1/2019	17:3:45	1201	600	6838	8652	0	655
3	T-B	14/1/2019	17:13:44	1801	600	6649	8582	0	522
4	T-B	14/1/2019	17:23:44	2401	600	6621	8481	0	503
5	T-B	14/1/2019	17:33:44	3001	600	6647	8480	0	495



Cosmic Hunter Software

- Based on SiPM detectors and plastic scintillating tiles.
- Up to 3 scintillating tiles management
- No fixed geometry
- No Needs SW interface
- SD card to download data
- **New Management Software**

Tools for Discovery Cosmic HUNTER v1.0.9 SP5621 CAEN

Dashboard | Tile Counts Charts | Coincidences Charts | Environment Data Chart | Configuration

Data Persistence Configuration

Data Saving Data

Output File Path
C:\Users\pbarba\Documents\cosmichunter Select Folder

Altitude Correction Configuration

Current Offset (m): 73,7 Change Altitude Correction

Altitude

Date/Time Configuration

Computer: 17:20:57 20/10/2022 Time

Device: 17:20:57 20/10/2022 Synchronize

Copyright 2022 - C.A.E.N. S.p.a.

Cosmics

Section	Subsection	Experiment	Equipment									
			SP5600C	★ SP5600D	SP5600E	SP5600AN	SP5600EMU	SP5700	SP5701	👤 SP5620CH	SP5630EN	SP5640
Particle Physics	Cosmic Rays	Muons Detection		*		*				*		
		Muons Vertical Flux on Horizontal Detector		*		*				*		
		Muons Spectrum		*		*						
		Zenith Dependence of Muons Flux		☉ ▲		☉ ▲				☉ ▲		
		Triple coincidence								☉		
		Cosmic Shower Detection								☉		
		Random Coincidence								*		
		Detection Efficiency								☉ +		
		Cosmic Flux as a function of the altitude								*		
		Environmental and Cosmic Radiation								*		
		Absorption Measurements								*		
		Solar Activity Monitoring								*		

★ Recommended kits



Alternative Choice



👤 This symbol suggests the kit use for young students!

Outreach & Balloon Experience

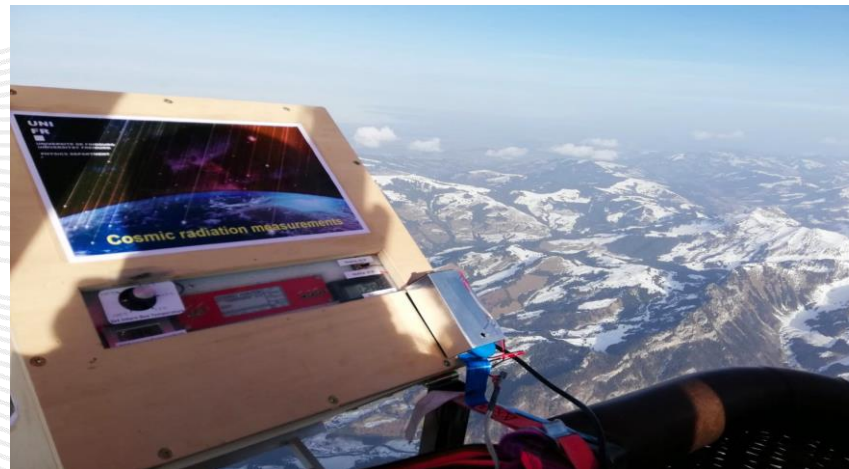


Les Rencontres de Physique de la Vallée d'Aoste

The INFN OCRA project involved many high school students with experiments measuring cosmic rays in Aosta (IT) and at high altitudes.



Commemorative balloon flight 25 January 2020



Prof. Hans Peter Back

Albert Einstein Center for Fundamental Physics, University of Bern

High School Student's Educational Experience



Jonas Hoecker
Mr. Fluckiger

Volée 2019-2021

Radiographie terrestre avec les rayons cosmiques

Quel est le taux de rayons cosmiques en fonction de l'altitude et de la profondeur géologique? Théorie et mesures.



8369 mots

Rendu le 7 Janvier 2020

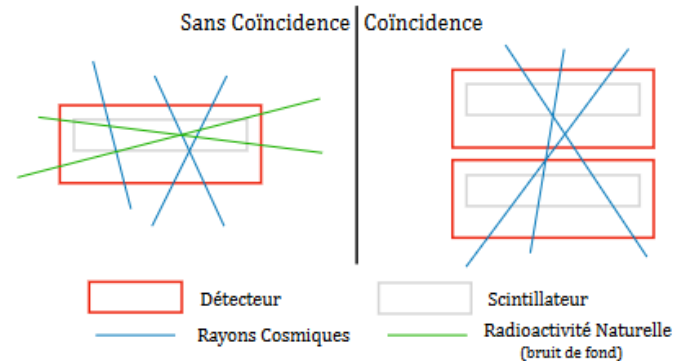
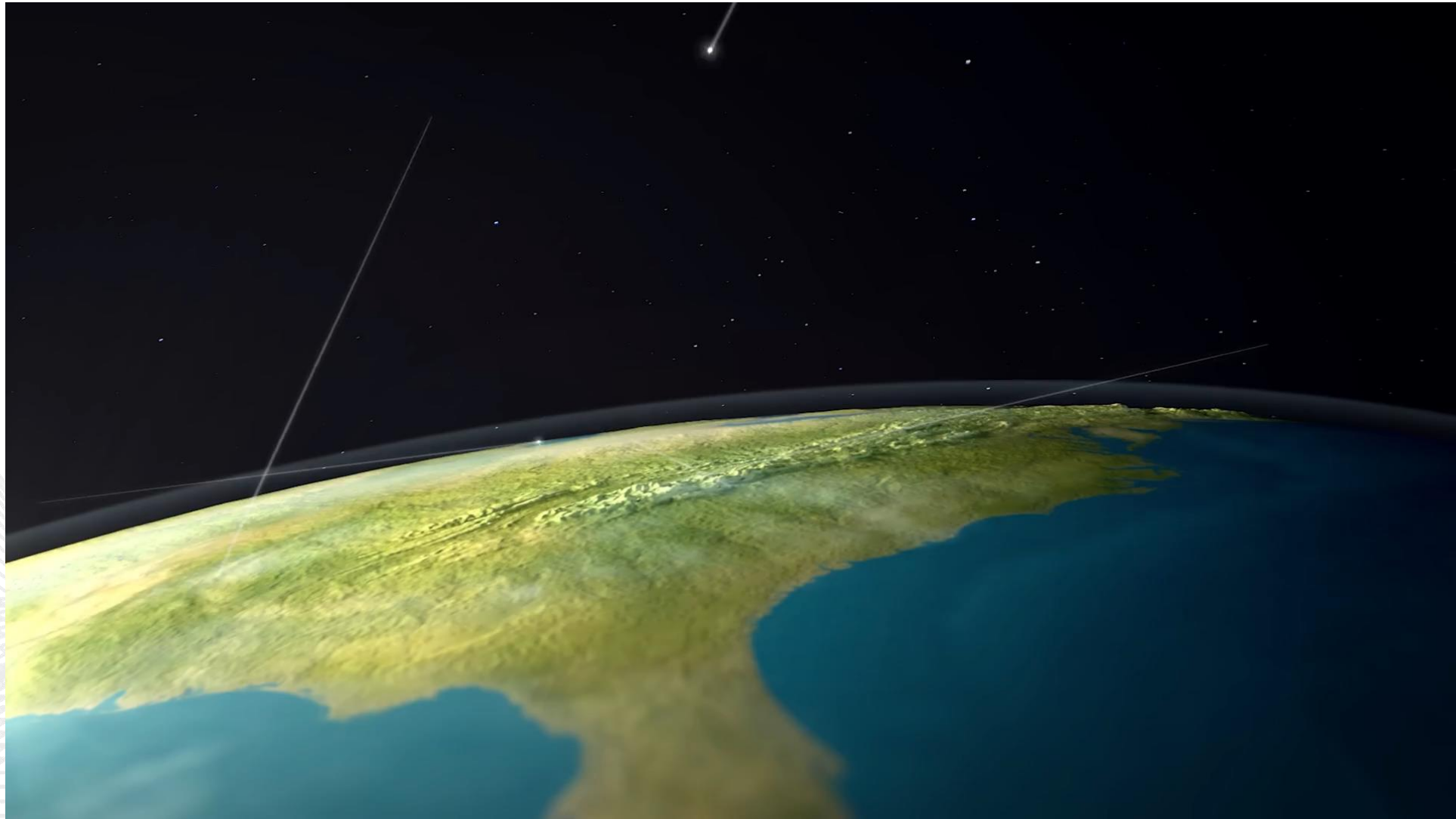


FIGURE 18 – Schéma expliquant le principe de coïncidence entre deux détecteurs à rayons



FIGURE 26 – Photos de prises de mesure.

Research Application



The Absurd
Search For
Dark Matter

Hands-on

SP5620CH – Cosmic Hunter



NEW Educational kits dedicated to the Cosmic Rays detection



SP5620CH - Cosmic Hunter



SP5622B - Detection System Plus

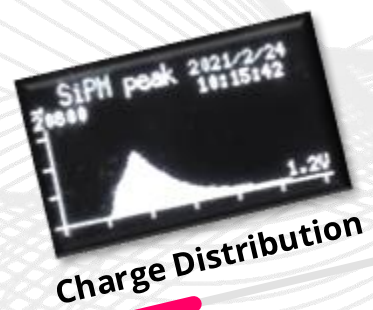
Coming SOON

Detection System Plus – SP5622B

Compact solution for Cosmic Rays Detection!

- Standalone
- External Trigger system for laboratory setups
- Fully compatible with Cosmic Hunter
- Based on SiPM detector and plastic scintillating tile
- Analog and Digital Outputs
- No Needs SW interface
- SD card to download data

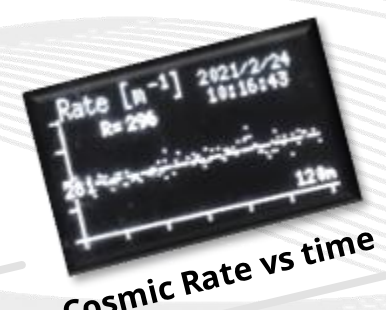
Coming SOON



Charge Distribution



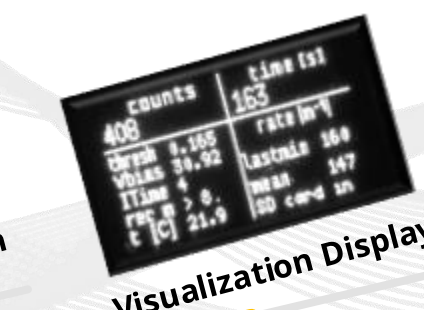
Time Distribution



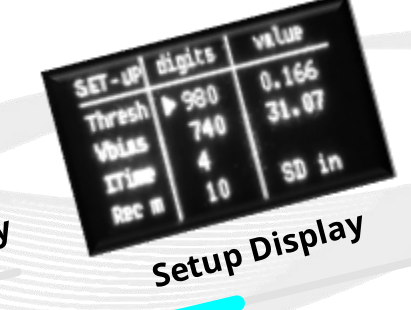
Cosmic Rate vs time



Mean Rate Histogram



Visualization Display



Setup Display

Thank you!



Spares Slides

SP5630EN

Samples

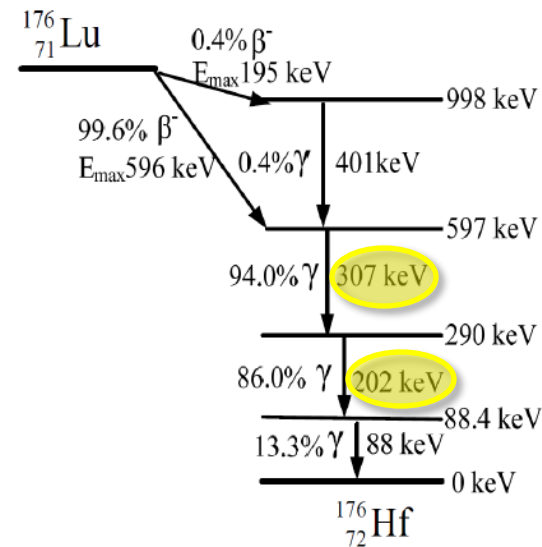
LYSO (Lu_{1.8}Y₂SiO₅:Ce) (Cerium-doped Lutetium Yttrium Orthosilicate) Scintillating Crystal

Scintillator based on Lutetium (Lu) like **LSO** and **LYSO (Lu_{1.8}Y₂SiO₅:Ce)** are usually used in PET applications thanks to their high stopping power (high Z), high light yield and very short decay time (very fast signals). It is a non-hygroscopic scintillator.

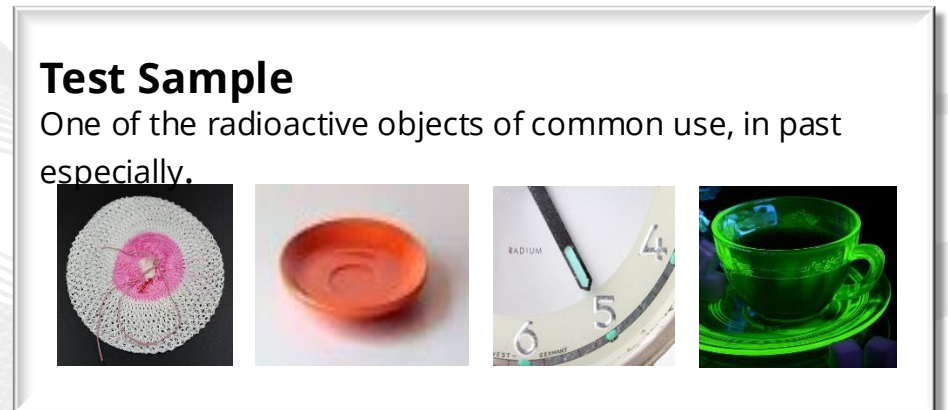
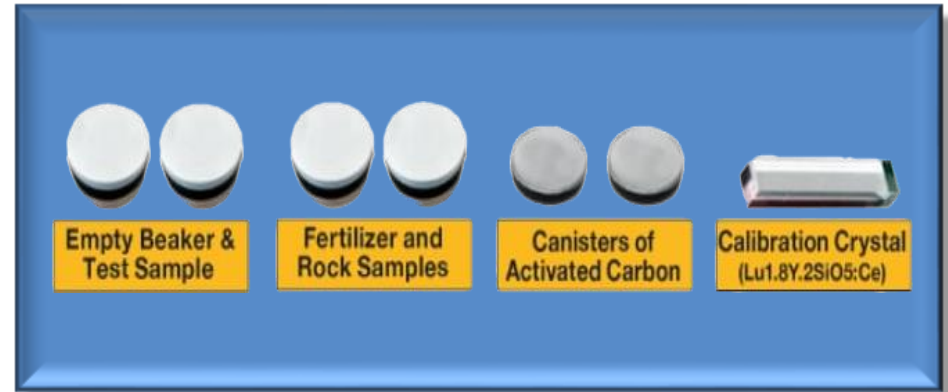


2,6% of the natural Lu is ¹⁷⁶Lu, a radioisotope with a long half life decaying via two different beta decays followed by gamma emissions.

$$T_{1/2} \sim 3,6 \times 10^{10} \text{ y}$$



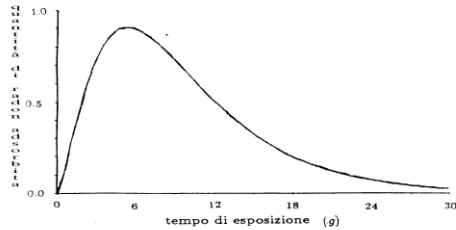
Properties	Value
Cleavage Planes	None
Decay Constant (ns)	40
Density (g cm ⁻³)	7.1
Emission Spectral Range (nm)	380-480
Melting Point (K)	2323
Peak Scintillation Wavelength (nm)	420
Photons/MeV	32000
Radiation Length (cm)	1.15
Refractive Index at Peak	1.81
Emission	
Solubility (g/100g H ₂ O @ 300K)	Insoluble
Stability	Good
Structure	Cubic



Samples (2)

Canisters of Activated Carbon for Radon Passive Measurements

- Diameter: 10 cm
- Height: 3 cm
- Content: 70 g of activated carbon



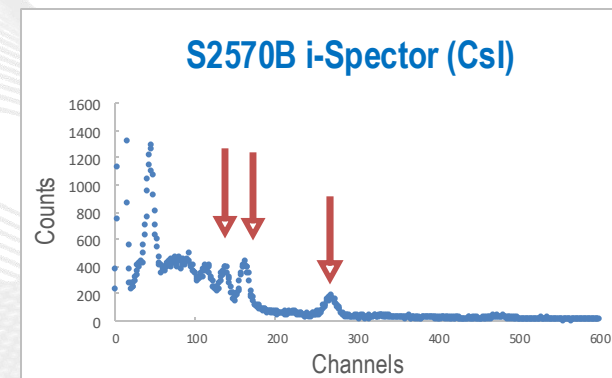
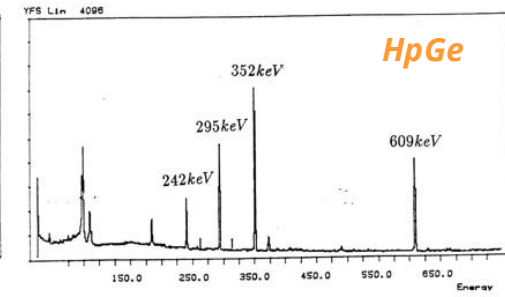
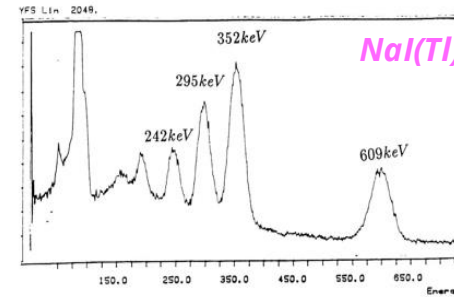
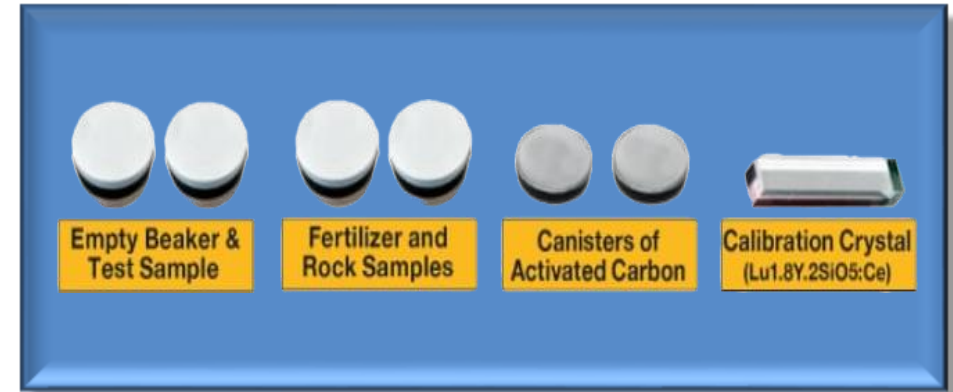
The activated carbons are enclosed in metal containers called "canisters". Covered by a thin double-mesh metal mesh (diffusive barrier). The diffusion barrier serves to eliminate the air flows inside the basket, which can favor the re-emission of radon.

The method consists in carrying out gamma spectrometry measurements on the baskets after the radon has been adsorbed by them!

After 6-7 days, the loss due to decay prevails over the accumulation by adsorption.

