

C.1.2 Statistics

SG6210D



Dedicated kit	
Description	pp.
SP5620CH Cosmic Hunter	182



Requirements

No other tools are needed.

Difficulty	Execution Time	Data Analysis	Radioactive Sources
    	   	NO	YES

Equipment

SP5620CH - Cosmic Hunter

Model	SP5621	SP5622 (x2)
Description	Coincidence Module  p. 182	Detection System  p. 187

Purpose of the experiment

Statistical properties of the cosmic rays.

See the
Application



Fundamentals

The event number in a given time interval is one of the most interesting points in many Physics phenomena. This number is often affected by statistical fluctuations around an average value determined by the type of phenomenon. Multiple factors may cause fluctuations and influence the measurement result. Thus, the exact value is not always the same (as in the case of particles that decay may derive from space or from a radioactive source).

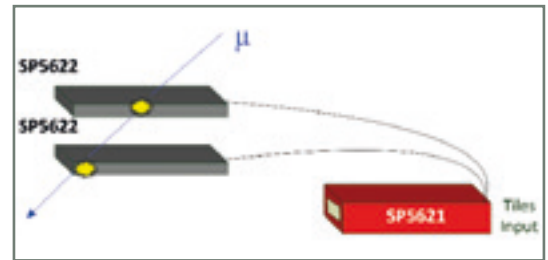
The most important goal in the experimental approach is to understand which values can occur in a series of measurements as well as their probability, i.e. the probability distribution. The Poisson distribution describes with good approximation events coming from radioactive phenomena or from counting cosmic rays. This distribution expresses the probability of a given number of events occurring in a fixed interval of time or space, and can be expressed as:

$$P_{\mu}(n) = (\mu^n / n!) * e^{-\mu}$$

where μ is the average number of events in a fixed interval and n is the number of events.

Carrying out the experiment

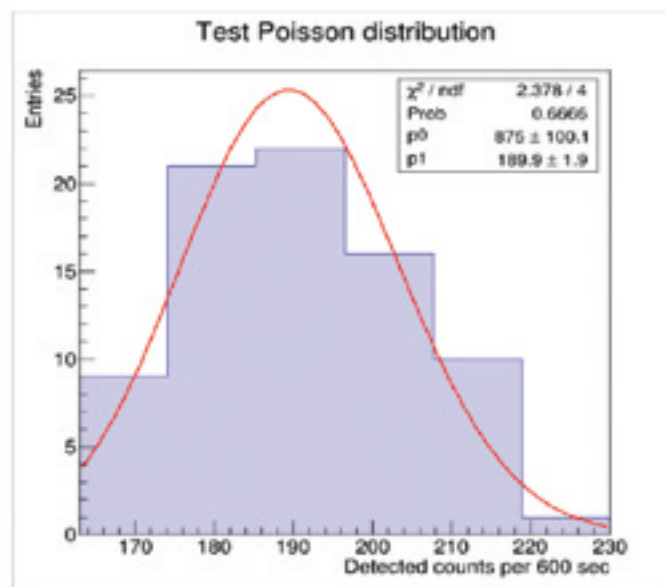
Connect the cable connector from each SP5622 Detection System to the input located on the rear panel of the SP5621 module. Power on the SP5621 module and start acquisition via the START button on the front panel. When a charged particle crosses the black tile its energy is converted into scintillation light. The photons produced are detected by the photosensor and converted into an electrical signal. The number of counts for each scintillator is available via the SP5621 display. Note that spurious electrical signals will likely also be detected by the photosensor, thus producing noise. Coincidence between two SP5622 will greatly reduce the number of spurious events. However, the statistical fluctuations cannot be totally removed. Select the scintillators coincidence via the related button on the front panel, then select the integration time of the measurement. Before starting acquisition choose the system geometry. Be sure to keep this geometry constant for the duration of the experiment. Take and record more data to obtain statistical significance.



Experimental setup block diagram.

Results

The Poisson distribution of cosmic rays can be experimentally verified via data analysis and the treatment of their statistical uncertainty.



Poissonian distribution of cosmic rays [Fit: $y = p_0 * (p_1 x / x!) * e^{-p_1}$]

This experiment is also possible with the following kits

A SP5600D Educational kits



see
pp. 182,
179

B SP5600AN Educational kits ★ PREMIUM



see
p. 179