Features:

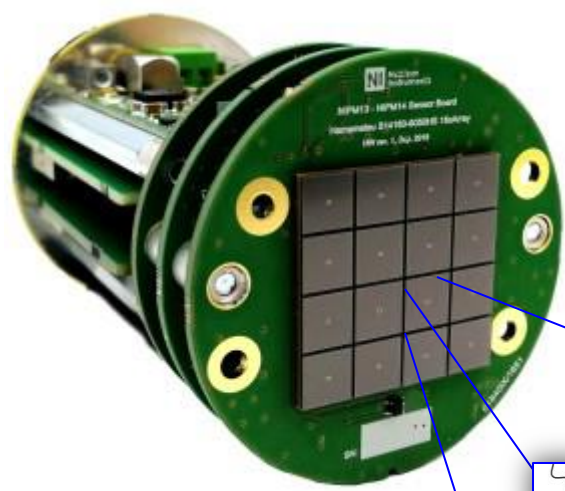
- ❖ useful for short-term measurements: 2 - 7 days
- ❖ strong dependence of the response on humidity



Gamma Energy lines:

- 295 keV and 352 keV from ^{214}Pb
- 609 keV from ^{214}Bi

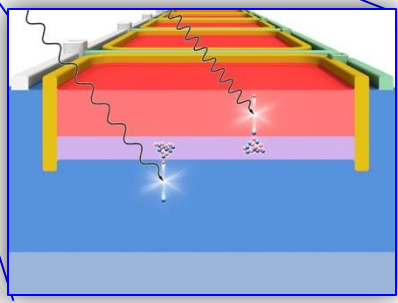
i-Spector Digital [S2570B-S2570D]



- The systems are based on a SiPM area 18×18 mm²
Hamamatsu S14160-60520HS

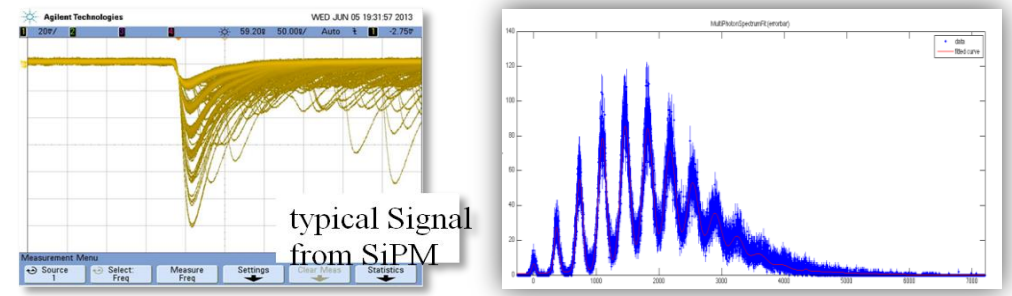
All SiPMs of the area are connected in parallel to increase the active area of the matrix.

Silicon Photomultiplier (SiPM) are high density (up to 10⁴/mm²) matrix of diodes with a common output, working in Geiger-Müller regime. Each cell is a pixel with a binary signal. It can detect a single photon!



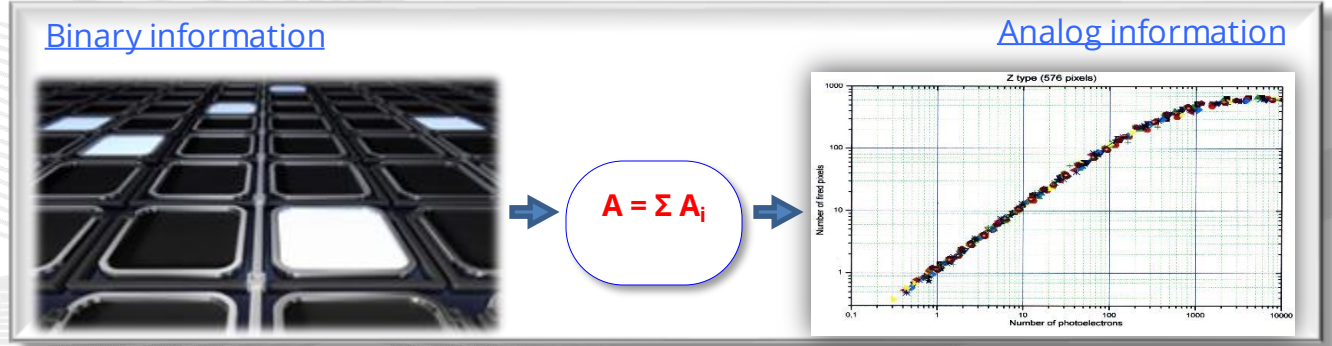
- High Gain
- Low Voltage
- High photon number resolving power
- Wide dynamic range
- Good timing capability
- Low cost
- Withstanding to magnetic field

The high uniformity of pixel structure guarantees no avalanche fluctuations

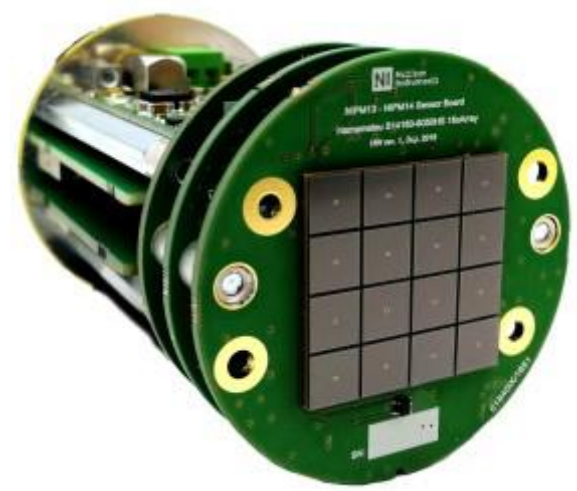


Linear response if the average number of photoelectrons/pixel is less than one

Number of pixel determines the SiPM **dynamic range**



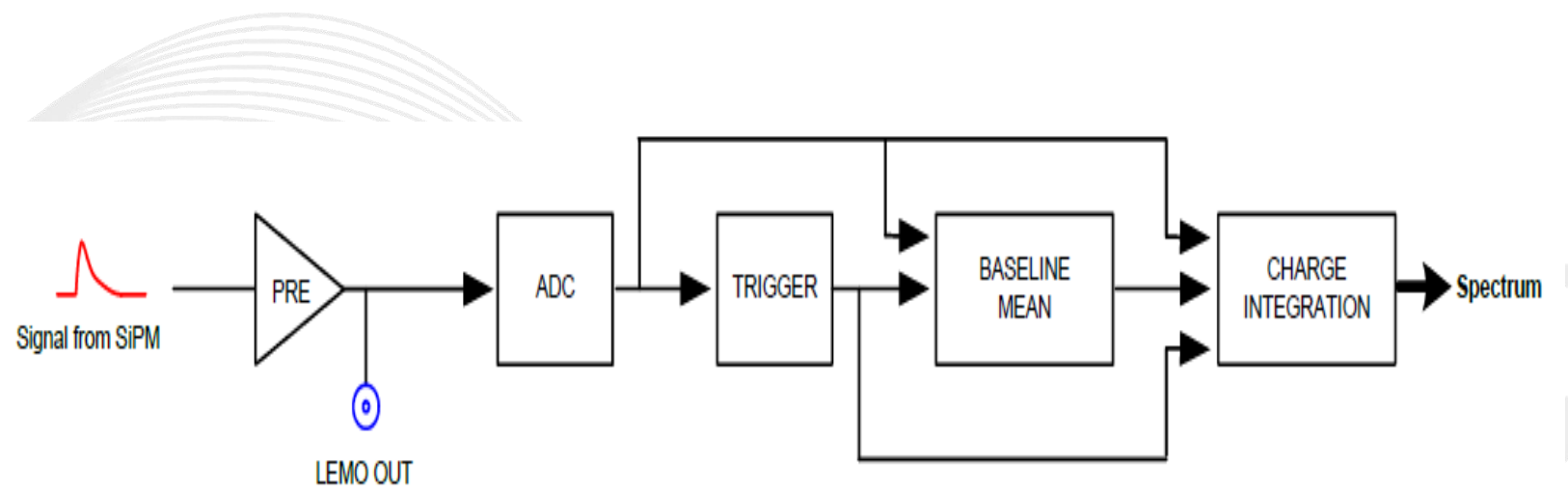
i-Spector Digital [S2570B-S2570D]



- The systems are based on a SiPM area $18 \times 18 \text{ mm}^2$
Hamamatsu S14160-60520HS
- All SIPMs of the area are connected in parallel to increase the active area of the matrix.
- They integrate a shaper, a peak stretcher and a peak ADC to implement a simple MCA (4K).
- Scintillator Crystal: CsI $18 \times 18 \times 30 \text{ mm}^3$
- Energy Range: 30 keV to 3 MeV
- Energy Resolution (FWHM): $<6 \%$ @ 662 keV
- Connectivity: Ethernet
- Software: Web GUI

TECHNICAL SPECIFICATIONS

Supply Voltage	8-13V (12 V typ.)
Power consumption	3W max.
Preamplifier bandwidth	$>1\text{GHz}$
Preamplifier gain	x 5
Shaping time	180 ns
Output signal	- 4 ... +4 V, 170 mA
HV Power supply	20-80 V (10mA)
HV accuracy	1 mV
Thermal feedback accuracy	$0.01^\circ\text{C} - 1\text{mV}$
MCA nr. of channels	4096



i-Spector Digital block diagram

i-Spector Digital [S2570B-S2570D]



- The systems are based on a SiPM area $18 \times 18 \text{ mm}^2$
Hamamatsu S14160-60520HS
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HV Power supply	20-80 V (10mA)
HV accuracy	1 mV
Thermal feedback accuracy	$0.01^\circ\text{C} - 1 \text{ mV}$
MCA nr. of channels	4096

Web-based GUI for unit control and data analysis

i-Spector Digital can be easily controlled through its dedicated web graphical user interface, with no needs to install a dedicated software. The user can configure the module and visualize the acquired spectrum.

Thanks to the internal circular memory buffer, the last 1-hour recording can then be downloaded by the web interface.



i-Spector Digital [S2570B-S2570D]



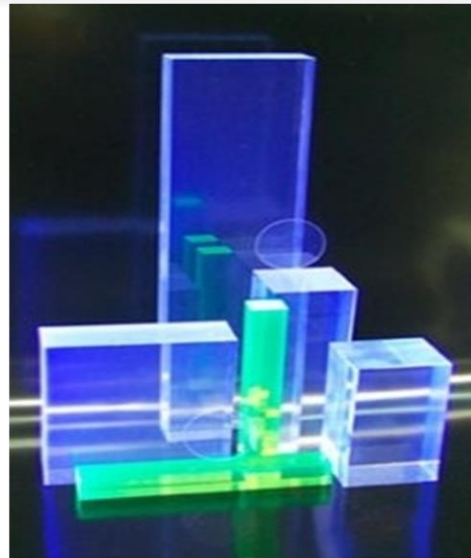
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- Energy Range: 30 keV to 3 MeV
- Energy Resolution (FWHM): $<6 \%$ @ 662 keV
- Connectivity: Ethernet
- Software: Web GUI

CsI(Tl) information:

This scintillator offers a high light yield and emits at a wavelength very suitable for silicon photomultipliers (SiPMs). Typical applications include arrays of this material used in security imaging systems, such as baggage scanners.

BGO information:

A relatively hard, high density, non-hydroscopic crystal with good gamma ray absorption. Often used for PET imaging and high energy physics applications as Compton shields.



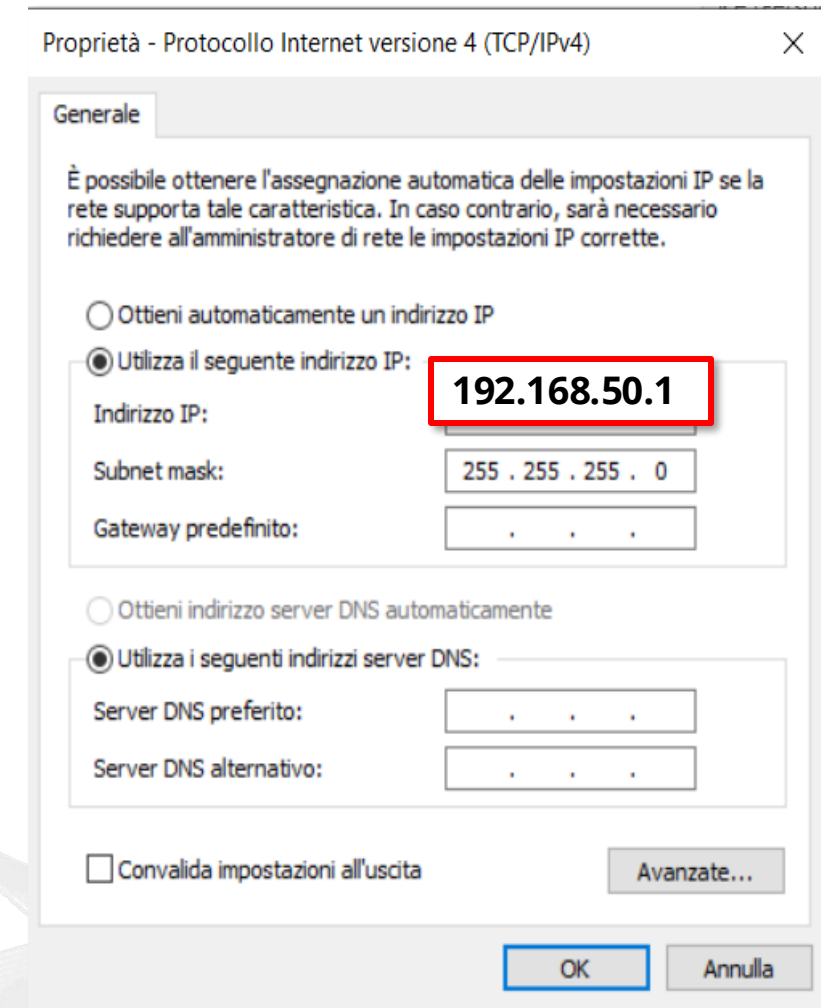
Properties	CsI	BGO
Cleavage Planes	None	None
Decay Constant (ns)	1000	300
Density (g cm^{-3})	4.51	7.13
Emission Spectral Range (nm)	350-725	350-650
Gamma and X-ray absorption coefficients (cm^{-1})	0.48 at 660keV 10.00 at 100KeV	-
Melting Point (K)	894	1323
Peak Scintillation Wavelength (nm)	550	480
Photons/MeV	52000	8500
Radiation Length (cm)	1.86	1.13
Refractive Index at Peak Emission	1.78	2.15
Solubility (g/100g H_2O @ 300K)	44.0	Insoluble
Stability	Slightly Hygroscopic	Good
Structure	BCC	Cubic
Thermal Conductivity ($\text{W}\cdot\text{m}^{-1}\cdot\text{K}^{-1}$) @ 300K	1.13	-
Transmission Range (nm)	240-70000	470-7500

Environmental Kit QuickStart

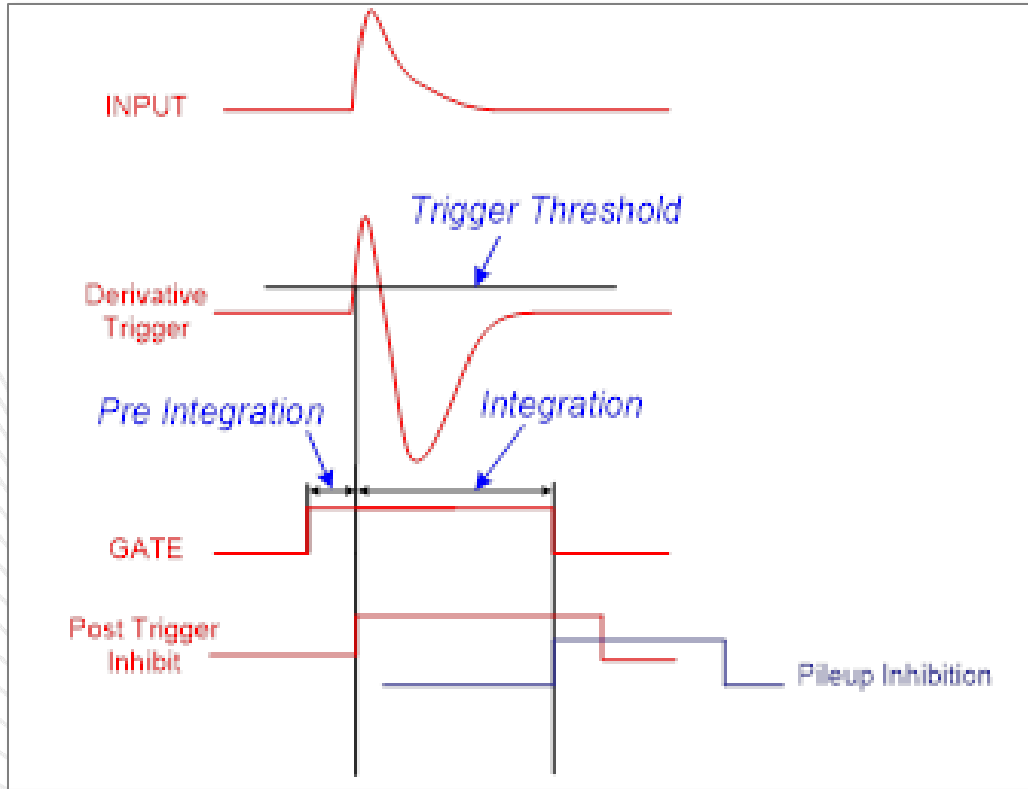
- 1. Unboxing and Assembling**
- 2. Software Setup**
3. Energy calibration of the system based on LYSO crystal (time base = 10')
4. Calibration verification and tuning with Potassium Chloride sample (time run = 30')
5. Background measurement (time run = 30')
6. Rock sample Spectrum (time run = 30')
7. Test sample radiation identification
8. Analysis of spectra and superposition
9. Passive Radon Measurements

Software Setup

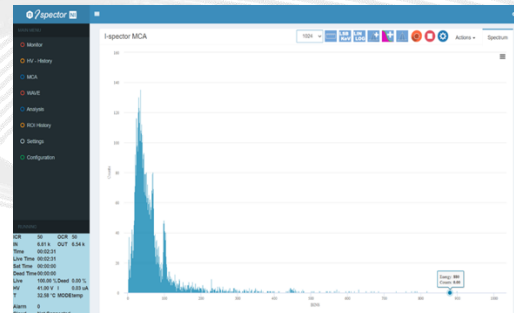
- **IP address** of the i-Spector for Ethernet connection is **192.168.50.2**
- Configure the Ethernet network of the PC from the “Network and Sharing Center”
- Open a web browser (**Microsoft Edge browser is suggested**) and enter the web address **192.168.50.2**. The homepage of the graphical web interface will open.



Software Setup - Parameters of the charge integration algorithm

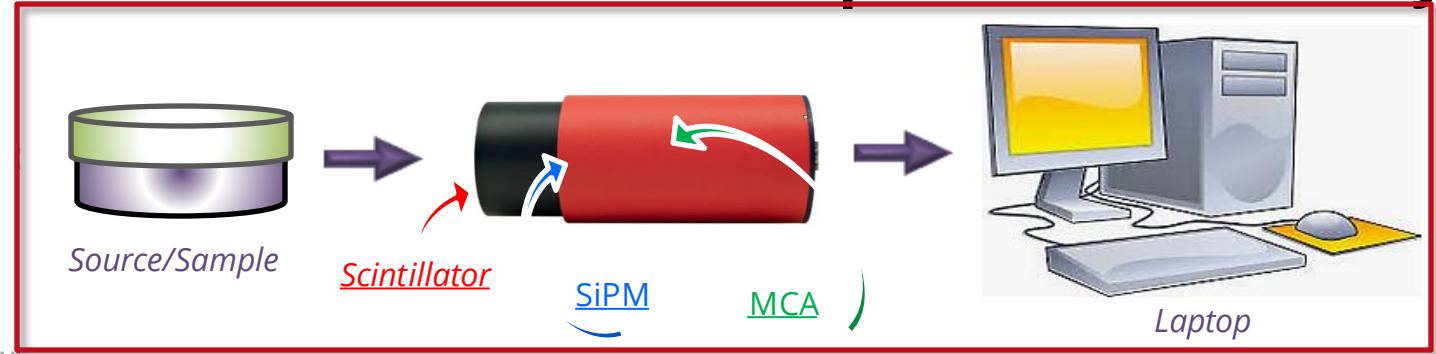


- **Trigger Threshold (LSB):** threshold for the derivative trigger
- **Post Trigger Inhibit (ns):** set the time after a trigger for which any other trigger is inhibited
- **Pre-Integration (ns):** set how much time before the trigger the charge integration is started
- **Integration (us):** set the charge integration gate
- **Gain:** set the energy digital gain to be applied to the spectrum
- **Pileup Inhibition (us):** set the time after the integration gate for which the acquisition of any other event acquisition is inhibited
- **Pileup Penalty (us):** set the trigger inhibition gate to be opened after a pile up
- **Baseline Inhibition (us):** set the time after the integration gate for which the baseline is not calculated
- **Baseline Length (samples):** set the number of samples used to calculate the baseline
- **Target Mode:** set the acquisition mode as Free Running or with a target in Time (ms) or Counts
- **Target Value:** set the target value in time or counts, accordingly to the Target Mode

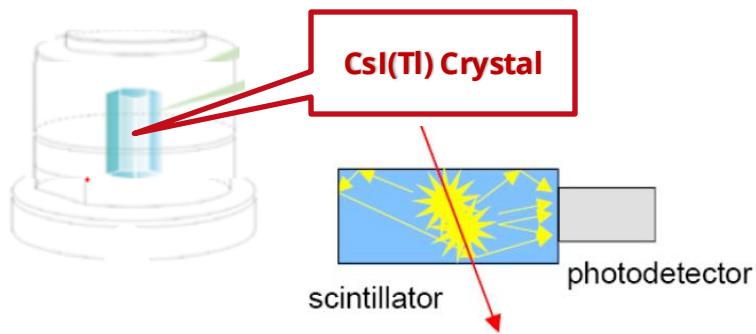


Experimental activity

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO
- 3) Background measurement (time run = 30')
- 4) Calibration verification and tuning with Potassium



Scintillator



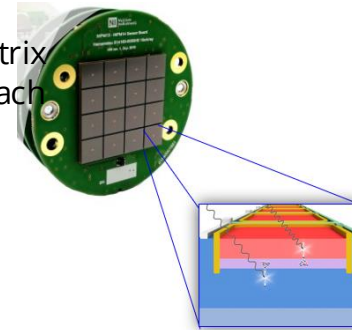
Energy deposition by an ionizing particle:

- Generation of light
- Transmission of scintillation light
- Detection

CsI(Tl) has a light output of 54 photons/keV and average decay time of about 1 μ s for γ -rays

Silicon Photomultiplier (SiPM)

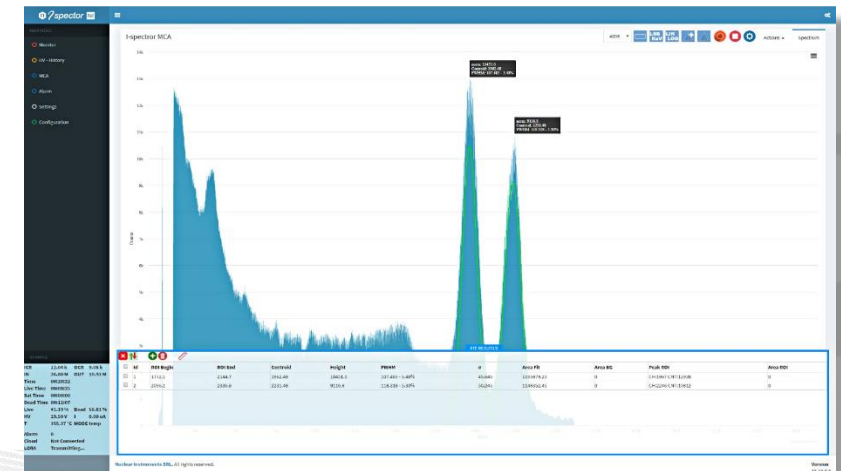
Silicon Photomultiplier (SiPM) is detector made of a matrix of silicon cells (diodes). Each diode is a pixel with a binary signal. It can detect a single photon!



Photosensors detect and transform the light produced by the scintillator into an **electrical signal**. This signal is proportional to the energy released inside the crystal by the interacting particle

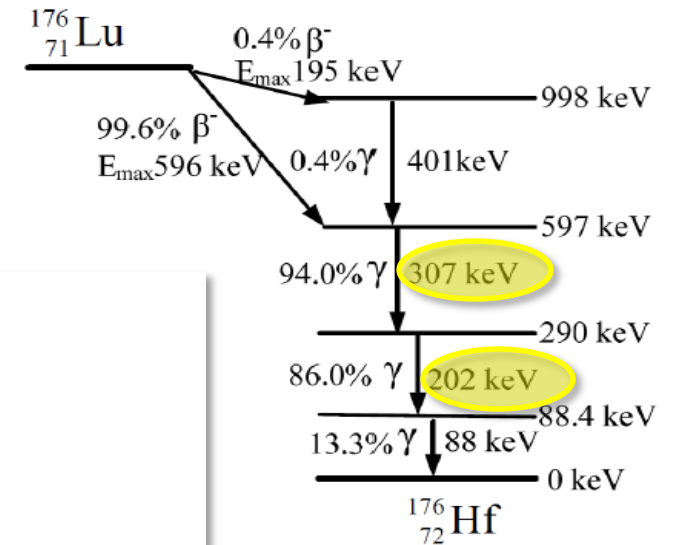
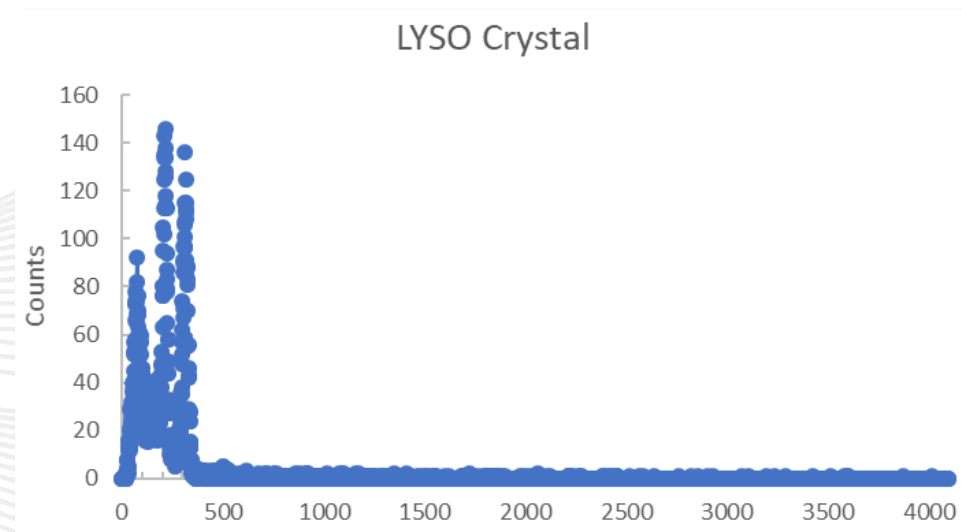
Electronics & Analysis Software

The output pulses are proportional to the energies of the incident radiation, the ADC is used combined to a Multichannel Analyzer (MCA) to generate energy distributions (spectra) of radioactive samples



Experimental activity

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO crystal (time base = 10')**
- 3) Background measurement (time run = 30')
- 4) Calibration verification and tuning with Potassium Chloride sample (time run = 30')
- 5) Rock sample Spectrum (time run = 30')
- 6) Test sample radiation identification
- 7) Analysis of spectra and superposition
- 8) Passive Radon Measurements

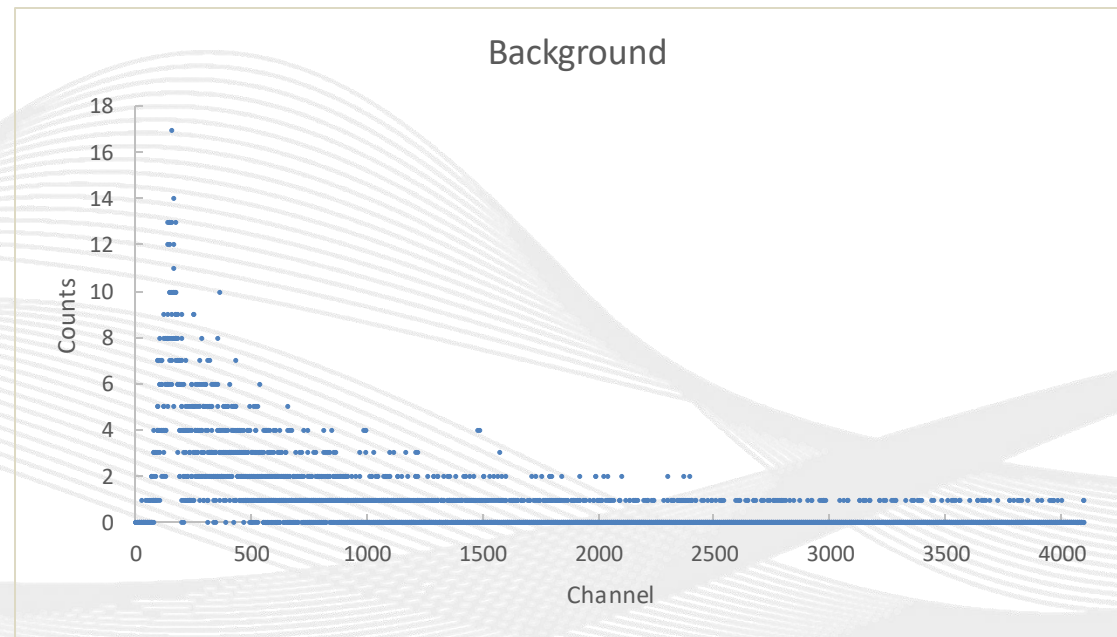


Scintillator based on Lutetium (Lu) like LYSO (Lu_{1.8}Y_{0.2}SiO₅:Ce) has an high stopping power (high Z), high light yield and very short decay time (very fast signals).

2,6% of the natural Lu is ¹⁷⁶Lu, a radioisotope with a long half life decaying via two different beta decays followed by gamma emissions.

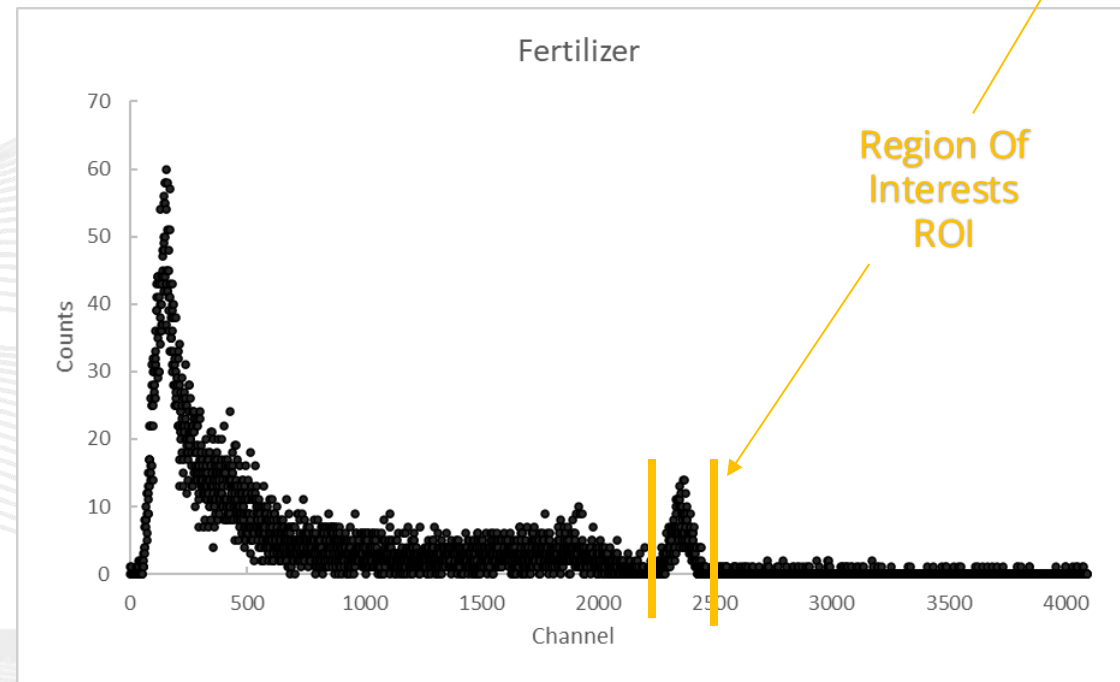
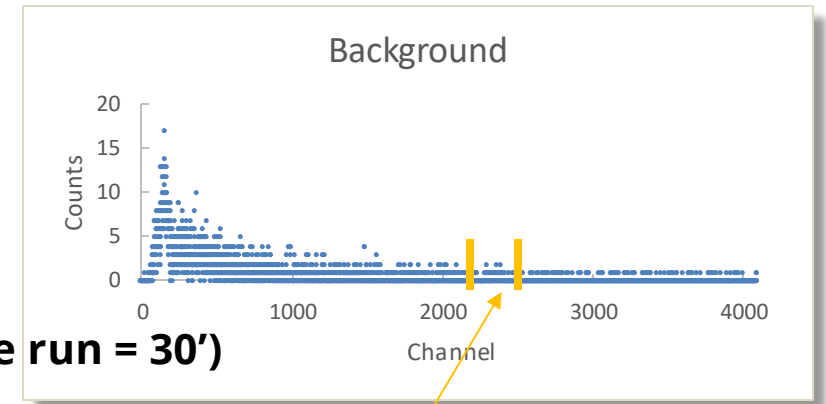
Experimental activity

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO crystal (time base = 10')
- 3) Background measurement (time run = 30')**
- 4) Calibration verification and tuning with Potassium Chloride sample (time run = 30')
- 5) Rock sample Spectrum (time run = 30')
- 6) Test sample radiation identification
- 7) Analysis of spectra and superposition
- 8) Passive Radon Measurements



Experimental activity

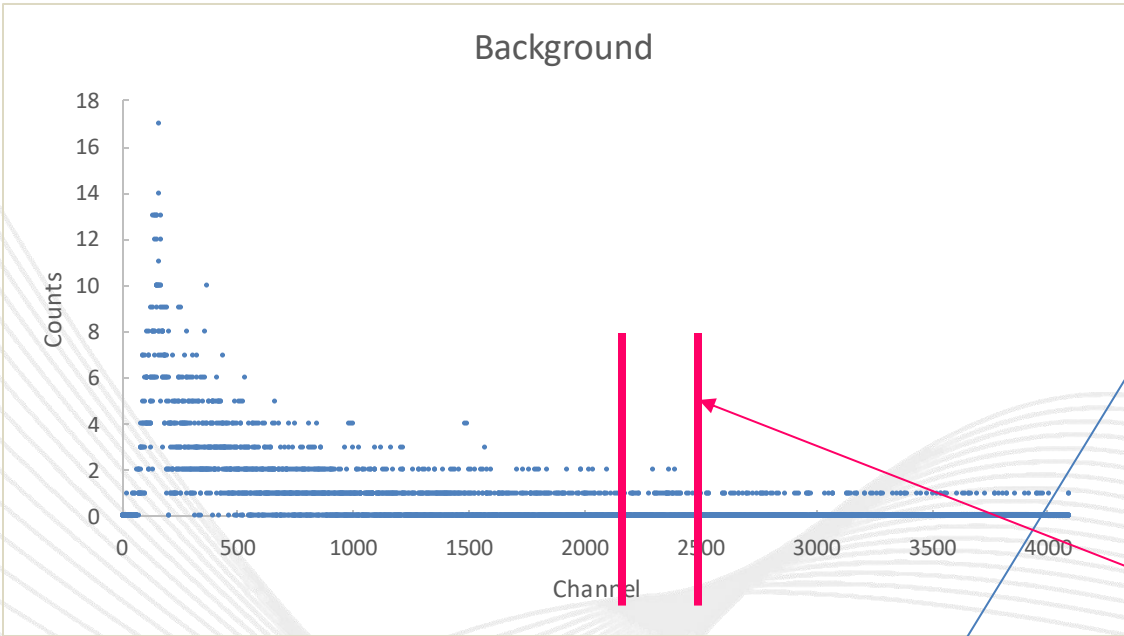
- 1) How to use the i-Spector Digital
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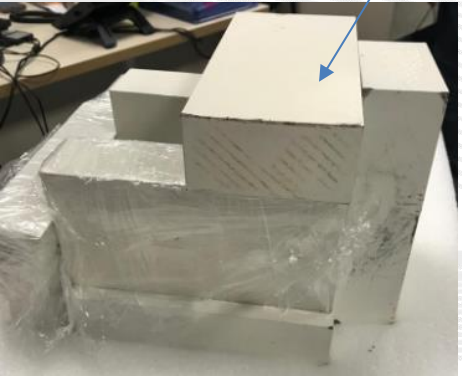
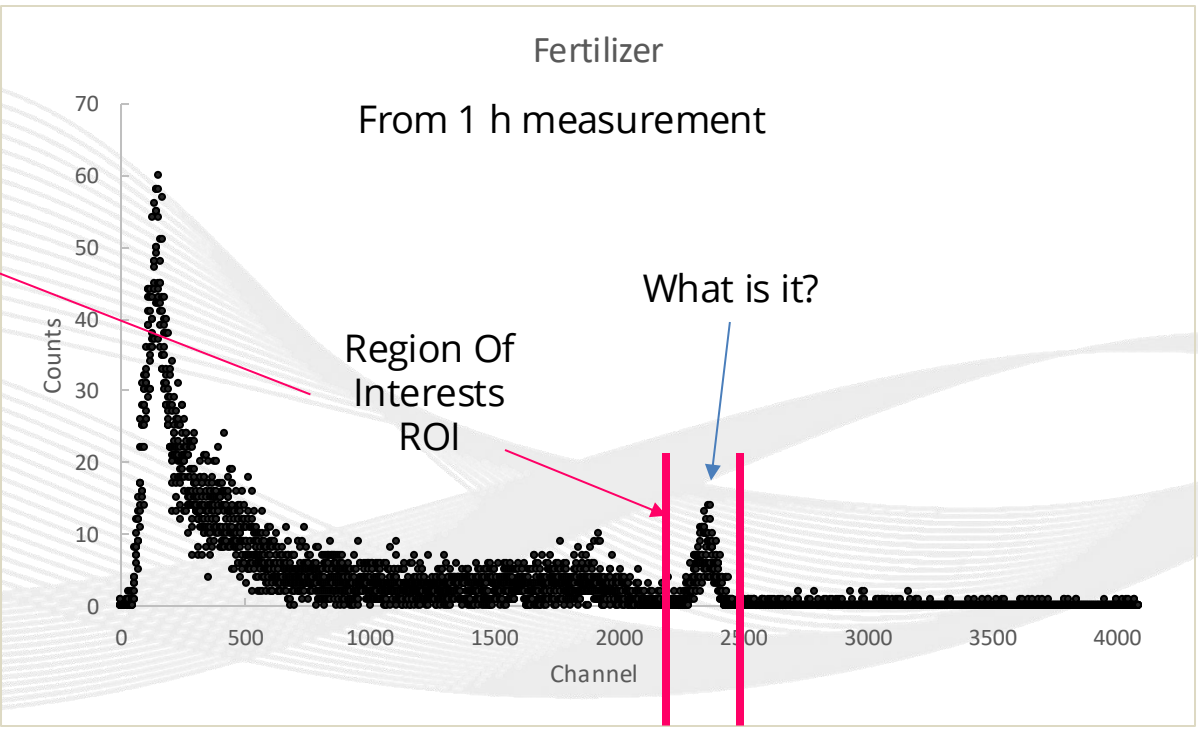
Experimental activity

Fertilizer – Potassium Chloride

For our hands on take a 30 min spectra with the Potassium Chloride sample



Fertilizer

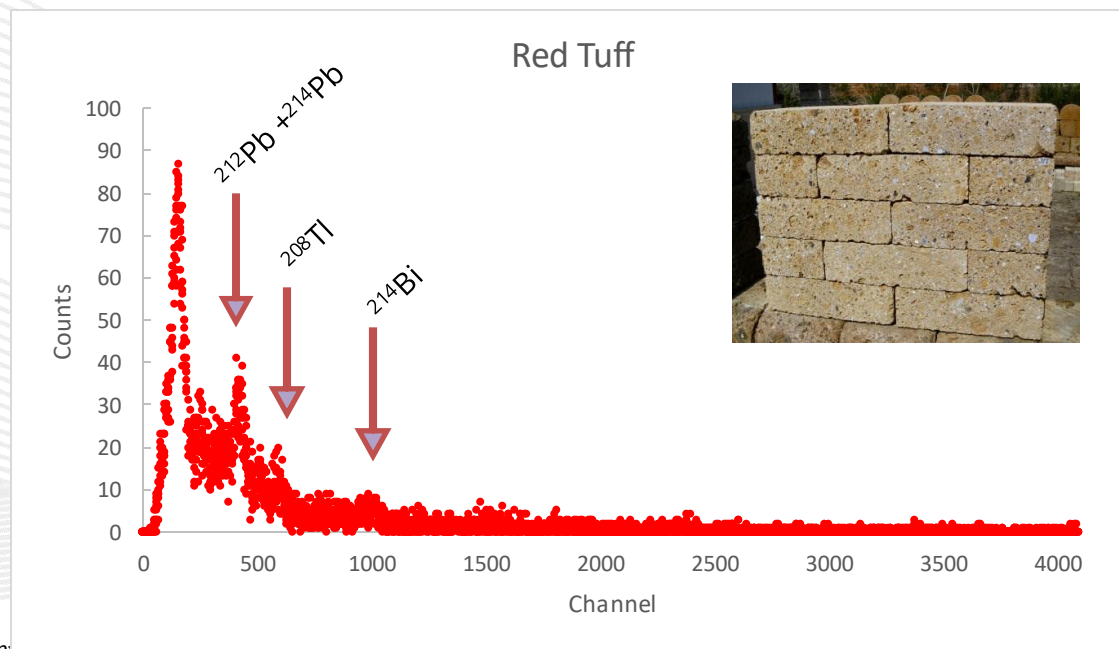


Experimental activity

U-238

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO crystal (time base = 10')
- 3) Background measurement (time run = 30')
- 4) Calibration verification and tuning with Po
- 5) Rock sample Spectrum (time run = 30')**
- 6) Test sample radiation identification
- 7) Analysis of spectra and superposition
- 8) Passive Radon Measurements

Th-232		U-238	
99.98%	0.02%	26.8 minutes	53.2 (1.1%)
Pb214			242.0 (7.46%)
			295.2 (19.2%)
			351.9 (37.1%)
			785.9 (1.09%)
	At218	2 seconds	
Bi214		19.7 minutes	609.3 (46.1%)
			768.4 (4.89%)
			806.2 (1.23%)
			934.1 (3.16%)
Bi212	60.6 minutes		1120.3 (15.0%)
			1238.1 (5.92%)
			1377.7 (4.02%)
64.06%	35.94%		1408.0 (2.48%)
Po212	304 nsec		1509.2 (2.19%)
	Tl208	3.1 minutes	1764.5 (15.9%)
			277.4 (6.31%)
			510.77 (22.6%)
			583.2 (84.5%)
			763.1 (1.81%)
			860.6 (12.4%)



Experimental activity

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO crystal (time base = 10')
- 3) Background measurement (time run = 30')
- 4) Calibration verification and tuning with Potassium Chloride sample (time run = 30')
- 5) Rock sample Spectrum (time run = 30')
- 6) Test sample radiation identification**
- 7) Analysis of spectra and superposition
- 8) Passive Radon Measurements



*Thorium
Lantern Mantle*



*Rare earth
uranium oxide*



Uranium Glazed Pottery



Uranium glass beads

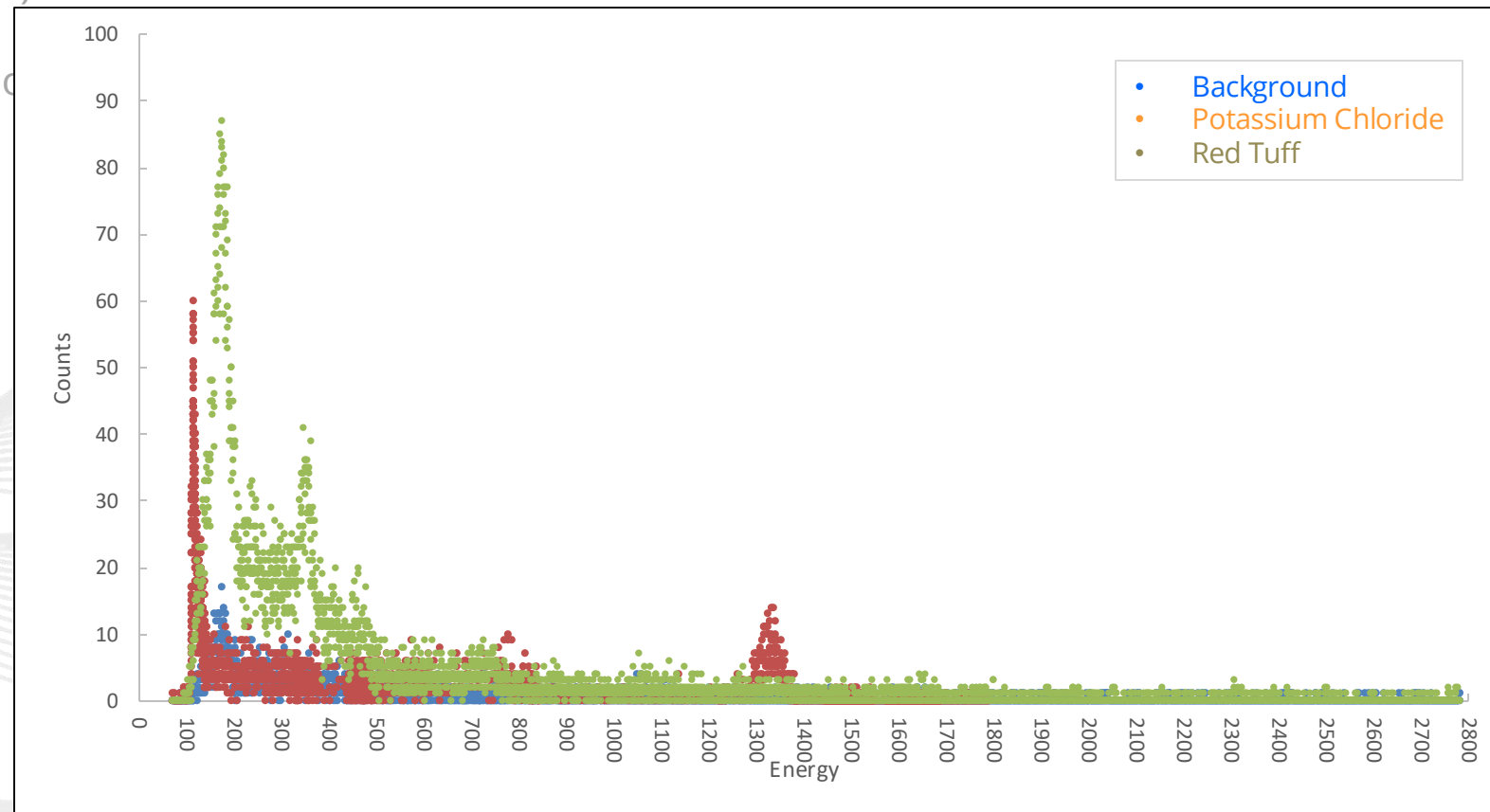


*Decades-Old
Lenses*

Experimental activity

Superimposed spectra

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO crystal (time base = 10')
- 3) Background measurement (time run = 30')
- 4) Calibration verification and tuning with Po
- 5) Rock sample Spectrum (time run = 30')
- 6) Test sample radiation identification
- 7) Analysis of spectra and superposition**
- 8) Passive Radon Measurements



Experimental activity



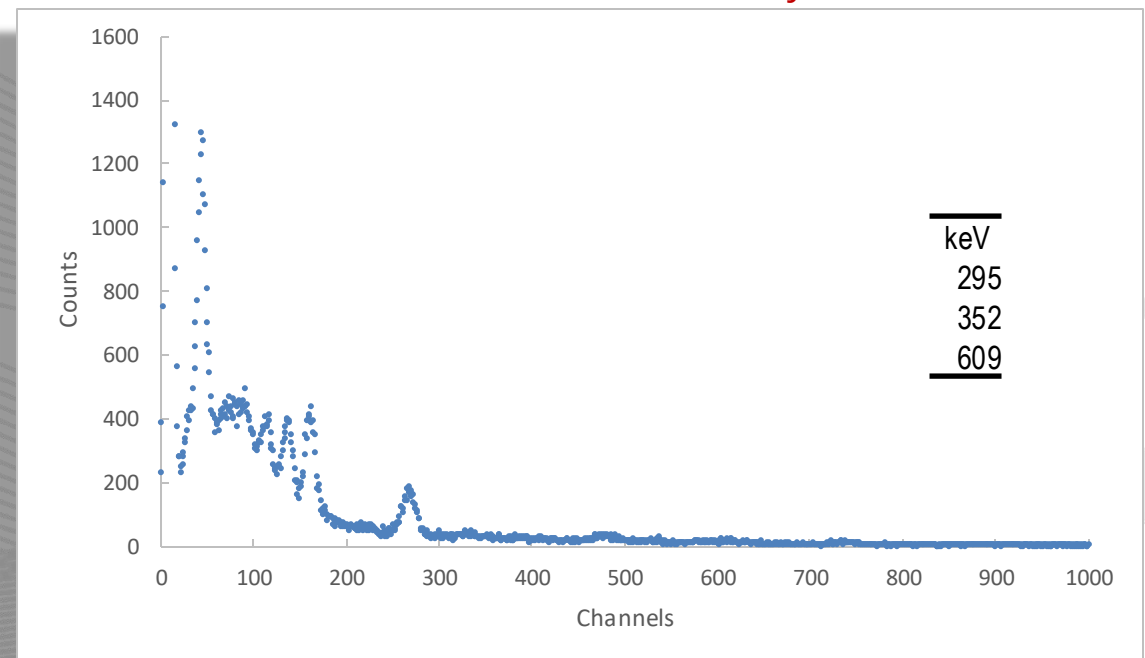
Canisters of Activated Carbon

- 1) How to use the i-Spector Digital
- 2) Energy calibration of the system based on LYSO crystal (time base = 10')
- 3) Background measurement (time run = 30')
- 4) Calibration verification and tuning with Potassium Chloride sample (time run = 30')
- 5) Rock sample Spectrum (time run = 30')
- 6) Test sample radiation identification
- 7) Analysis of spectra and superposition
- 8) Passive Radon Measurements (time run=60')**

The amount of adsorbed radon by the canisters can be evaluated via the detection of the gamma rays emitted by the ^{214}Pb e dal ^{214}Bi .

Among the many available gamma emissions, the following nuclei are used as they are formed in a short time from the decay of Radon:
 ❖ 295 keV and 352 keV from ^{214}Pb
 ❖ 609 keV from ^{214}Bi

Measurement of the adsorbed radon activity



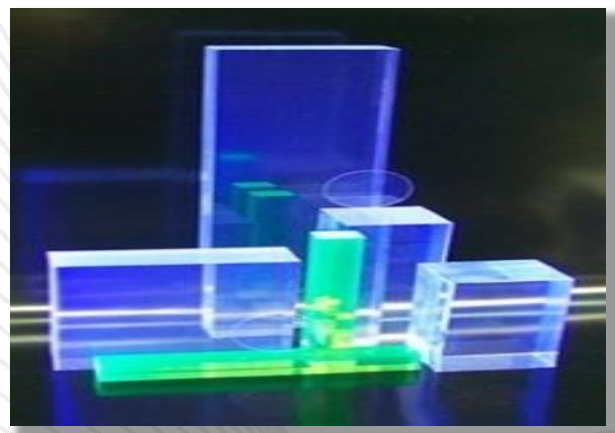
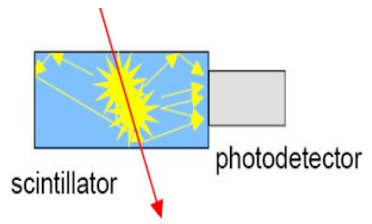
Spares Slides

SP5640

GammaEDU Description - Scintillator

Energy deposition by an ionizing particle:

- Generation of light
- Transmission of scintillation light
- Detection

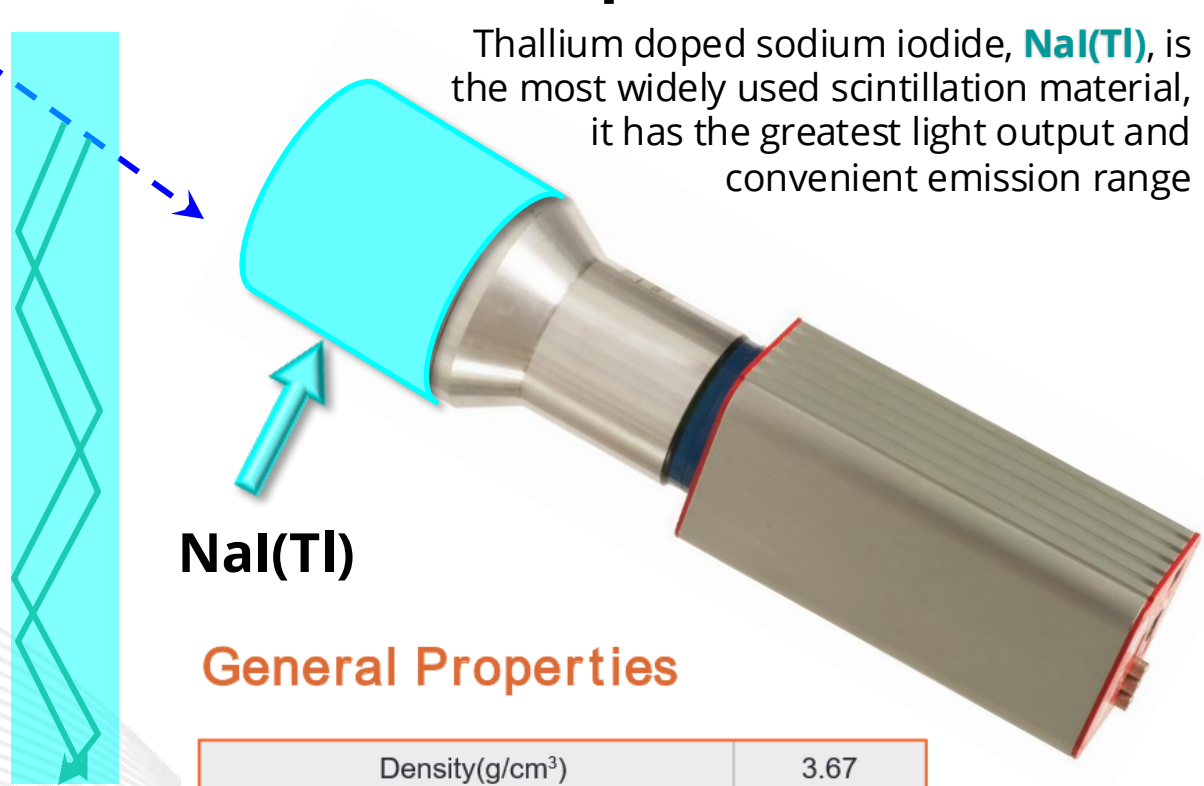


What are scintillators used for?

- To measure the energy released
- To measure the passage time of radiation

Inorganic (crystalline structure)
 Up to 40000 photons per MeV
 High Z
 Large variety of Z and ρ
 Un-doped and doped
 ns to μs decay times
 Expensive

Organic (plastics or liquid solutions)
 Up to 10000 photons per MeV
 Low Z
 ρ~1g/cm³
 Doped, choice of emission wavelength
 ns decay times
 Relatively inexpensive



Thallium doped sodium iodide, **NaI(Tl)**, is the most widely used scintillation material, it has the greatest light output and convenient emission range

NaI(Tl)

General Properties

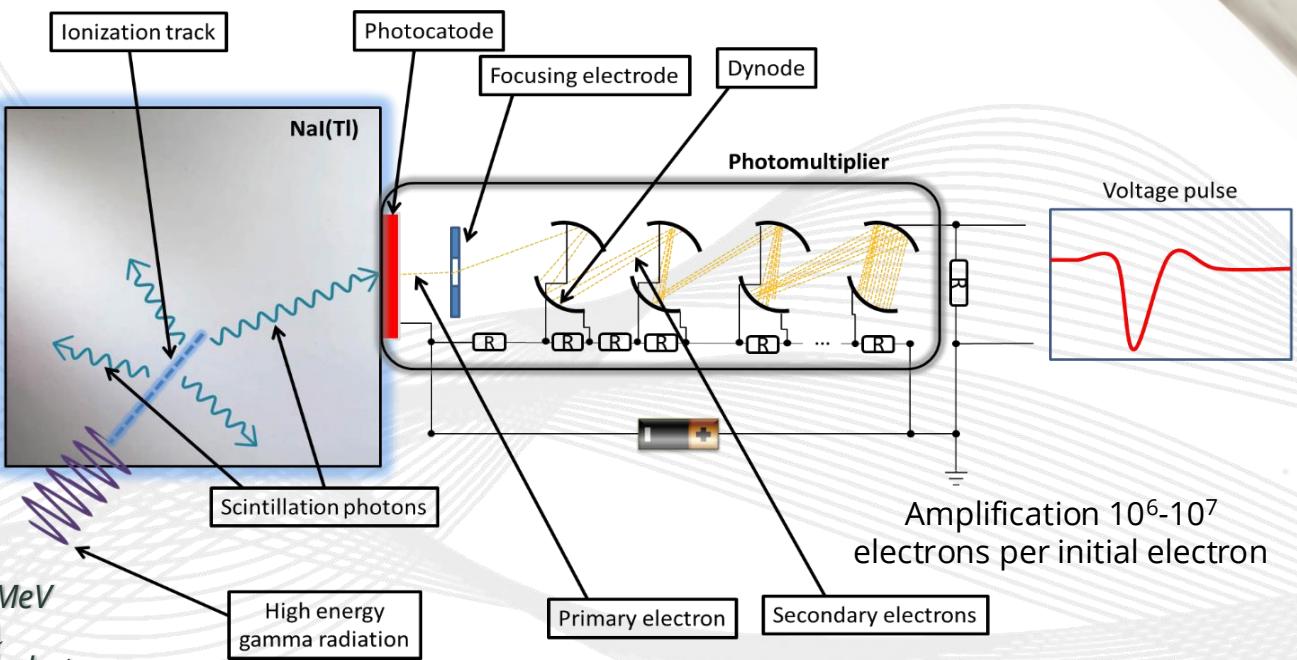
Density(g/cm ³)	3.67
Melting point(K)	924
Wavelength of emission peak(nm)	415
Light output(Photons/MeV)	40,000
Decay time(ns)	264
Cleavage plane	(100)
Hygroscopic	Yes
Refractive index	1.85
Hardness(Mho)	2

GammaEDU Description - Photodetector

Photodetector → From photons to electric current!

Photomultiplier Tubes (PMT) are composed of a photocathode, collection optics and multiplier section. The overall electrical signal is collected at the anode.

❑ *PMT collects and transforms the light produced by the scintillator into an electrical signal*



❑ *The intensity of the output current pulse is proportional to the energy of the incident photon!*



1 γ MeV
↓
 $4.3 \cdot 10^4$ photons

GammaEDU Description - γ stream

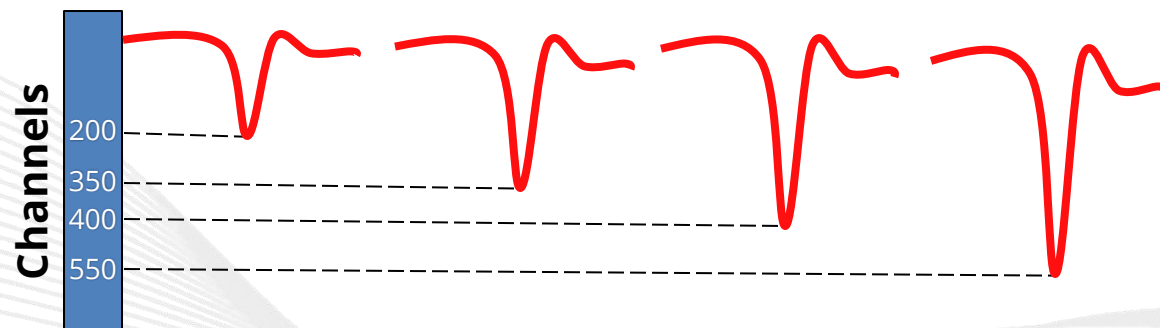
CAEN Gamma stream [S2580] is a compact and portable system for gamma ray spectroscopy with scintillation detectors, which provides an active **Multi-Channel Analyzer** (MCA) integrated in a 14-pin photo-multiplier tube (PMT) base.

Gamma *stream* fully integrates in a stand-alone device the high voltage to bias the PMT, the preamplifier to shape the signal from detector, and the MCA for a complete Pulse Height Analysis online.

Gamma *stream* makes easy the measurements with scintillation detectors **NaI(Tl)** [0.3l] with no need of additional cables.



- High Voltage Power Supply (0 ÷ +1500V/500 μ A)
- Charge Sensitive Preamplifier
- digital Multi-Channel Analyzer (12-bit and 62.5 MHz ADC) for scintillation spectroscopy
- Specialized for NaI(Tl), LaBr₃(Ce), and CeBr₃ with standard 14-pin and 10-8 stages PMTs
- Full stand-alone operation with embedded CPU, data storage (SSD) unit, and power supply for up to 6÷8 hours operation
- Wired and wireless connectivity via USB, Ethernet, Wifi and Bluetooth
- Acquisition modes: PHA, PHA with time stamp, Signal



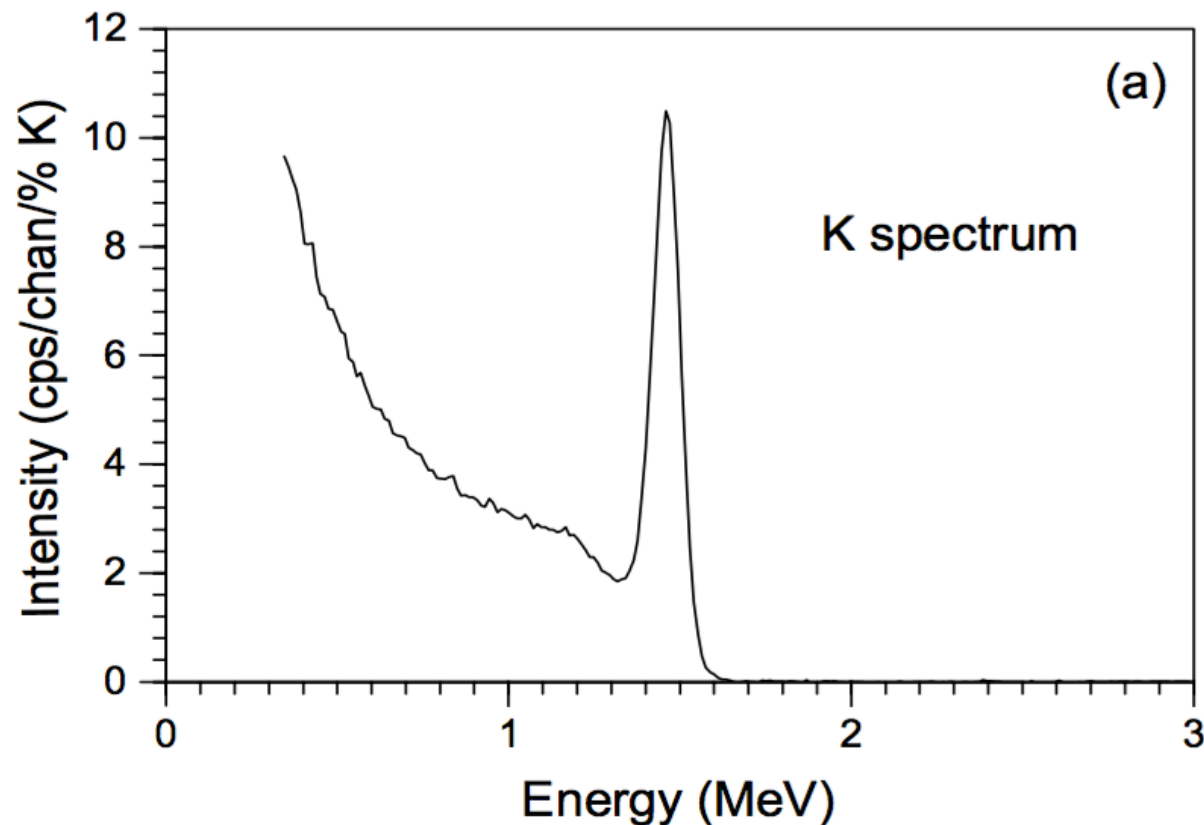
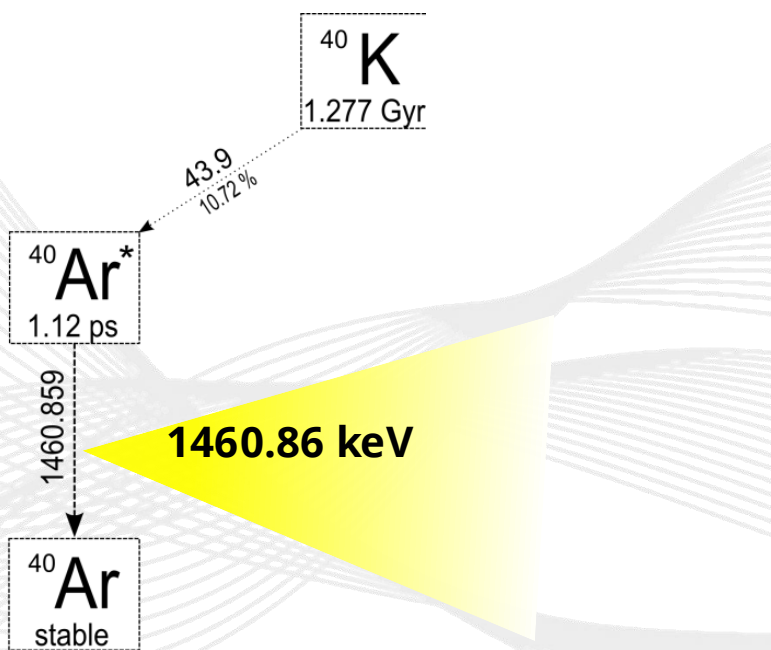
The acquisition channel is proportional to the energy of the incident photons!

Energy Spectrum

The **MultiChannel Analyzer MCA** classifies input pulses base on their height saving them in a memory and are associated to an ADC. The output of every channel can be visualized in a pulse amplitude spectra.

An Analog-to-Digital Converter (ADC) generates a digital signal proportional to the amplitude of an input pulse. Since these output pulses are proportional to the energies of the incident radiation, the ADC can be used combined to a MultiChannel Analyzer (MCA) to generate energy distributions (spectra) of radioactive samples.

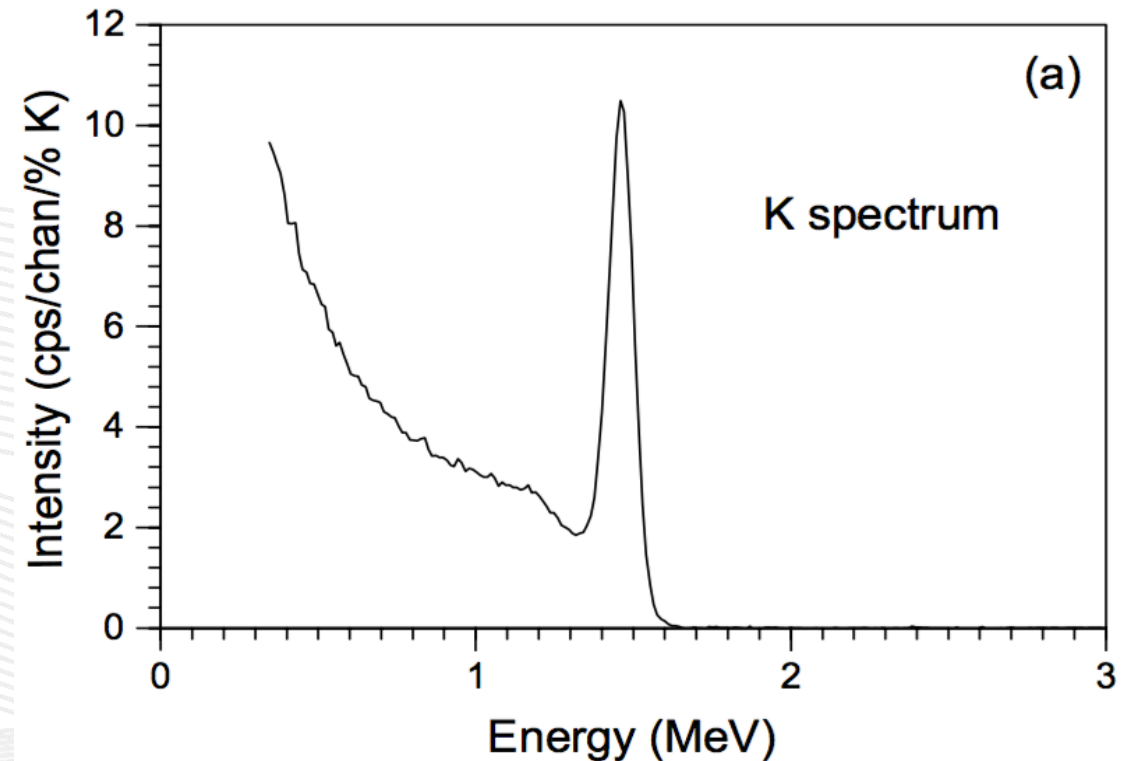
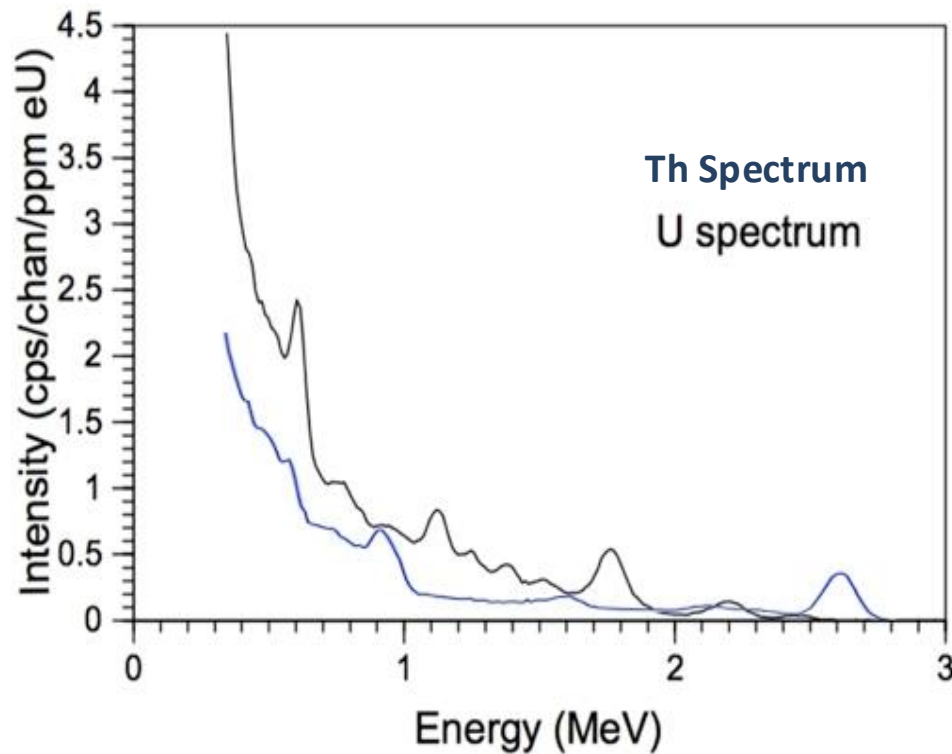
Ex. Monochromatic photons beam



Energy Spectrum

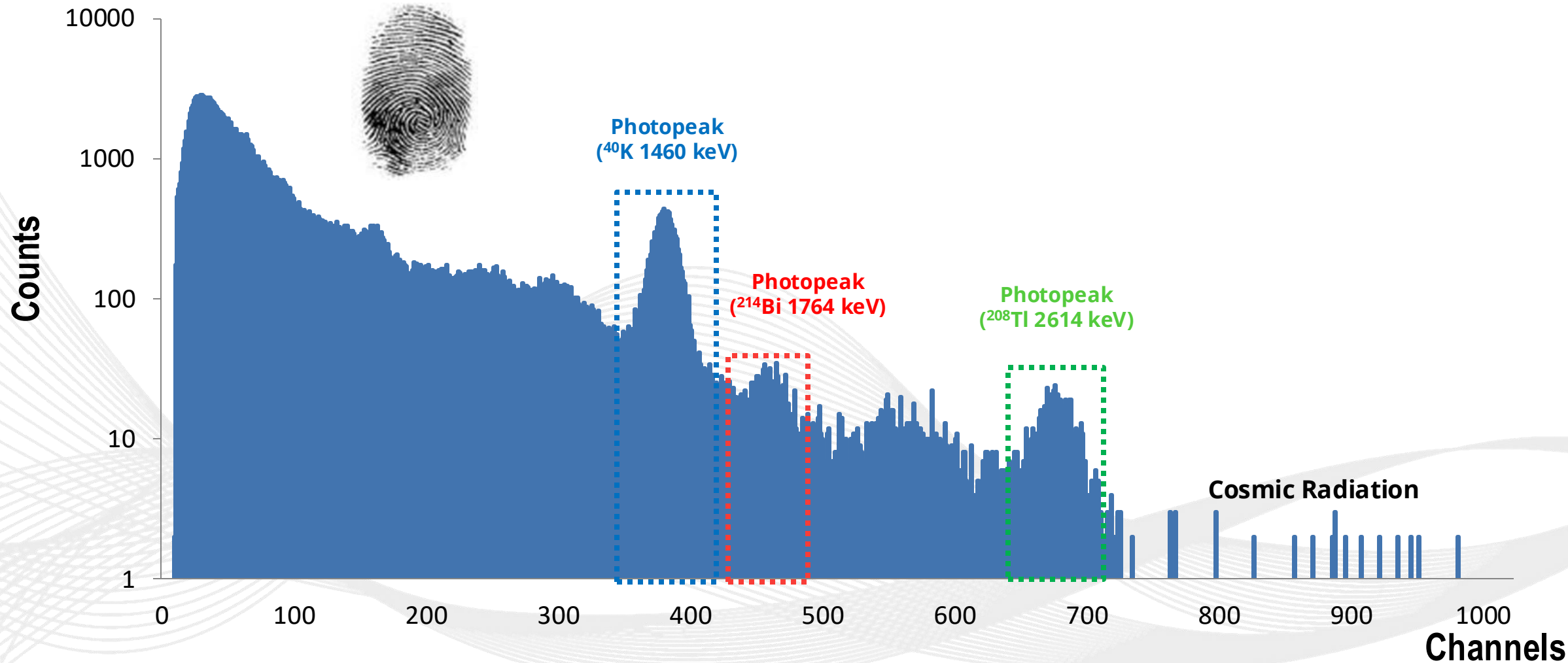
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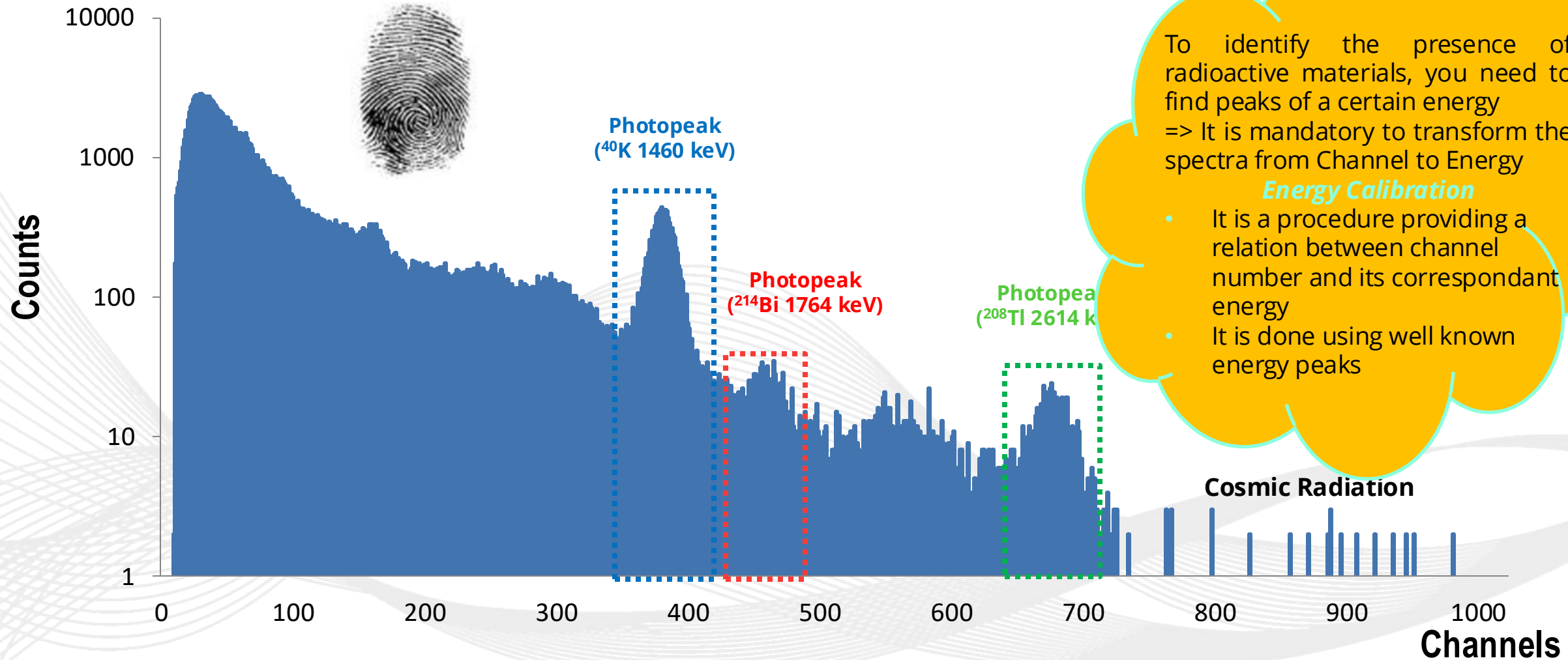
Typical Gamma Spectra

The photopeaks characterize the gamma spectrum. Each photopeak corresponds to the photons coming into detector with an energy value equal to the emission ones. These photons release all their energy into detector.



Typical Gamma Spectra

The photopeaks characterize the gamma spectrum. Each photopeak corresponds to the photons coming into detector with an energy value equal to the emission ones. These photons release all their energy into the detector.



Energy calibration

In the energy range of the environmental measurements the calibration in energy corresponds to a linear transformation

=> Knowing the energy of 2 peaks it is possible to extract the equation of the line from 2 points

$^{40}\text{K} \Rightarrow E_K = 1460 \text{ keV}$
 $^{208}\text{Tl} \Rightarrow E_{Th} = 2614 \text{ keV}$

^{208}Tl is coming from the radioactive chain of the ^{232}Th and is the highest energy gamma from natural sources

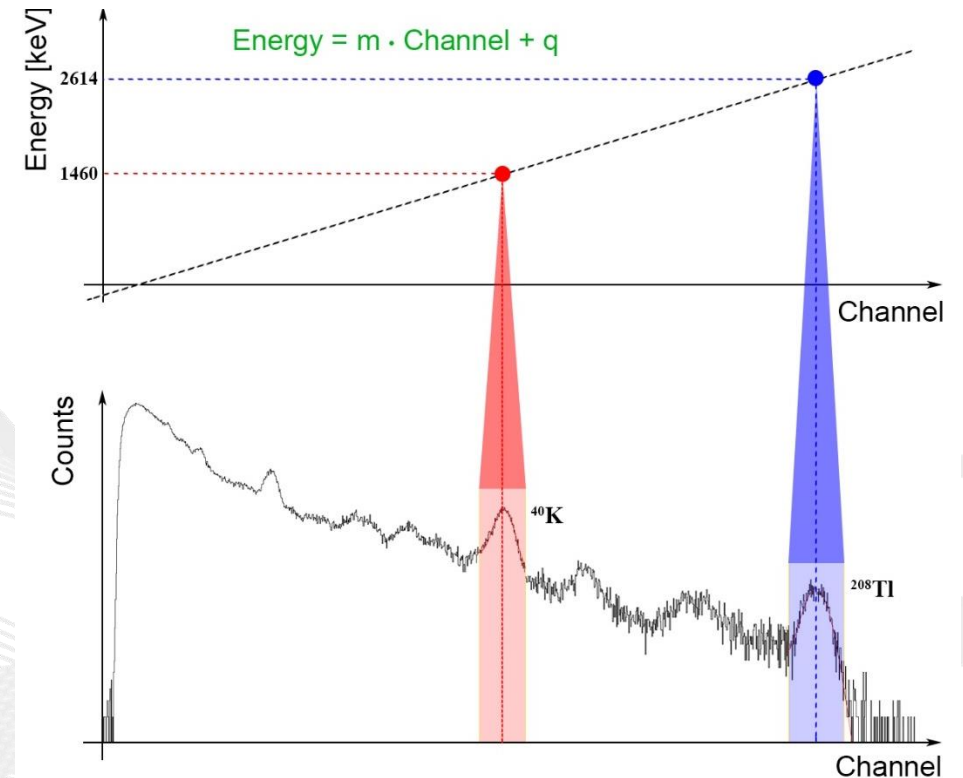
$$A = (E_{Th}, Ch_{Th})$$

$$B = (E_K, Ch_K)$$

$$\frac{Ch - Ch_K}{Ch_{th} - Ch_k} = \frac{E - E_k}{E_{th} - E_k}$$

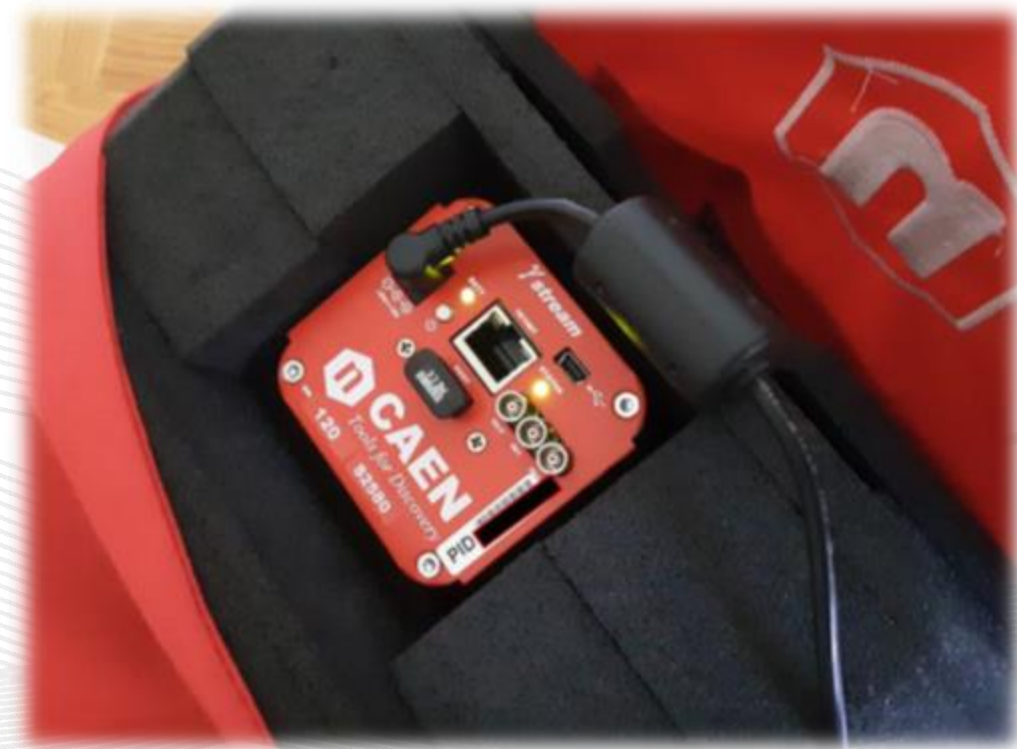
$$\frac{Ch - Ch_K}{Ch_{th} - Ch_k} = \frac{E - 1460 \text{ keV}}{(2614 - 1460) \text{ keV}}$$

Multichannel Calibration



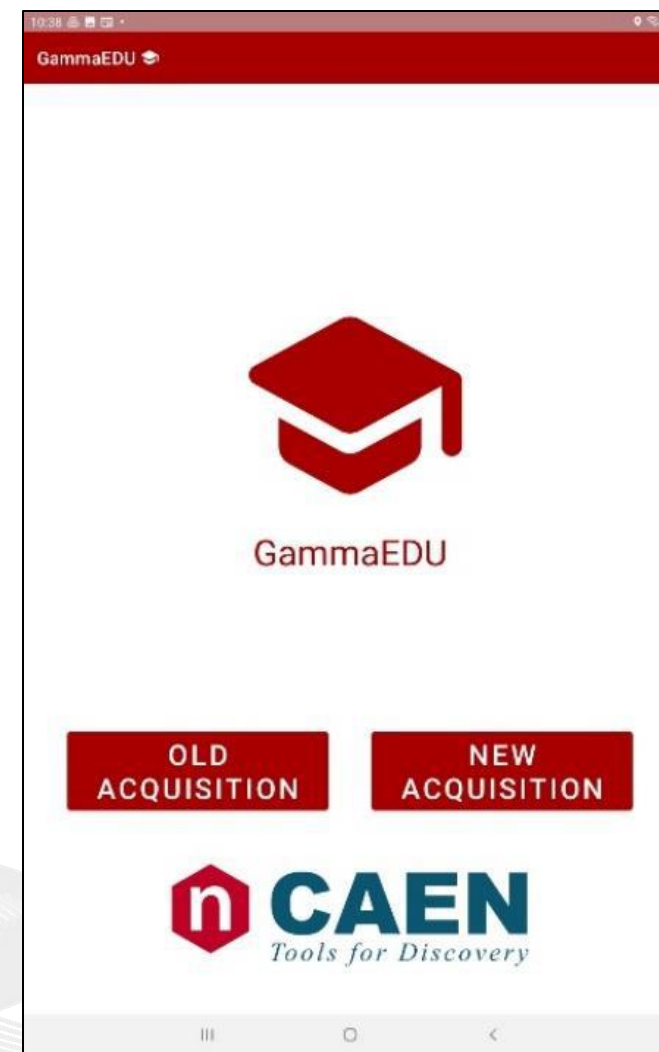
GammaEDU use

- Press the power button
- Verify that the status light is green
- Place the backpack at the point of interest



GammaEDU use (2)

- Activate the tablet and connect it via Bluetooth to the instrument
- Connect the tablet to a WIFI network
- Launch GammaEDU App
- Select «New Acquisition»



GammaEDU use (3)



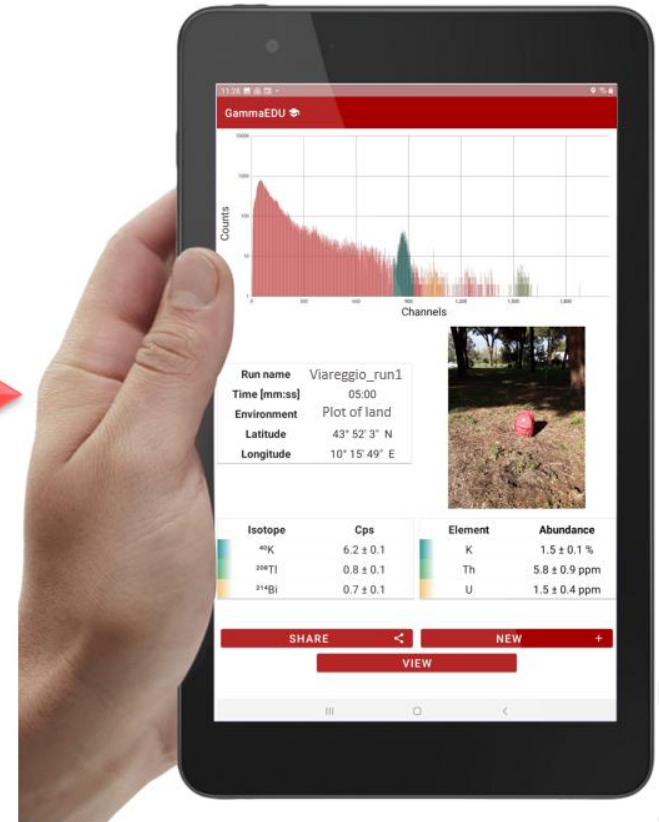
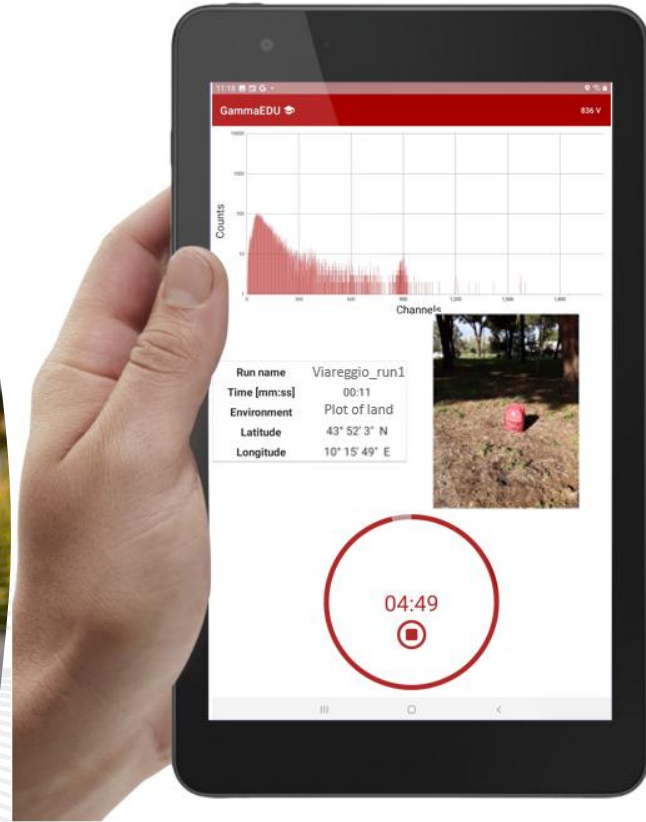
Set the measurement name

Take note of the surrounding environment

Take a measurement situ picture

Set the acquisition time

GammaEDU use (4)



Google Earth

Google Earth Pro

File Modifica Visualizza Strumenti Aggiungi Guida

Ricerca

es.: 94043

Ottieni indicazioni stradali Storia

Luoghi

- I miei luoghi
 - Tour panoramico
- Luoghi temporanei
 - 2022_05_02_17_03_Viar...
 - 2022_05_02_17_16_viar...
 - 2022_05_02_17_25_viar...
 - 2022_05_02_17_31_viar...
 - 2022_05_02_17_40_cae...
 - 2022_08_30_11_28_cae...

Livelli

- Database principale
- Annunci
- Confini ed etichette
- Luoghi
- Foto
- Strade
- Edifici 3D
- Tempo meteorologico
- Galleria
- Altro
- Rilievo

INAIL - Dipartimento Territoriale Viareggio

Via della Vetrata

Giocoheria

Google Earth

1985 43°52'20.59"N 10°15'32.67"E elev 0 m alt 226 m

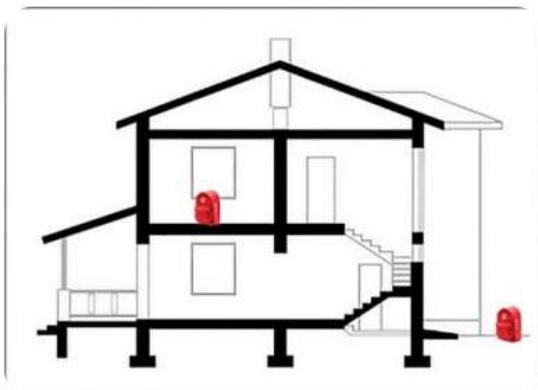
Date and time: 2022_05_02_17_25
Environment: parck pineta

Total activity concentration and Isotopic abundances			
Total Activity Bq/kg	K %	eU ppm	eTh ppm
525 ± 31	1,5 ± 0,1	1,5 ± 0,4	5,8 ± 0,9

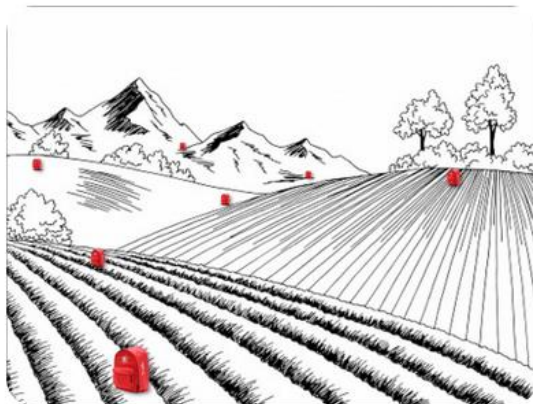
Typical abundancies

40K 0.02 g/g [2%]
 238U 3 µg/g [ppm]
 232Th 10 µg/g [ppm]

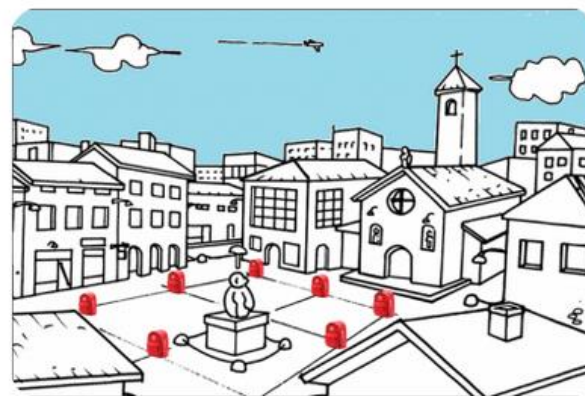
Suggested Experiments



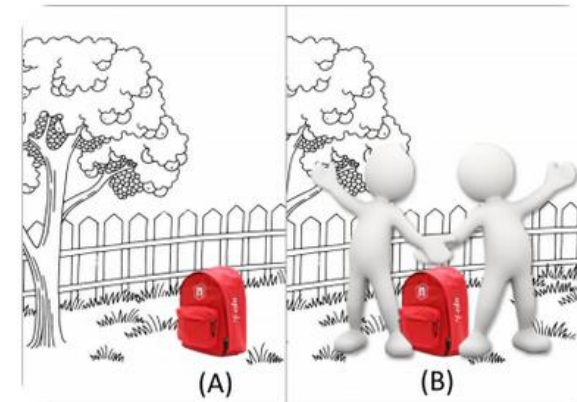
Radiological evaluation of the building materials



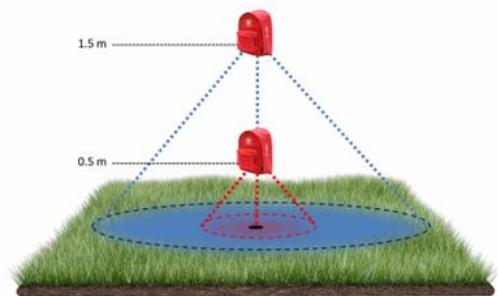
Geochemical and mineral exploration



Radioactivity maps production



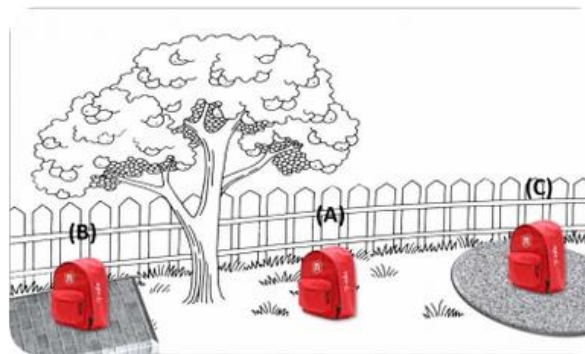
Human body Radioactivity



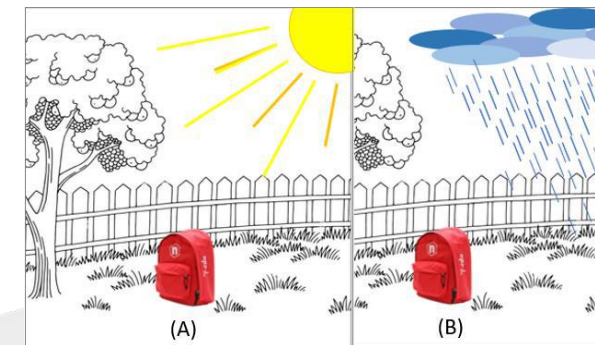
γ Environmental detection as a function of the soil distance



Environmental monitoring in field

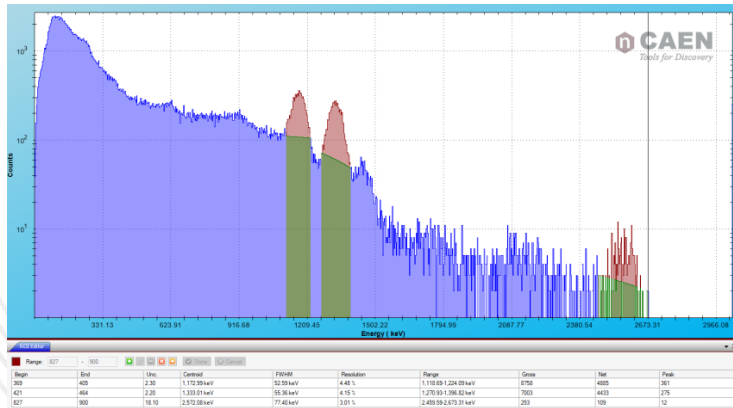


Ground coverage Effect on the Environmental Monitoring

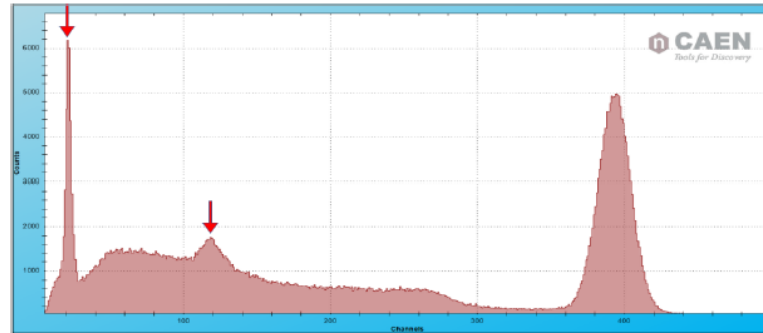


Soil water content evaluation with gamma ray spectroscopy

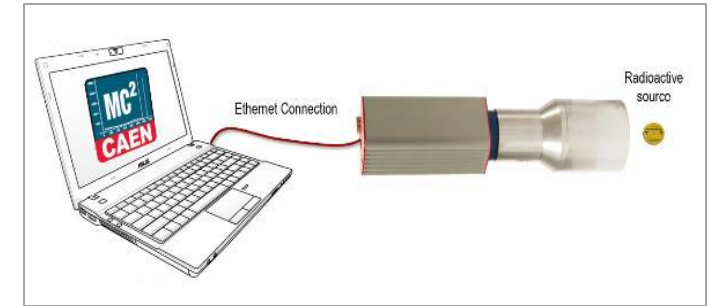
Additional Experiments



Activity of the ⁶⁰Co



Study of the ¹³⁷Cs spectrum: the backscatter peak and X rays



Measurement of Photomultiplier Plateau Curves

Detecting γ -radiation

Poisson and Gaussian Distribution

Energy Resolution

System Calibration: Linearity and Resolution

γ -Radiation Absorption

Photonuclear cross-section/Compton Scattering cross-section

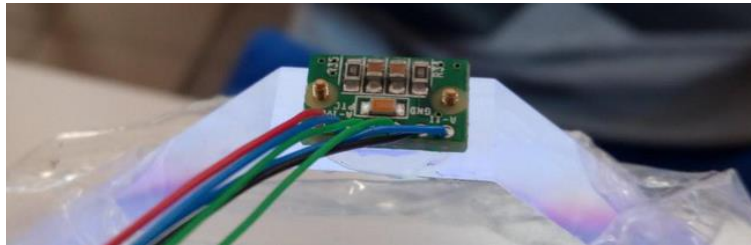
Spares Slides

SP5620CH

Detection System SP5622

Each unit consists of:

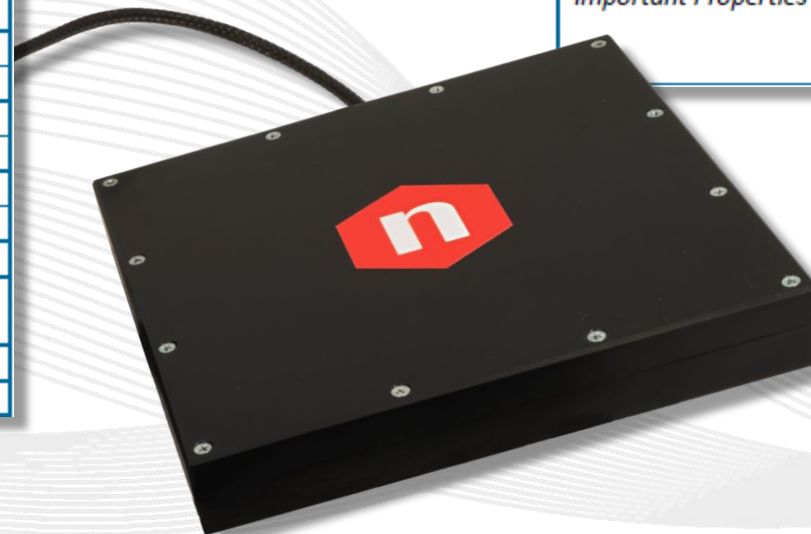
- Plastic scintillator, Polystyrene-based (15 x 15 x 1 cm²)
- Front-end electronic board (transconductance amplifier and a fast discriminator)
- SiPM ASD-NUV4S-P (4 x 4 mm²) mounted in the tile corner at 45°



Feature	Value
Effective active area	4 x 4 mm ²
Number of cells	9340
Cell size	40 μm x 40 μm
Cell fill-factor	60 %
Quenching resistance	800 kΩ
Cell capacitance	90 fF
Recharge time constant	70 ns
Photon Detection Efficiency	43 %
Breakdown voltage	Typical: 26 V Min: 24 V Max: 28 V
Recommended Overvoltage range	Min: 2 V Max: 6 V
Dark Count Rate	< 50 kHz/mm ² @ 2 V OV < 100 kHz/mm ² @ 6 V OV
Gain	3.6x10 ⁶

Geometrical, Electrical and Optical Typical Characteristics of **ASD-NUV4S-P** @ 20°C.

Feature	Value
Scintillator type	UPS-923A
Density	1.06
Refractive index	1.60
Absorption coefficient [cm ⁻³]	0.01-0.003
Softening [K]	355-360
Hygroscopic	no
Emission peak [nm]	425
Light Output [% of anthracene]	60
H/C ratio	1.0
Rise time [ns]	0.9
Decay time [ns]	3.3
Light attenuation length [cm]	400
Important Properties	<ul style="list-style-type: none"> • High light output • Good transparency • Short decay time

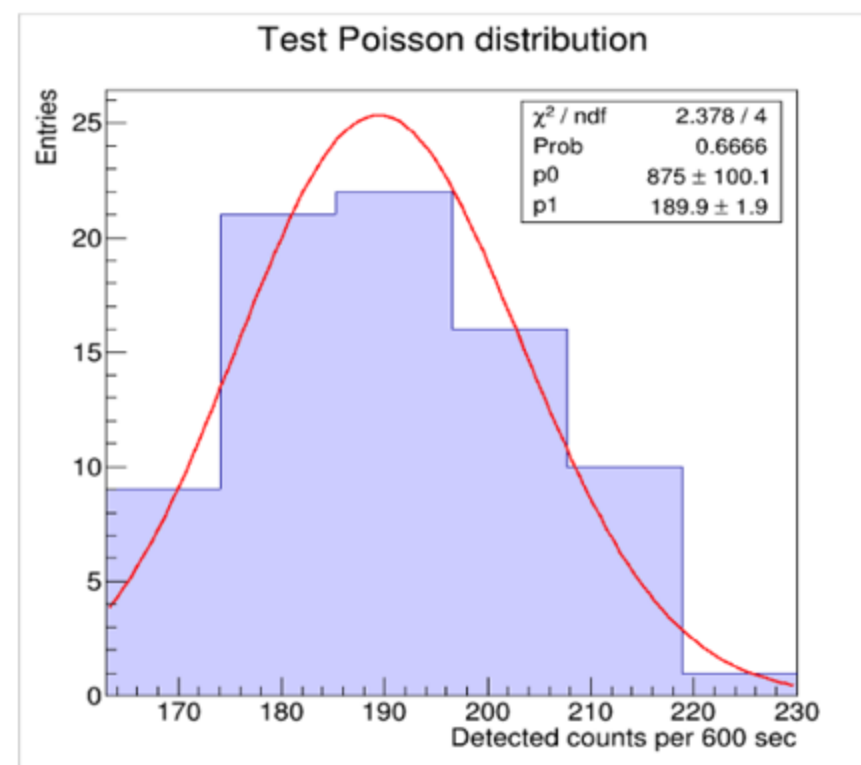


Experimental activity

1) Statistics

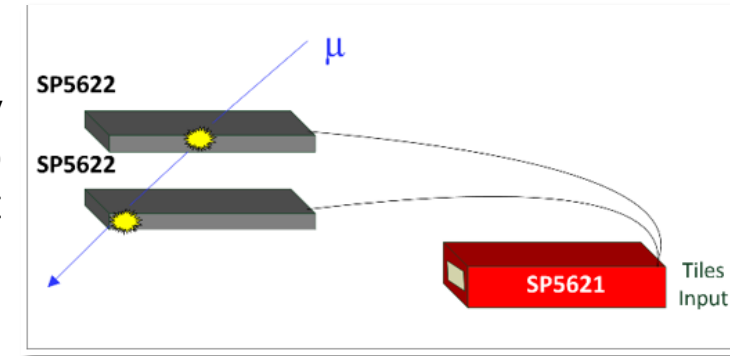
- 2) Muons Detection
- 3) Muons Vertical Flux on Horizontal Detector
- 4) Random Coincidence
- 5) Detection Efficiency
- 6) Cosmic Flux as a function of the altitude
- 7) Zenith Dependence of Muons Flux
- 8) Cosmic Shower Detection
- 9) Environmental and Cosmic Radiation
- 10) Absorption Measurements
- 11) Solar Activity Monitoring

The Poisson distribution of the cosmic rays can be experimentally verified via the data analysis and the treatment of their statistical uncertainty

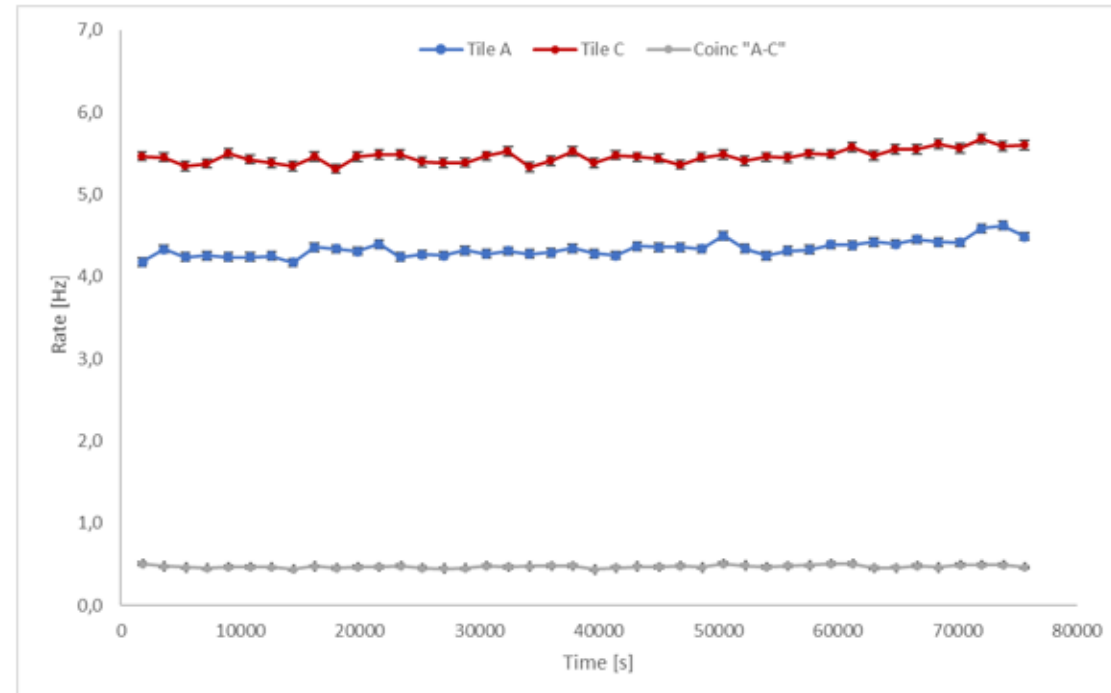


Poissonian distribution of cosmic rays [Fit: $y = p0 * (p1^x / x!) * e^{-p1}$].

Experimental activity



The double tile coincidence has a key role in the cosmic ray detection. It shall be used to reduce the random counts, select the solid angle, and measure the cosmic rate.



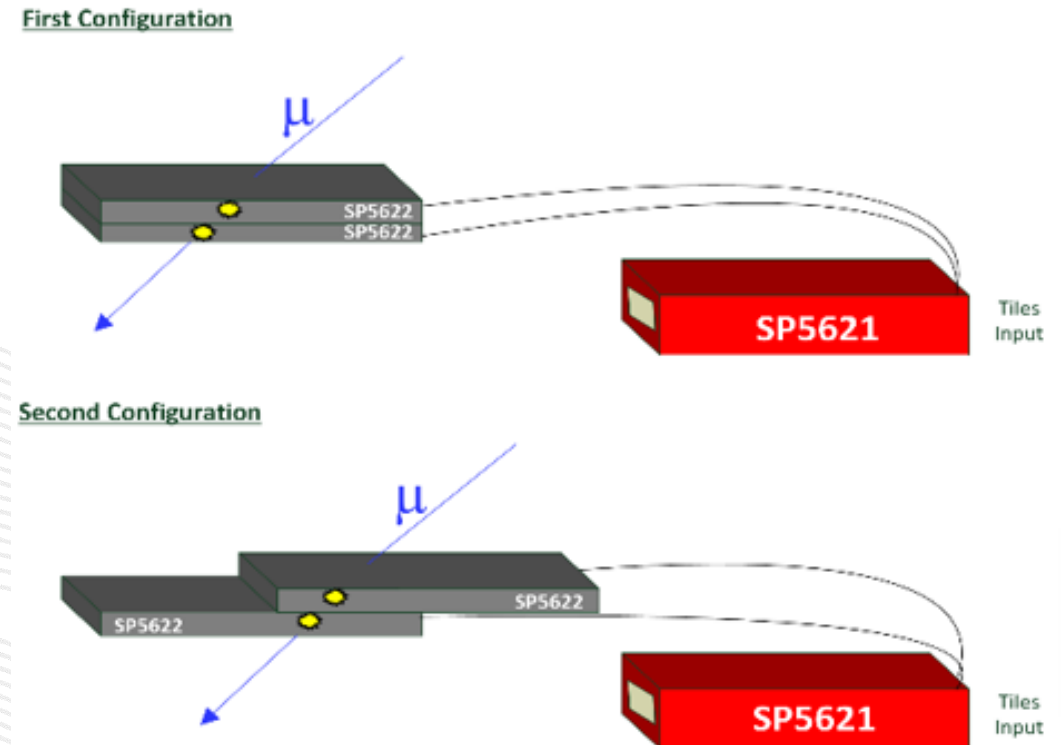
Counts Rate of the single tiles and their coincidence as a function of the time.

- 1) Statistics
- 2) Muons Detection**
- 3) Muons Vertical Flux on Horizontal Detector
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Experimental activity

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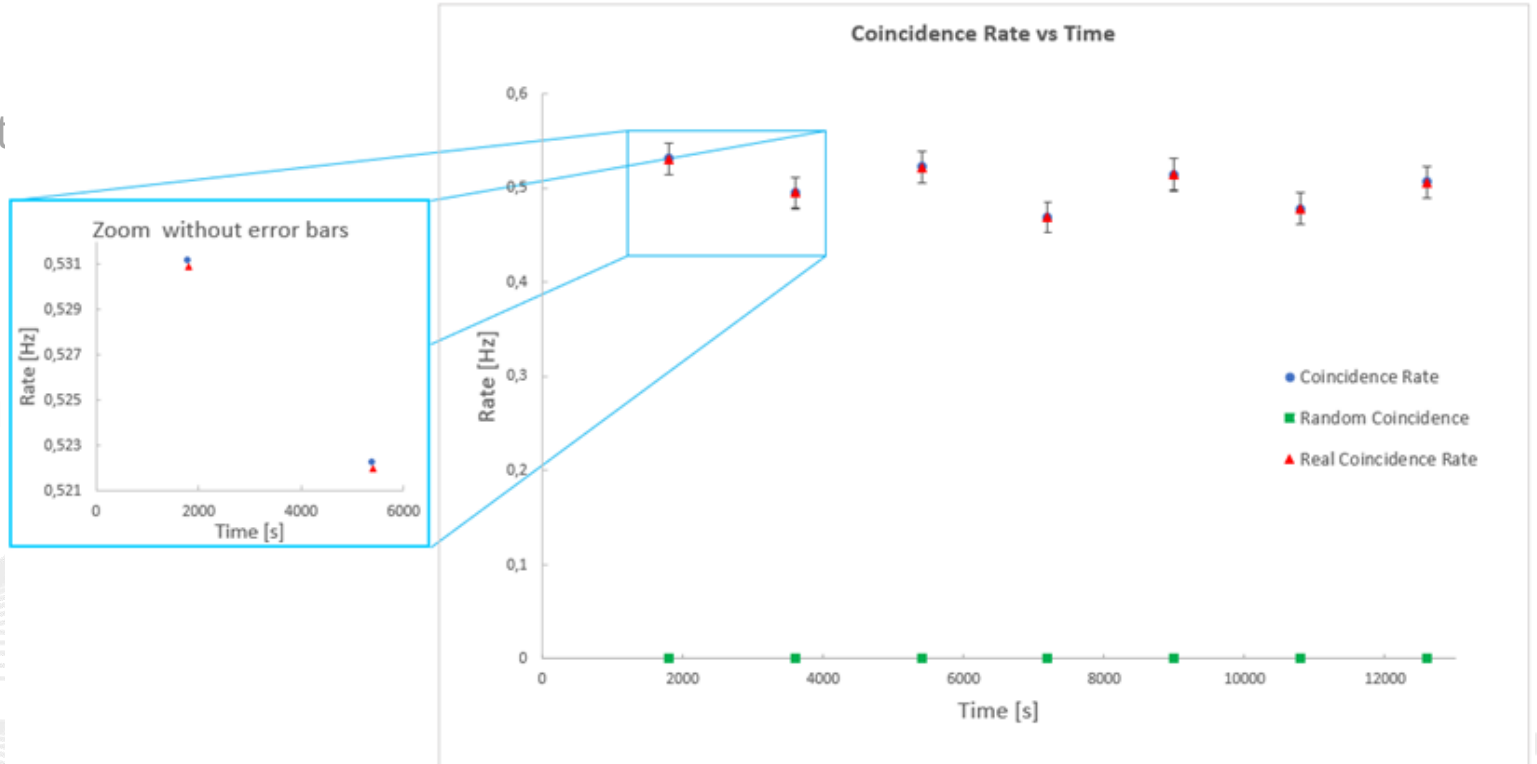
The integral intensity of vertical muons is: $I_v \approx 82 \text{ m}^{-2}\text{s}^{-1}\text{sr}^{-1}$ and their flux for horizontal detectors is $\approx 1 \text{ cm}^{-2}\text{min}^{-1}$ at energies higher than 1 GeV at sea level



Considering the integration over the solid angle, the expected cosmic rate due to the geometry system can be estimated and the detection efficiency can be evaluated.

Experimental activity

- 1) Statistics
- 2) Muons Detection
- 3) Muons Vertical Flux on Horizontal Det
- 4) Random Coincidence**
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- 11) Solar Activity Monitoring



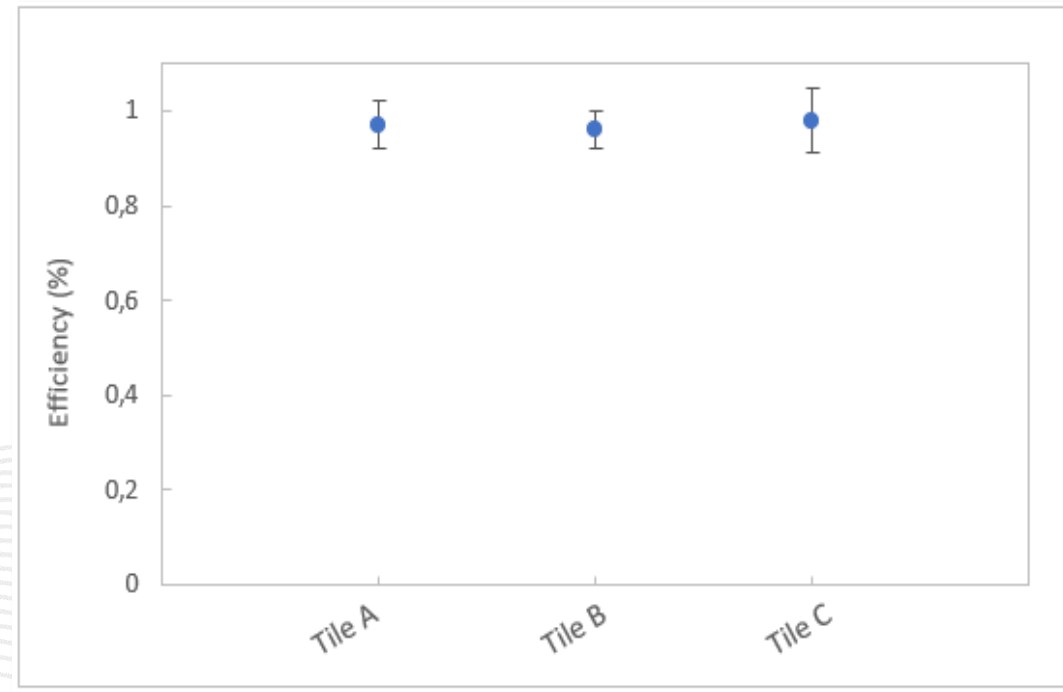
Trend of the Counts Rate and Random Rate as a function of the time.

The plot on the left side is an enlargement of the main plot and underlines the deviation between the measured coincidence rate and the real one, obtained via the random rate subtraction.

Experimental activity

- 1) Statistics
- 2) Muons Detection
- 3) Muons Vertical Flux on Horizontal Detector
- 4) Random Coincidence
- 5) Detection Efficiency**
- 6) Cosmic Flux as a function of the altitude
- 7) Zenith Dependence of Muons Flux
- 8) Cosmic Shower Detection
- 9) Environmental and Cosmic Radiation
- 10) Absorption Measurements
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$$N_{AC} = \epsilon_A * \epsilon_C * N_0 \quad \text{and} \quad N_{ABC} = \epsilon_A * \epsilon_B * \epsilon_C * N_0 \quad \rightarrow \quad \epsilon_B = N_{ABC} / N_{AC}$$



Additional tools

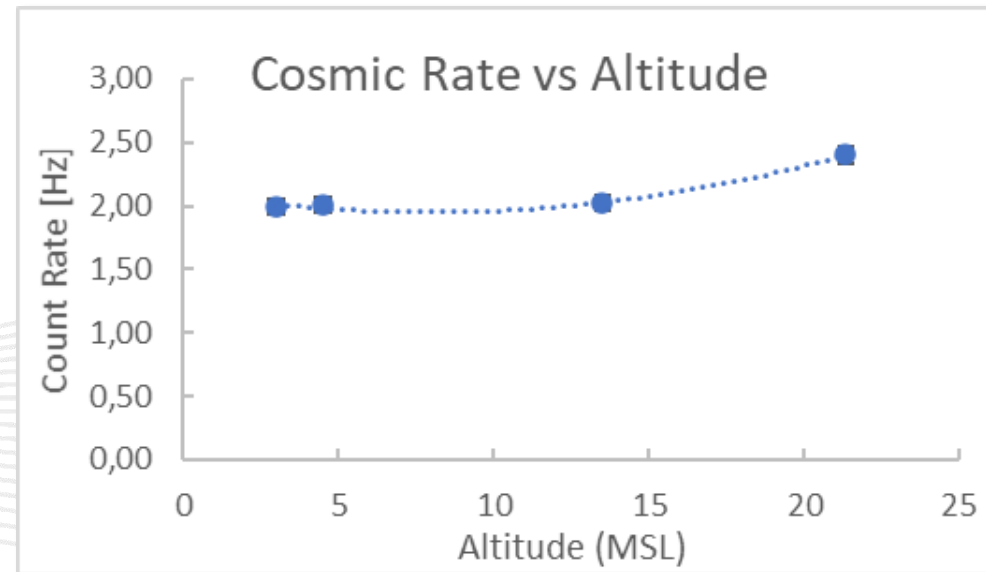
- n.1 SP5622 - Detection System
- n.1 DT1081A - Four-Fold Programmable Logic Unit and n.1 Cable Adapter



Experimental activity

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The experiment proves in a simple way the not terrestrial origin of the cosmic radiation. For a better comprehension of the cosmic flux behaviour as a function of the altitude, it is suggested to cover the floor with lead bricks.



<https://www.chateau-doex.ch/en/Z4237/festival-des-ballons-homepage>



Prof. Hans Peter Back

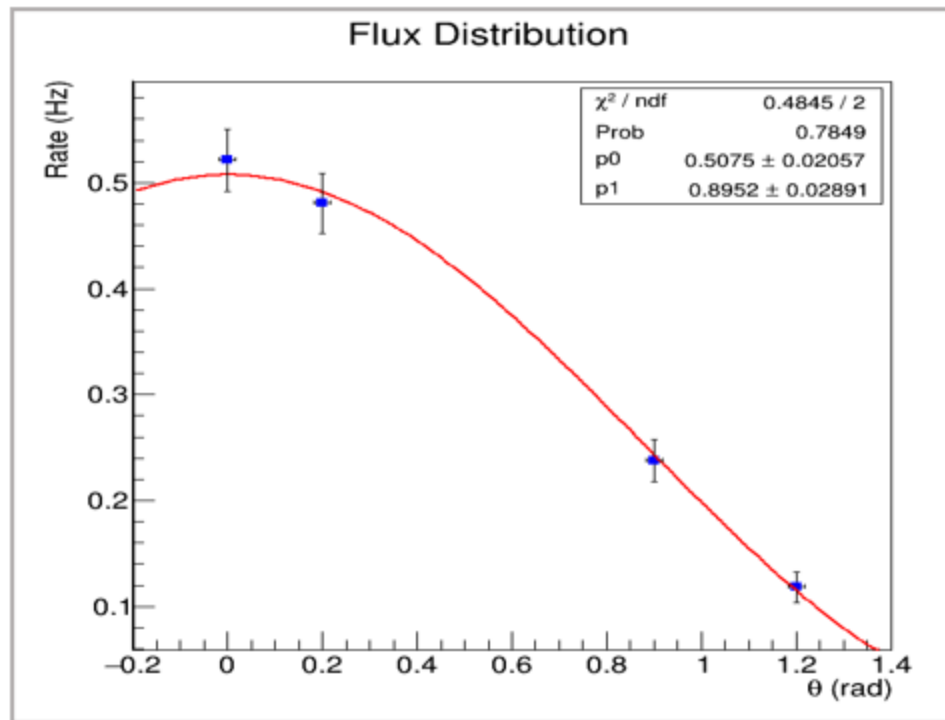
Albert Einstein Center for Fundamental Physics, University of Bern

[Link to customer](#)



Experimental activity

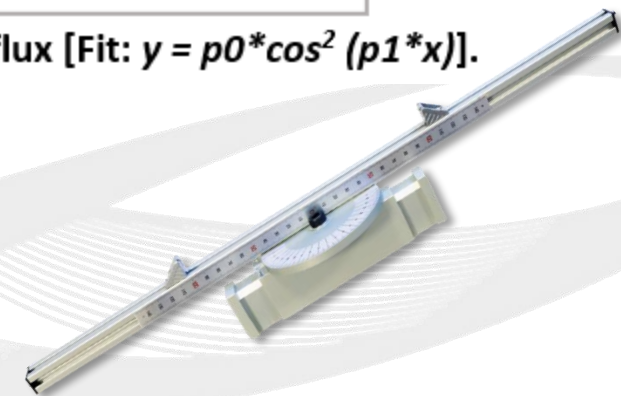
- 1) Statistics
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Zenith angle dependence of the muons flux [Fit: $y = p0 * \cos^2 (p1 * x)$].

Additional tools

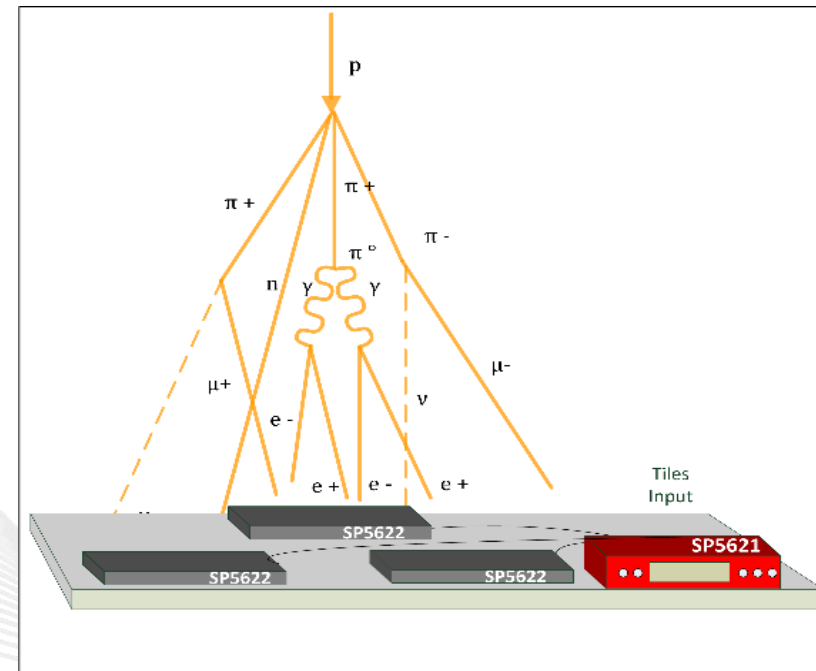
SP5609 - Telescope Mechanics



Experimental activity

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Observation of the cosmic ray showers, namely cascades generated by cosmic rays interacting in the atmosphere.



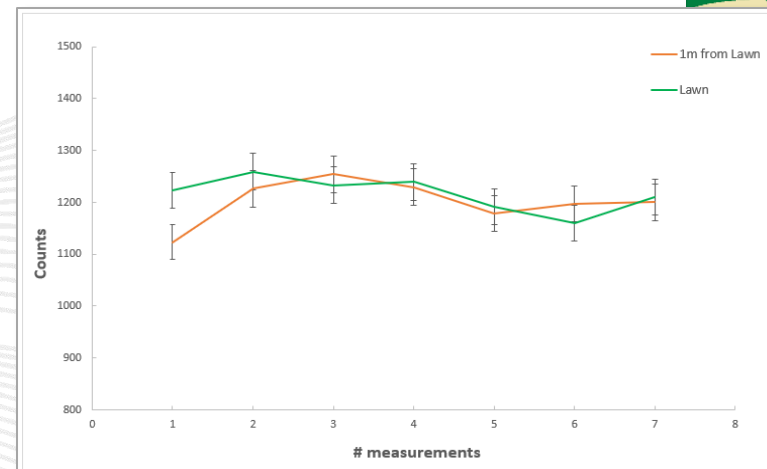
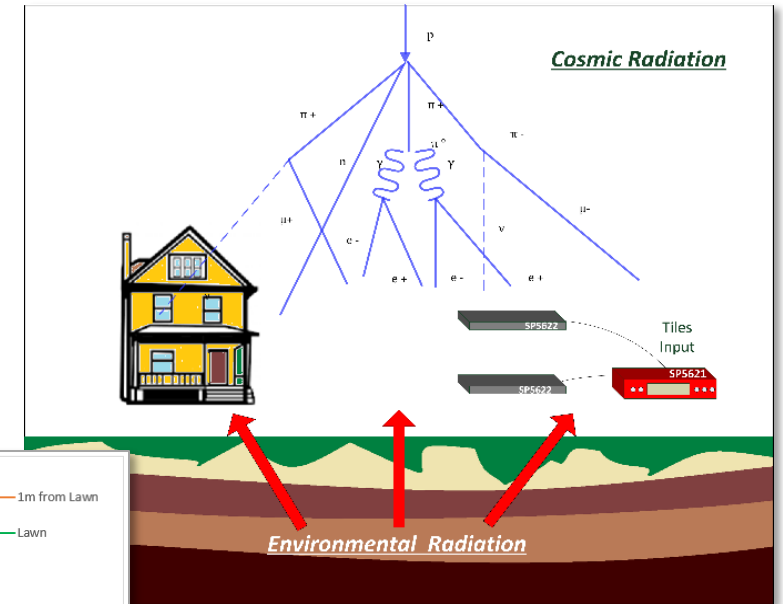
Additional tools

- SP5622 - Detection System



Experimental activity

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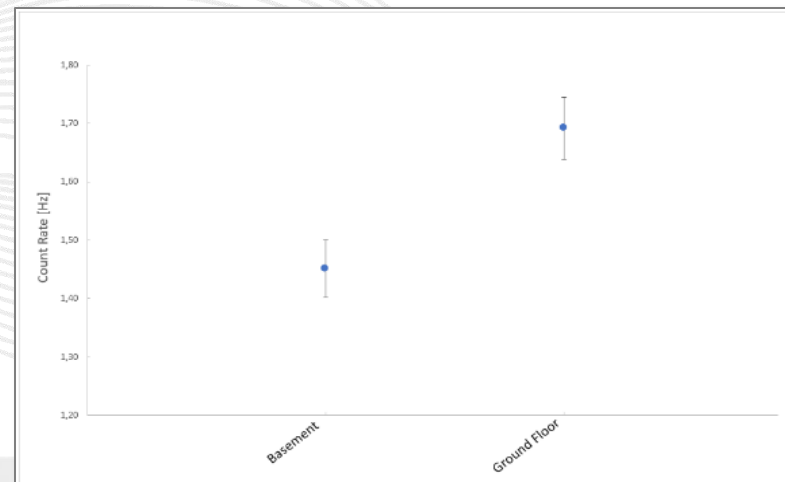
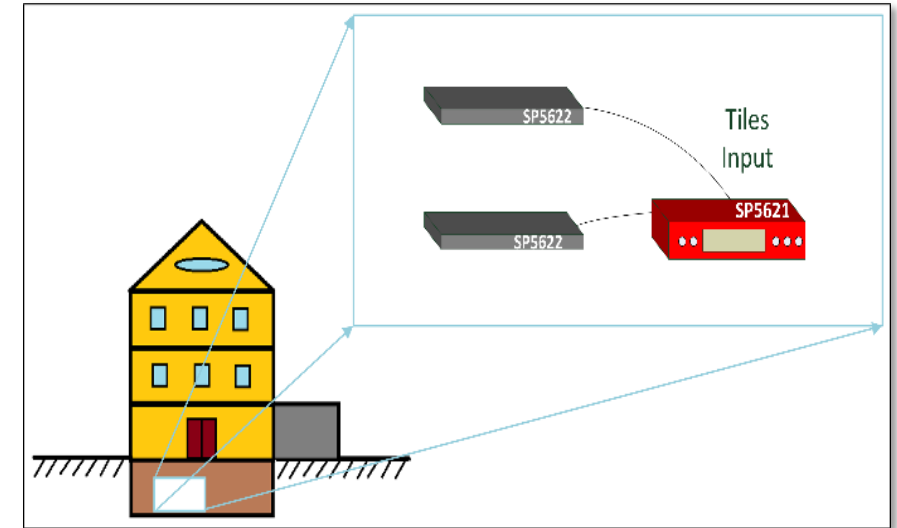


The students may get acquainted with the presence of natural radioactivity by identifying environmental and cosmic contributions via simple comparison of the counting measurements at different height.

Experimental activity

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The main goal of the experiment is the verification of the absorption of the cosmic rays passing through a matter thickness and the related observations about the crossed material.

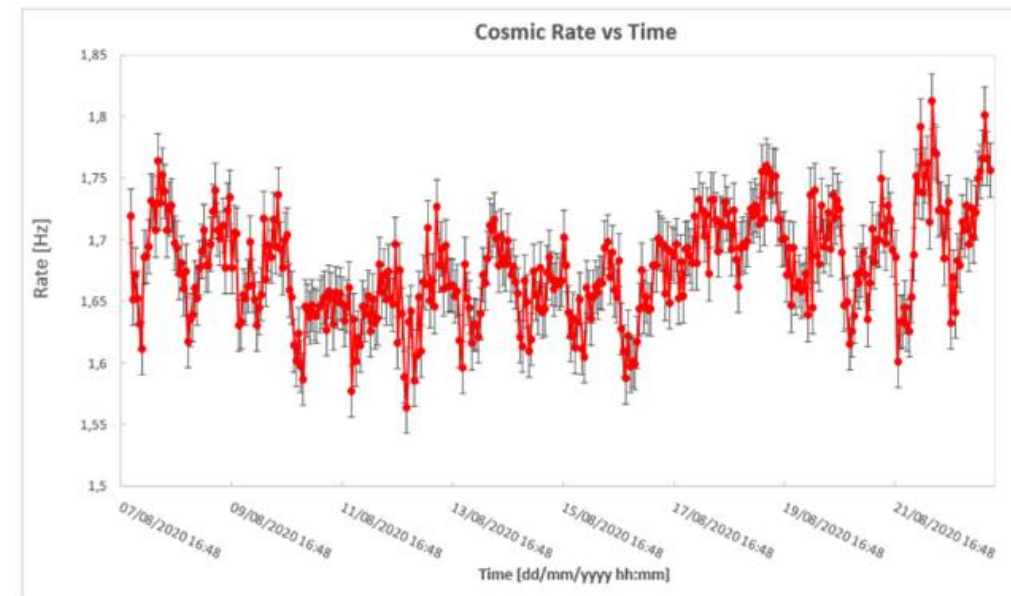
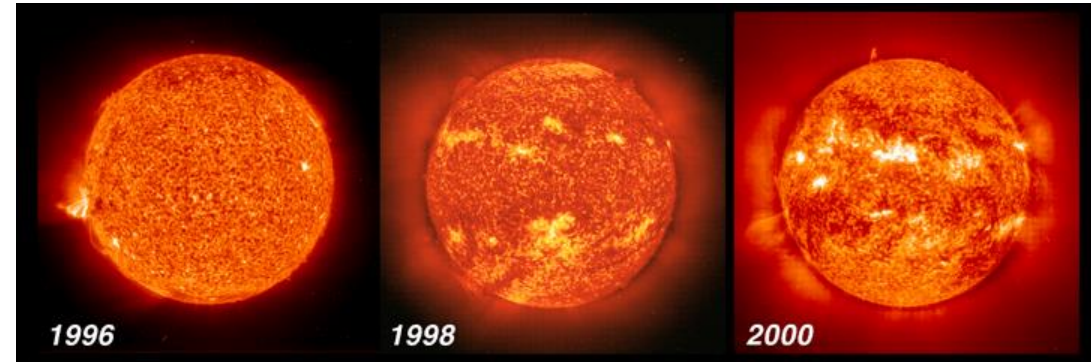


The students may estimate the absorption extent by comparing the results of the measurements performed underground or inside a building or a cave, and outside, without any matter barrier. Moreover, by knowing the thickness, some hypothesis about the average density of the material can be expressed.

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The typical cosmic rate night /day trend can be sometimes modified due to solar activity changes.

Spares Slides

SP5640

Educational Kits Description



SP5600E – Educational Photon kit



SP5600D – Educational Beta kit



SP5600C – Educational Gamma kit



SP5600AN – Educational kit Premium Version



SP5600 - Power Supply and Amplification Unit

- Two channels
- Independent biasing (max 120 V, 100 μ A)
- 2 stage amplification [500 MHz bandwidth, tunable gain up to \sim 50 dB]
- Fast leading-edge discriminator (\pm 2V)
- Coincidence logic
- Active feedback control on V_{bias} for Gain stabilization (granularity: 0.1 $^{\circ}$ C)
- USB 2.0 interface

DT5720A - Desktop Digitizer



- Digital Pulse Processing for Charge Integration DPP-CI
- Good timing resolution with fast signals (rise time < 100 ns)
- 2 channels
- Stand-alone
- 250 Ms/s, 12 bits
- \pm 1V input range
- Optical Link and USB 2.0 interfaces

Educational Beta Kit Description



SP5600E – Educational Photon kit



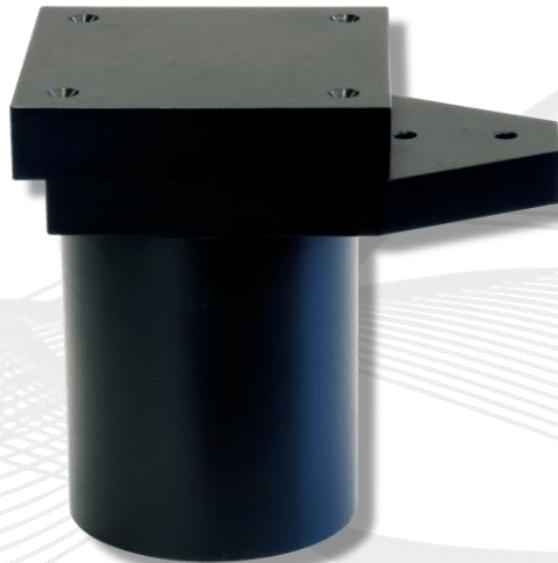
SP5600D – Educational Beta kit



SP5600C – Educational Gamma kit



SP5600AN – Educational kit Premium Version



SP5608 - Scintillating Tile

- Sensitive volume: 47 x 47 x 10 mm³
- Scintillator: polystyrene
- Directly coupled on HAMAMATSU MPPS S13360- 6050CS
 - Effective photosensitive area : 6 x 6 mm²
 - Pixel pitch : 50 μm
 - Number of pixels : 14400
- n° 20 Paper and Aluminum sheets
- Teflon tape

• Beta Spectroscopy
• Cosmics

Educational Beta Kit Description



SP5600E – Educational Photon kit



SP5600D – Educational Beta kit



SP5600C – Educational Gamma kit



SP5600AN – Educational kit Premium Version

Additional Tools

Additional SP5608 - Scintillating Tile



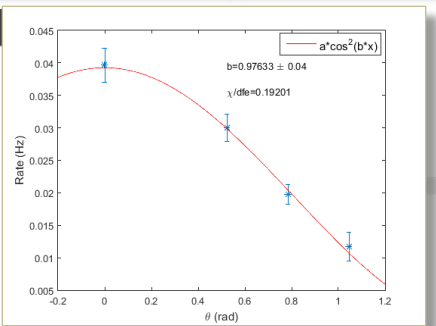
SP5609-Telescope Mechanics



Telescope Mechanics allows the easy construction of a muons telescope. It is composed of :

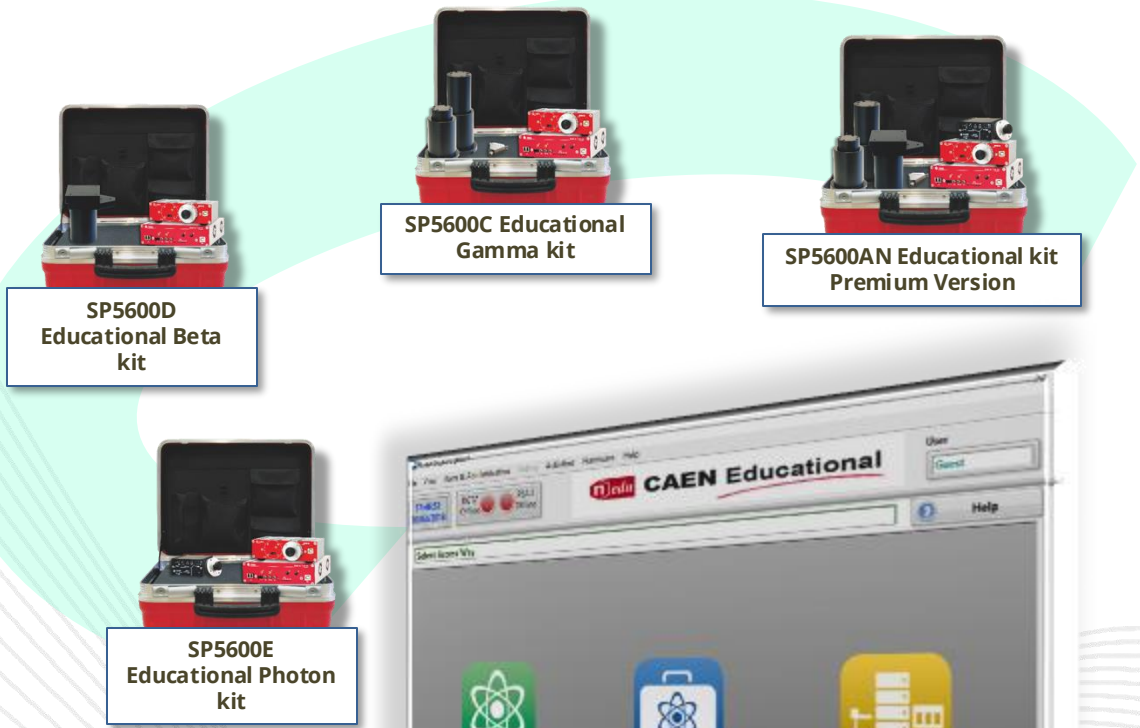
- Rotary axis with desk support
- Clamps with screws
- Angle brackets kit

Suggested Application



- **Double coincidence**
- **Zenith Dependence of Muons Flux**

Educational Kits - HERA Software



- ❑ **Help** and **QuickStart Guide** online
- ❑ **Three Access Levels** to the software functionalities:
 - **Level 1** - Hardware Management
 - **Level 2** - Hardware Management + Experiments
 - **Level 3** - Full Access (Analysis Tools)



Handy Educational Radiation Application

Analysis Tools are implemented in the software itself!

